

Neutron capture reaction of r-process in Common Envelop Jet Supernova

Shilun Jin

Institute of Modern Physics, Chinese Academy of Sciences

Rußbach, Austria
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- 1 Rapid neutron capture process
 - sites of r-process
 - r-process in common envelop jet supernova
- 2 Measurement of reaction rate on $^{59}\text{Fe}(n,\gamma)$
 - setup
 - preliminary results
 - HIAF in future



How were elements from Iron to Uranium made?

Nucleosynthesis process

N-deficient side:

rp-process

P-process:

p-nuclide ^{74}Se - ^{196}Hg

vp-process:

$^{92,94}\text{Mo}$ $^{96,98}\text{Ru}$

N-rich side:

s-process:

- slow (n,g) fast β decay

- AGB

- weak, strong/main

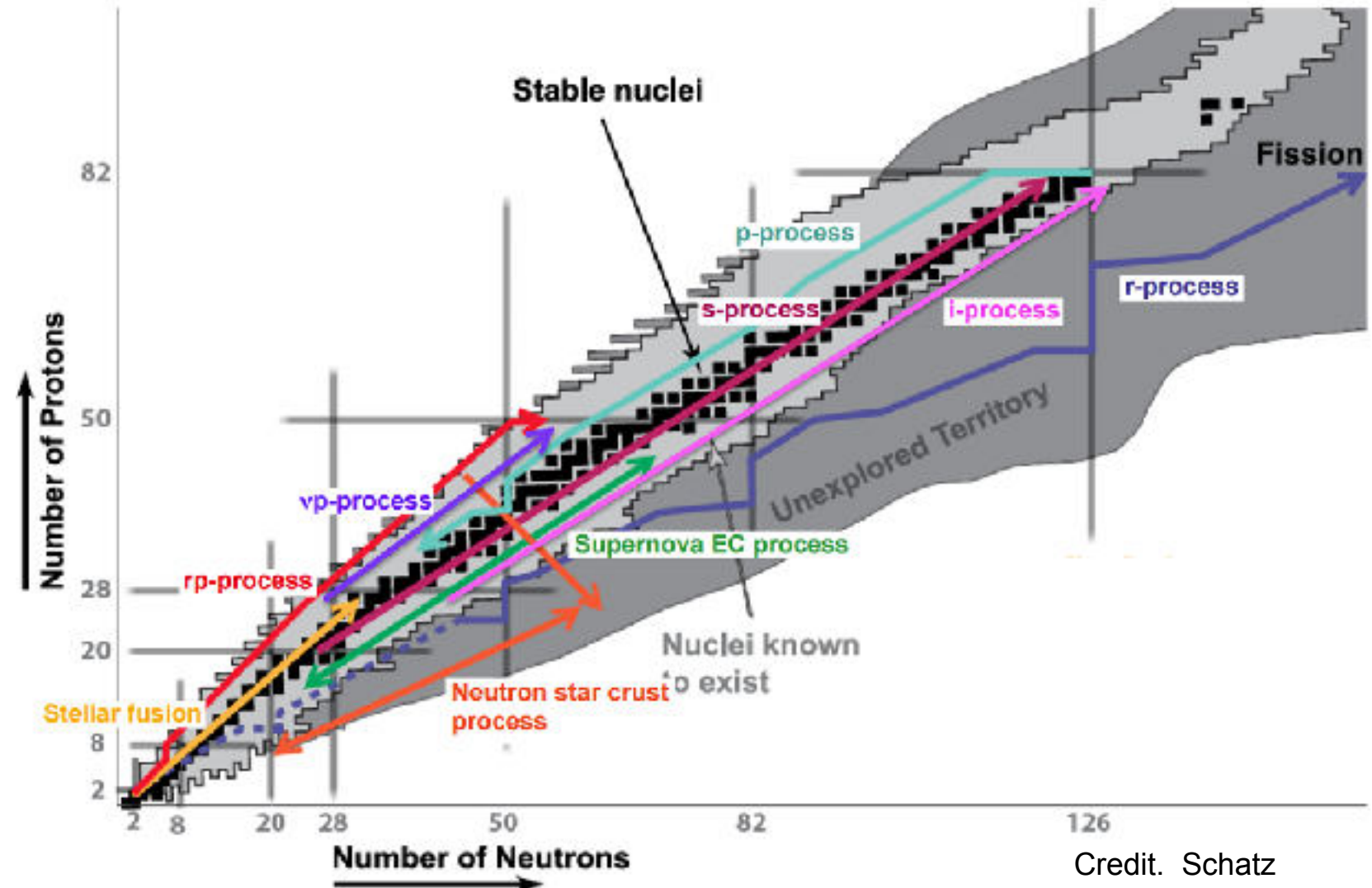
r-process:

- fast (n,g) slow β decay

- sites?

- weak, strong/main

i-process



Credit: Schatz



Neutron Star Merger

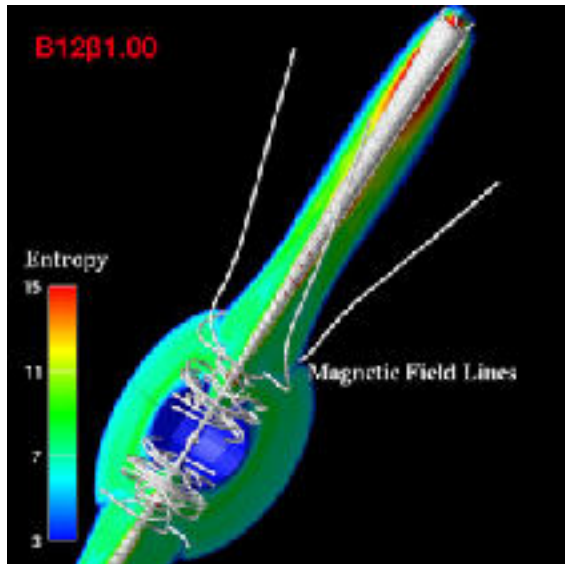
GW170817

- GW LIGO/VIRGO 2017
- Afterglow AT 2017gfo
- Lanthanides 2017
- Strontium 2019
- ^{191}Os , ^{234}Th 74days later 2019



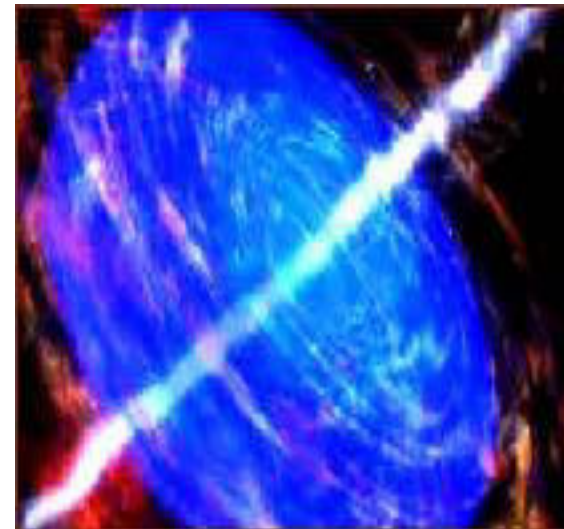
Core Collapse Supernovae

- death of massive star
- core collapse & shock wave
- neutrino driven wind
- no heaviest element
- can be proton rich



Magneto HydroDynamics Supernovae

- strong magnet 10^{12-15}G ,
- high spin
- 3D & 2D models
- 2021 UMP SMSS J200322.54-114203.3
- Weak & strong r-process

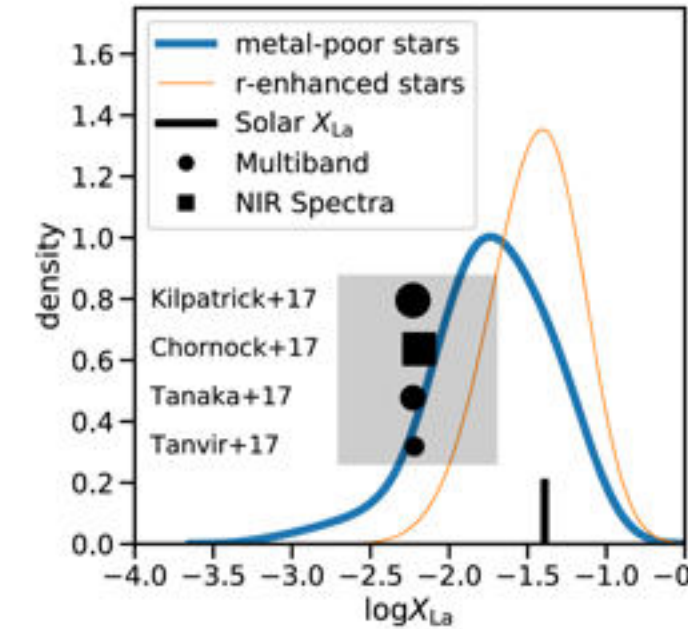


Collapsar

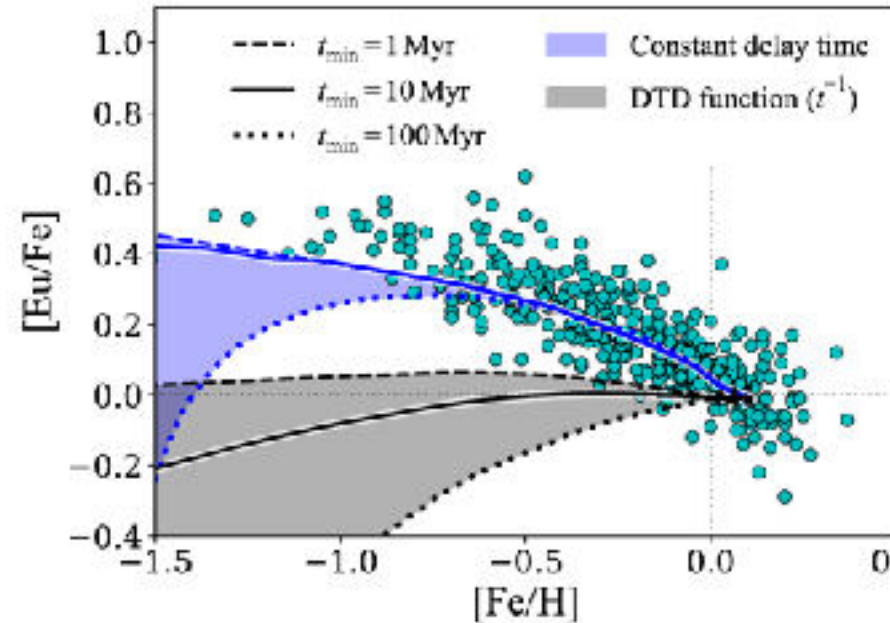
- massive, rapidly rotating star
- accretion onto the black hole
- produce a beamed jet
- long-duration γ -ray bursts
- 3D model 2019, Seigel et al



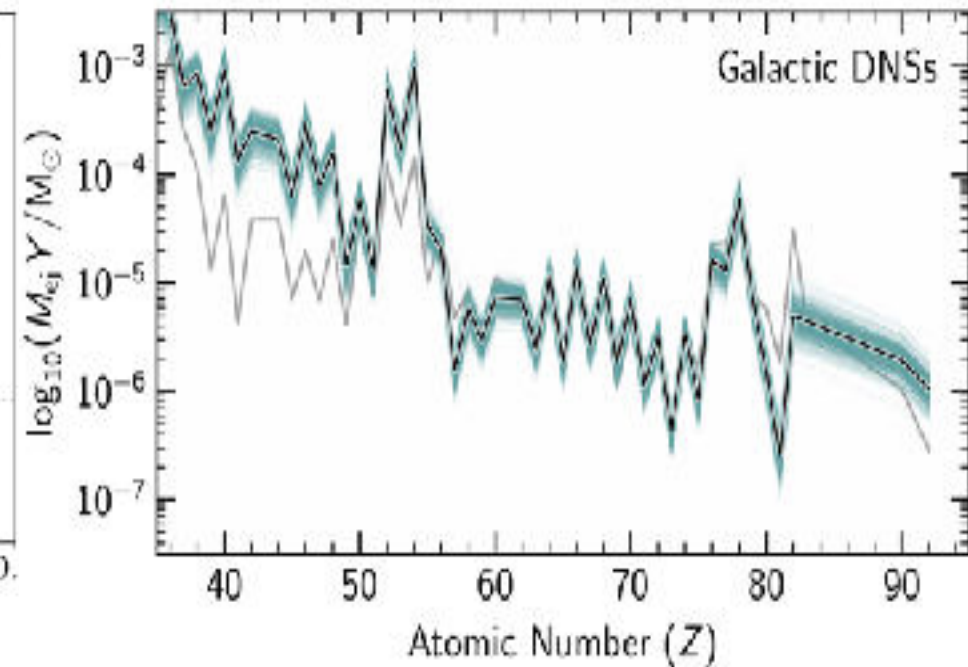
Is the neutron star merger only dominated site of r-process?



Ji et al. 2019
insufficient La



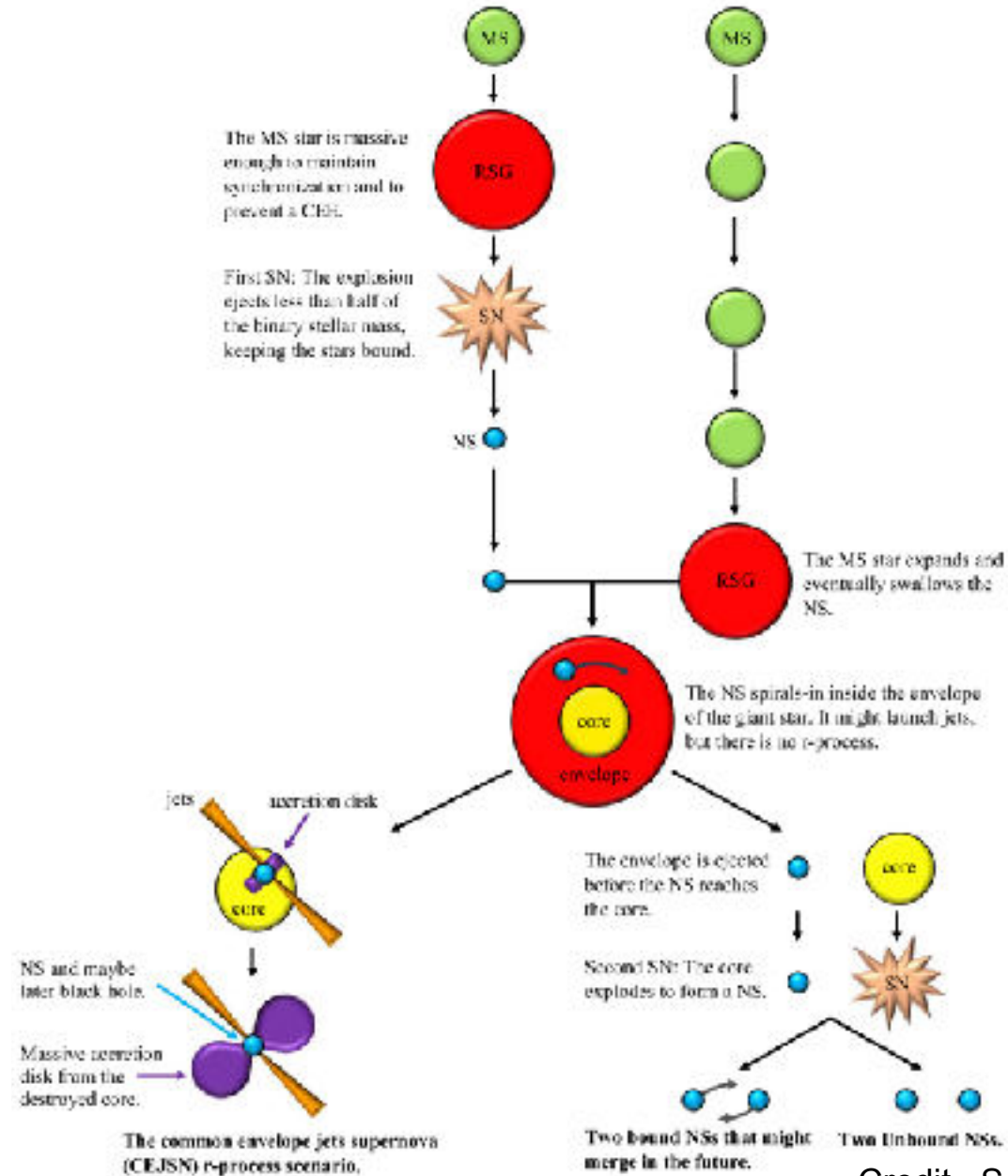
Côté et al. 2019
Inconsistent with $[\text{Eu}/\text{Fe}]$

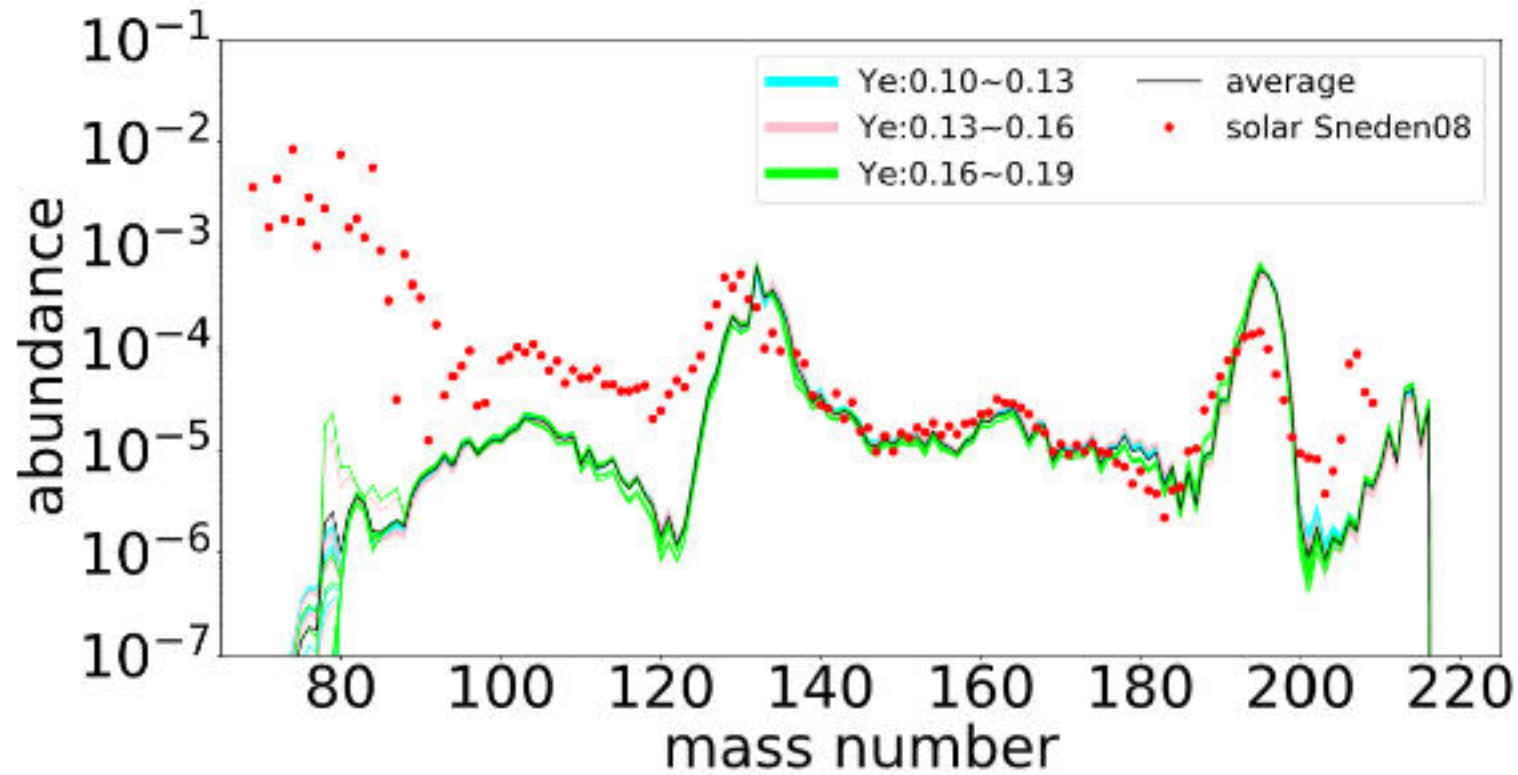


Holmbeck et al. 2023
Data of DNS, weaker than solar

Common Envelope Jet Supernovae

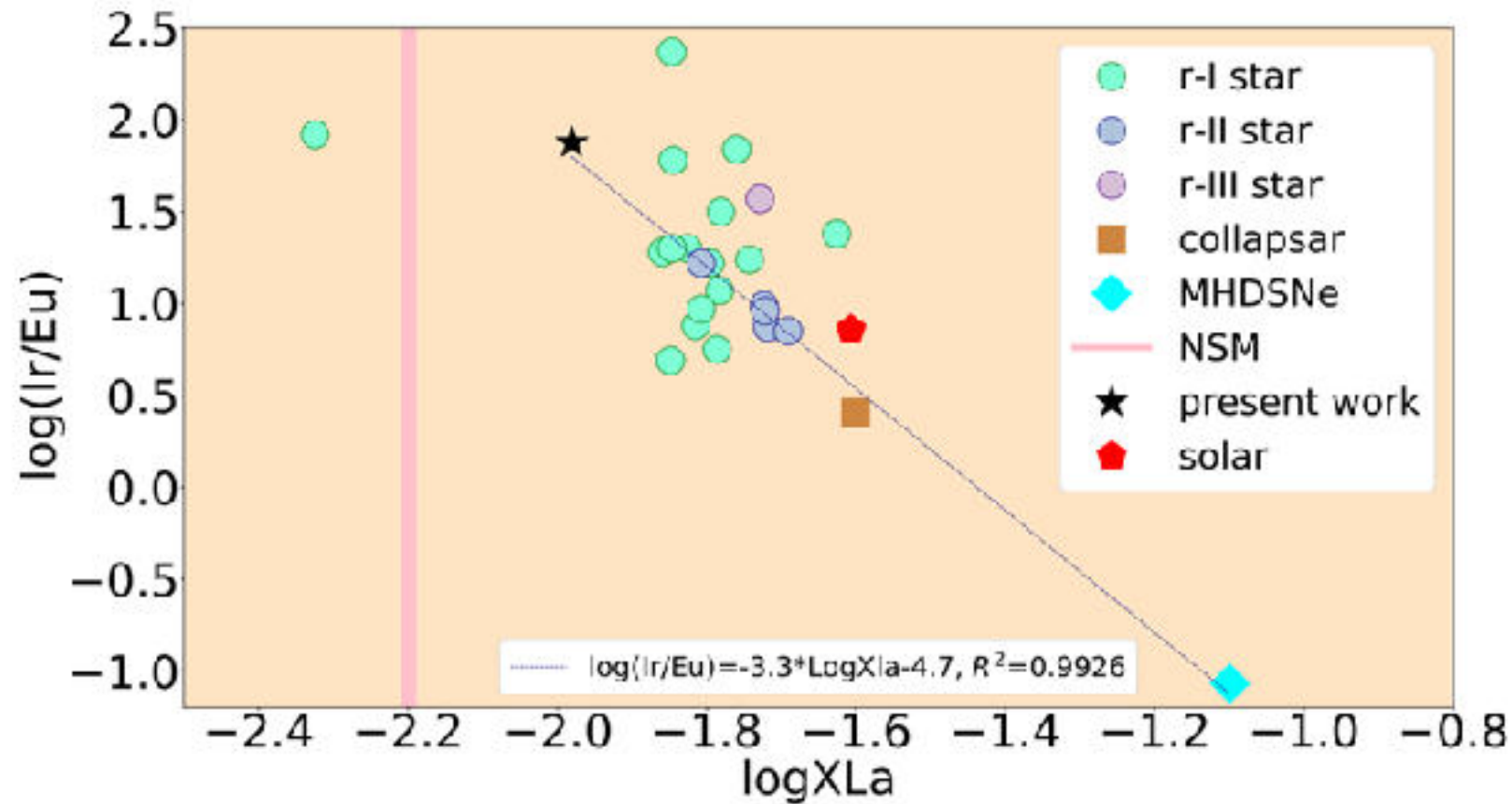
- death of two stars
- RSG and NS
- RSG swallows NS
- Ns spirals in core
- accretion disk formed
- core destroyed
- bipolar jet





- New density function
- Favor $A > 140$
- Agreement with La
- Strong 3rd peak

Jin & Soker 2024



- Define $\log(\text{Ir}/\text{Eu})$
 - Explain r-enhanced stars
 - Actinide Boost Star
- 2MASSJ09544277+5246414
age ~11b year

Jin & Soker 2024



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Heavy Ion Research Facility in Lanzhou (HIRFL)

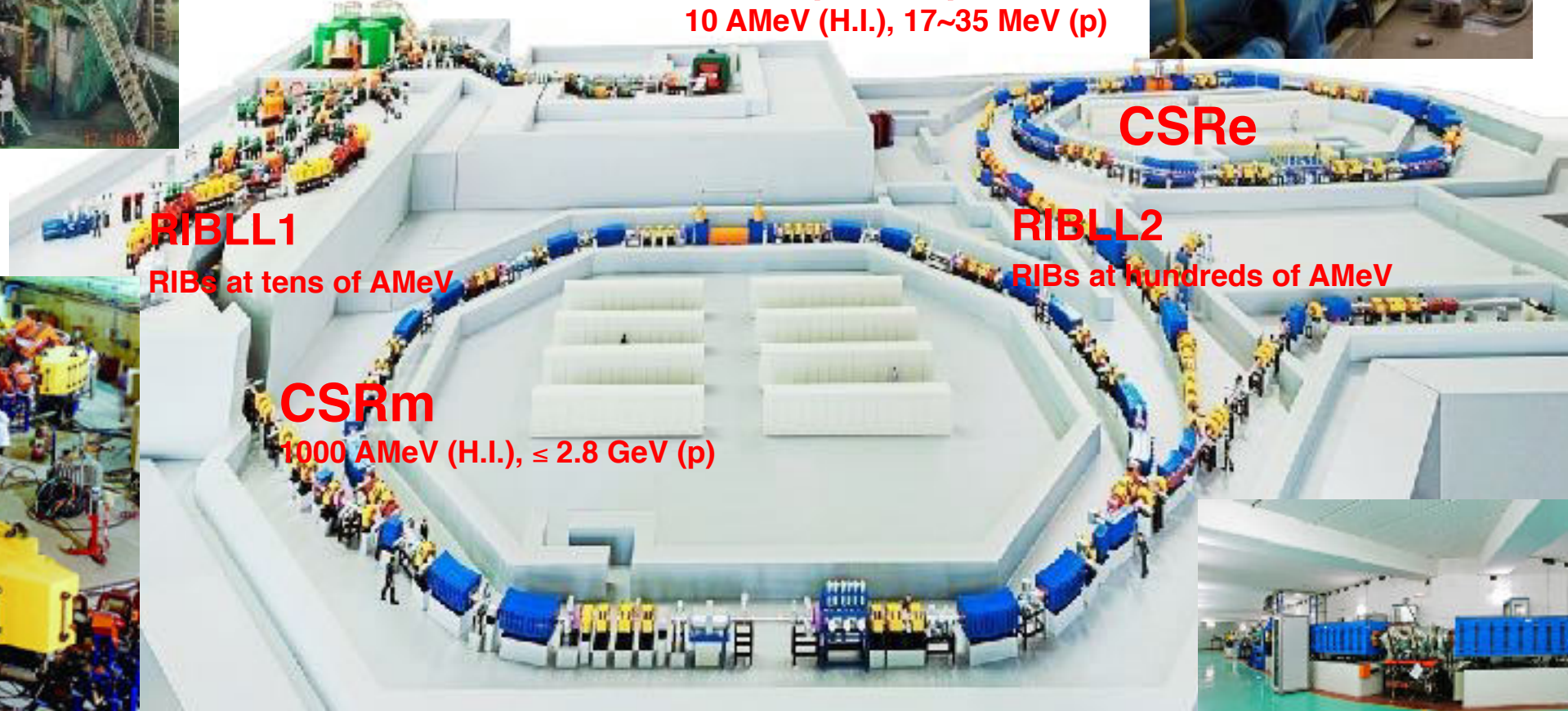


SSC(K=450)

100 AMeV (H.I.), 110 MeV (p)

SFC (K=69)

10 AMeV (H.I.), 17~35 MeV (p)



RIBLL1

RIBs at tens of AMeV

RIBLL2

RIBs at hundreds of AMeV

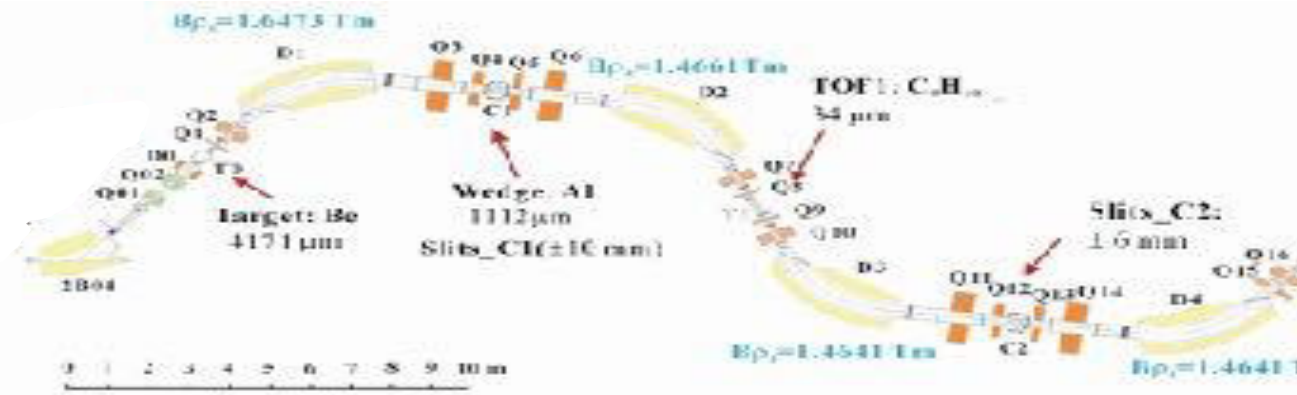
CSRm

1000 AMeV (H.I.), ≤ 2.8 GeV (p)

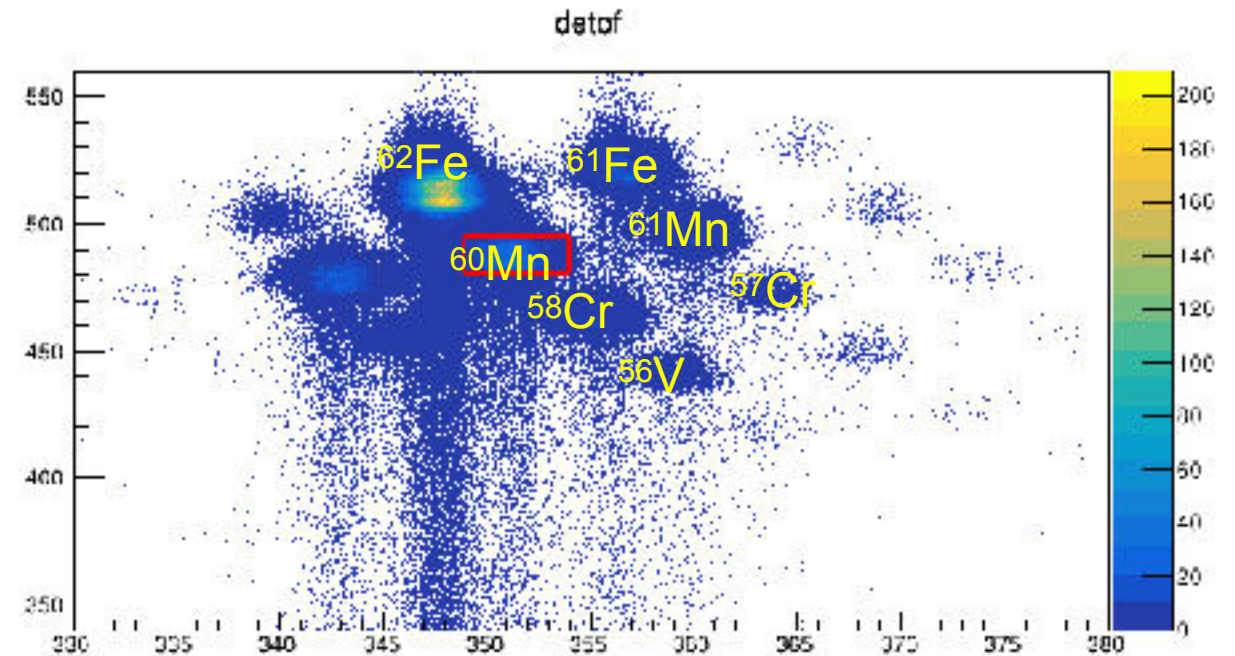
CSRe



Radioactive Ion Beam Line in Lanzhou (RIBLL)



- Primary Beam: $^{64}\text{Ni}@60\text{MeV/u}$
- Fast Beam(PF): ^{60}Mn
 $^{60}\text{Mn} \xrightarrow{\beta\text{decay}} ^{60}\text{Fe}$ (NLD& γ SF)
- Intensity: $\sim 1.2\text{pps}$
- Purity: $\sim 11\%$
- Beam time: $\sim 200\text{ hr}$

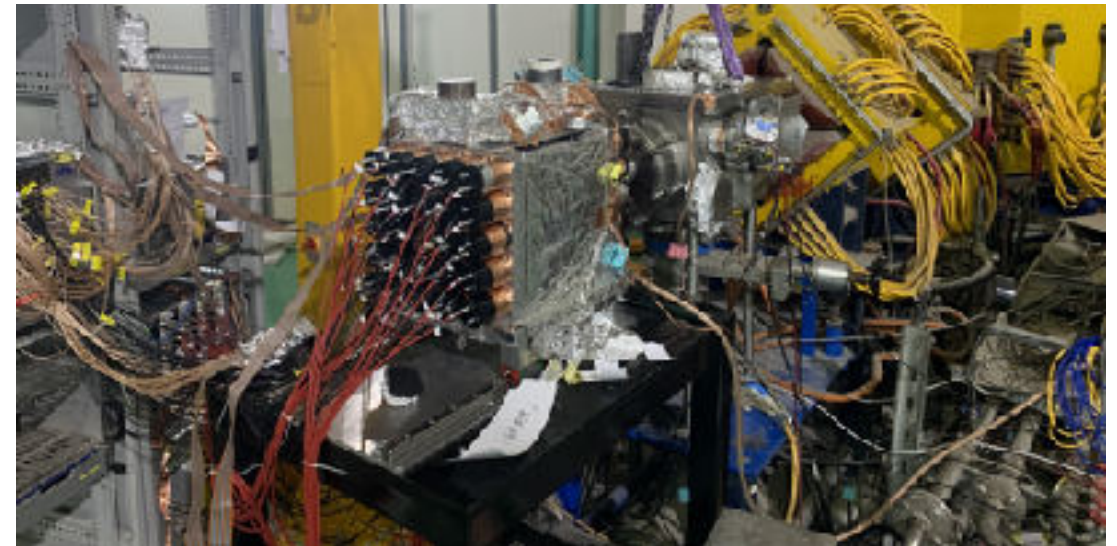
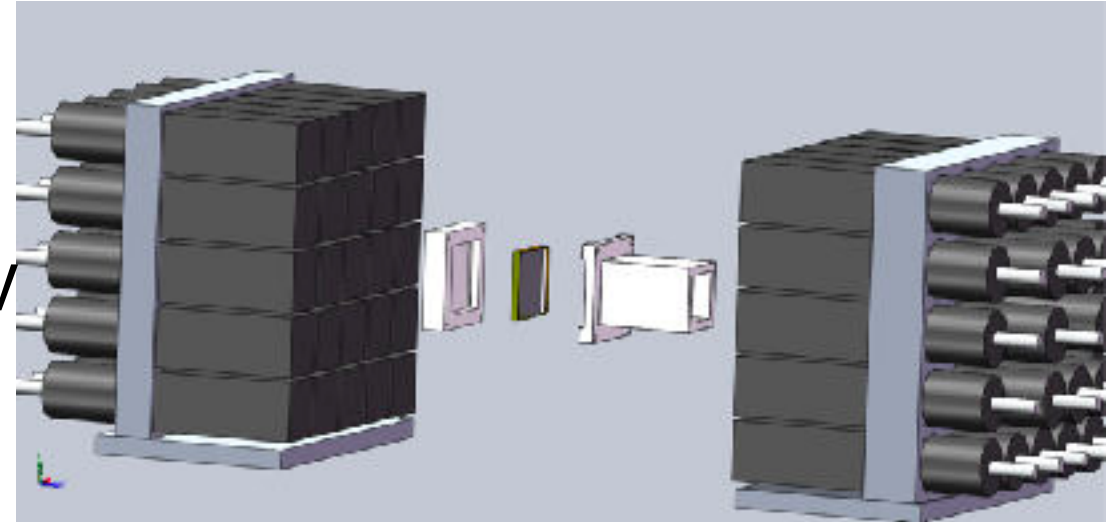




LAMBDA (**L**arge-scale **M**odular **B**GO **D**etector **A**rray)

49 BGO modules

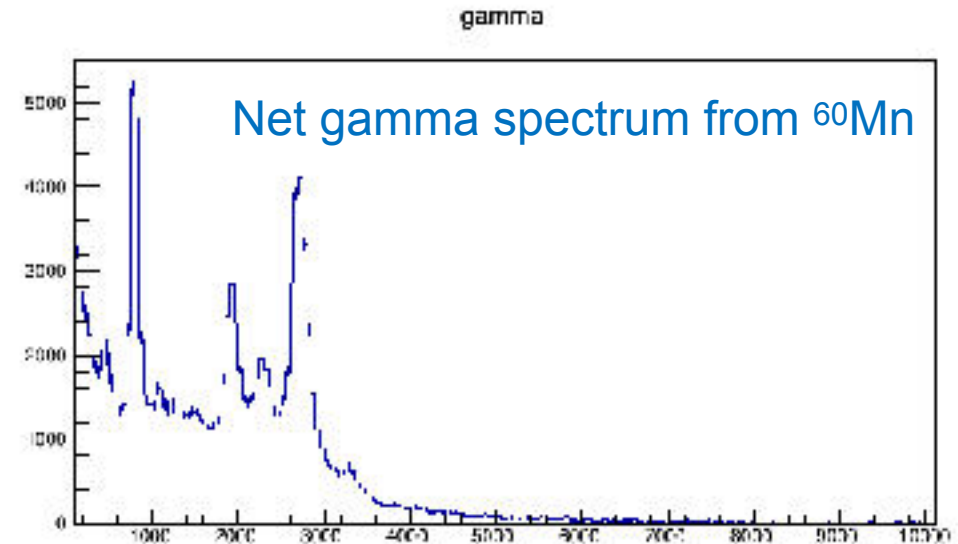
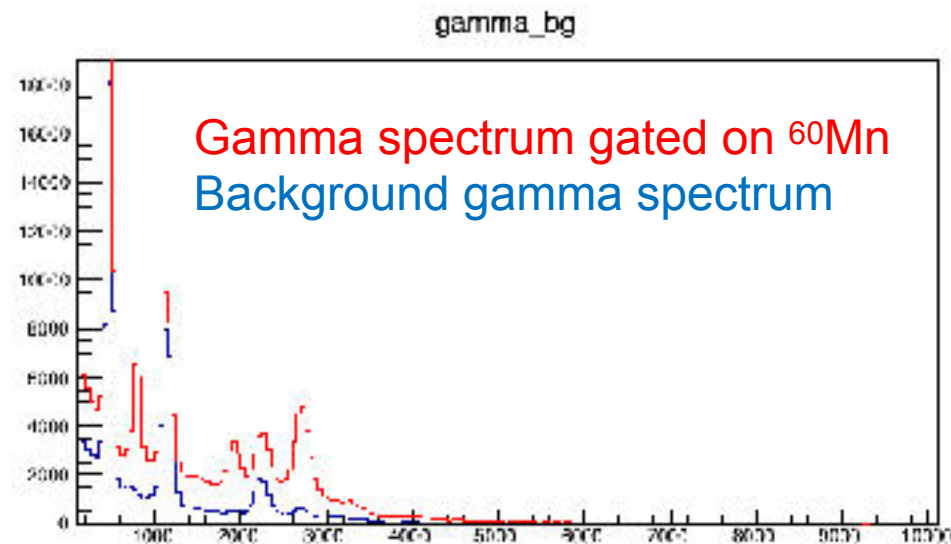
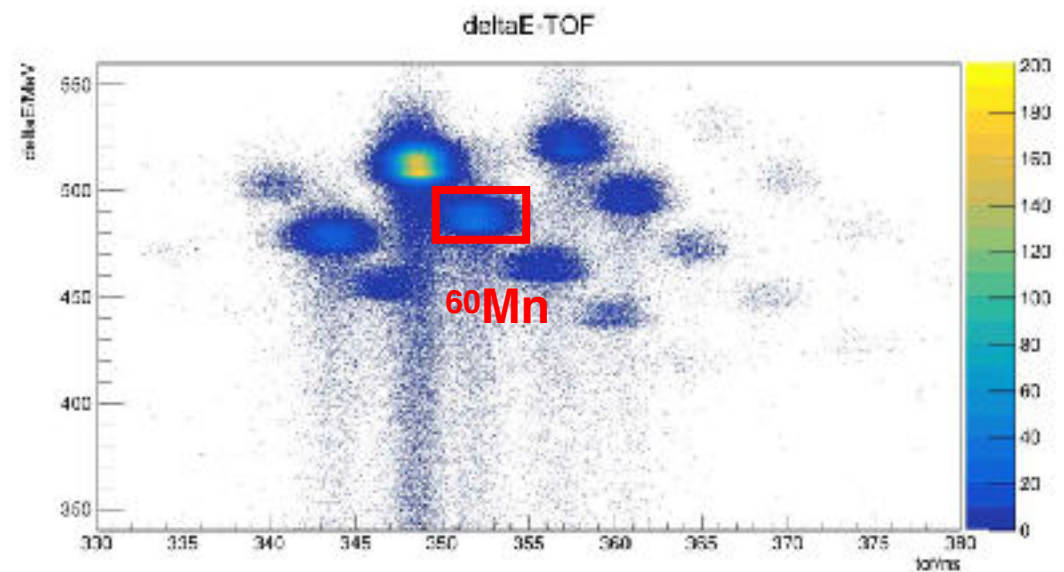
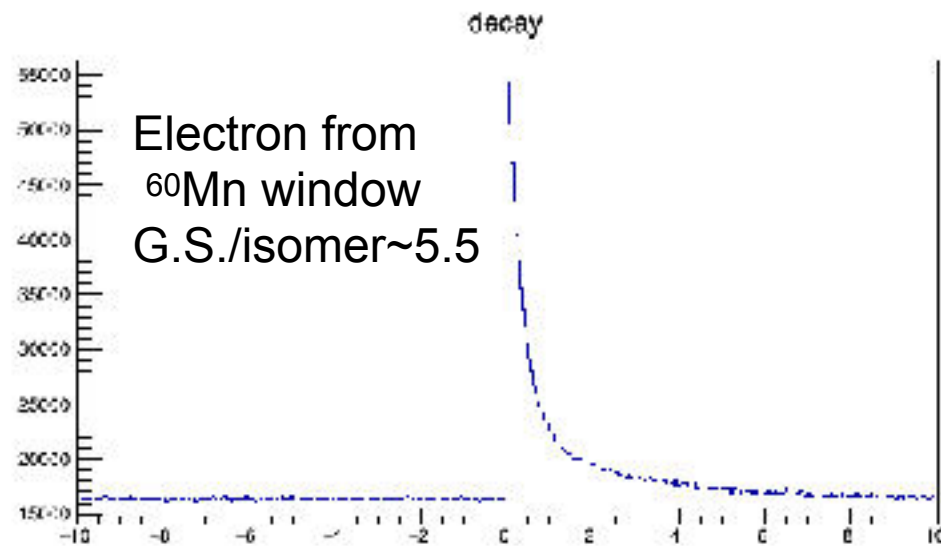
- Size: 6cm×6cm×12cm
- Energy resolution: 9.7% @ 662keV
- Total efficiency: 90% @ 1MeV 80% @ 10MeV





^{60}Mn β -Oslo experiment

Institute of Modern Physics, CAS





HIRFL



HIAF



Booster Ring:

- Circumference: 569 m
- Rigidity: 34 Tm
- Accumulation
- Cooling & acceleration

HFRS

Spectrometer Ring

Spectrometer Ring:

- Circumference: 277.2 m
- Rigidity: 15 Tm
- Electron cooler
- Stochastic cooler

45 GHz, 20 KW
ECR ion source

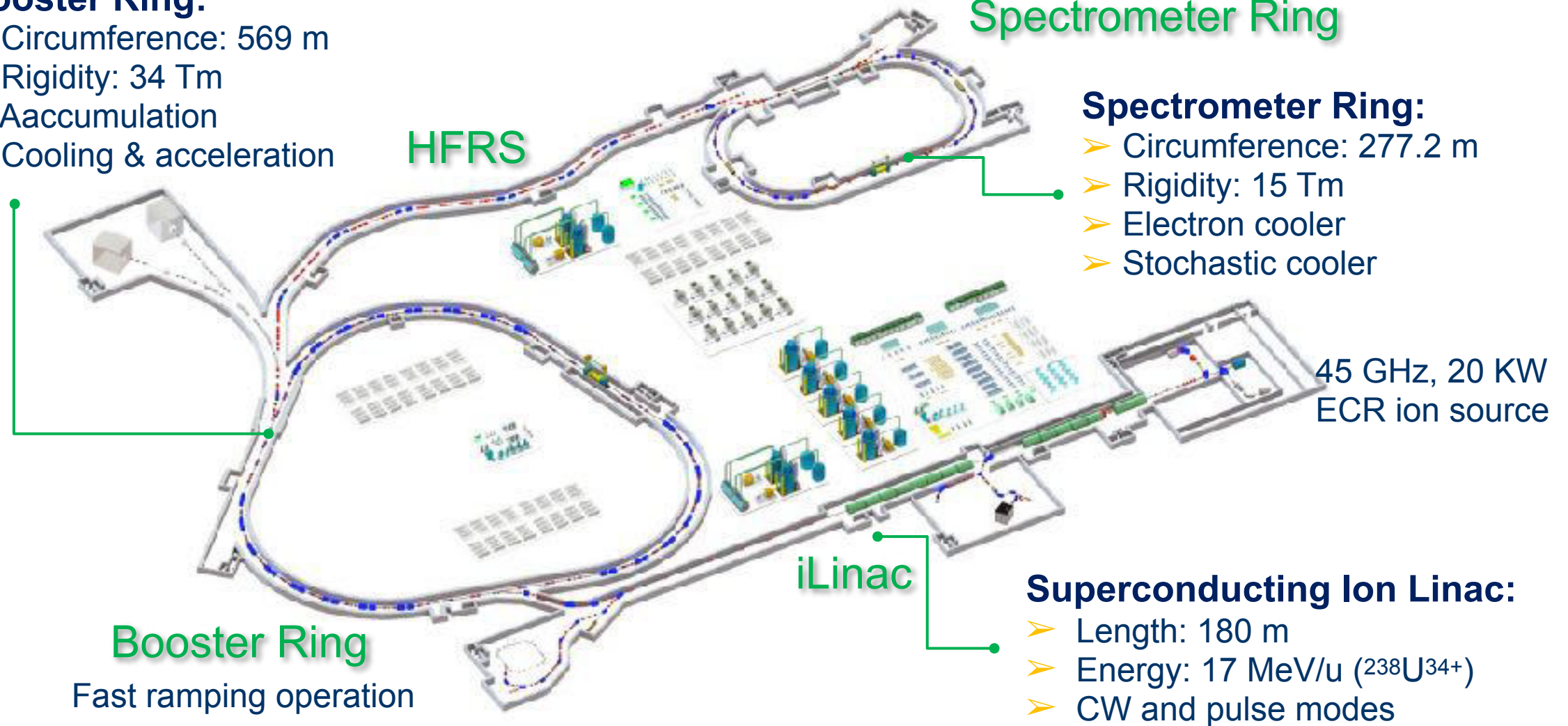
Booster Ring

Fast ramping operation

iLinac

Superconducting Ion Linac:

- Length: 180 m
- Energy: 17 MeV/u ($^{238}\text{U}^{34+}$)
- CW and pulse modes





2018, Ground-breaking Ceremony



December 2018



November 2020



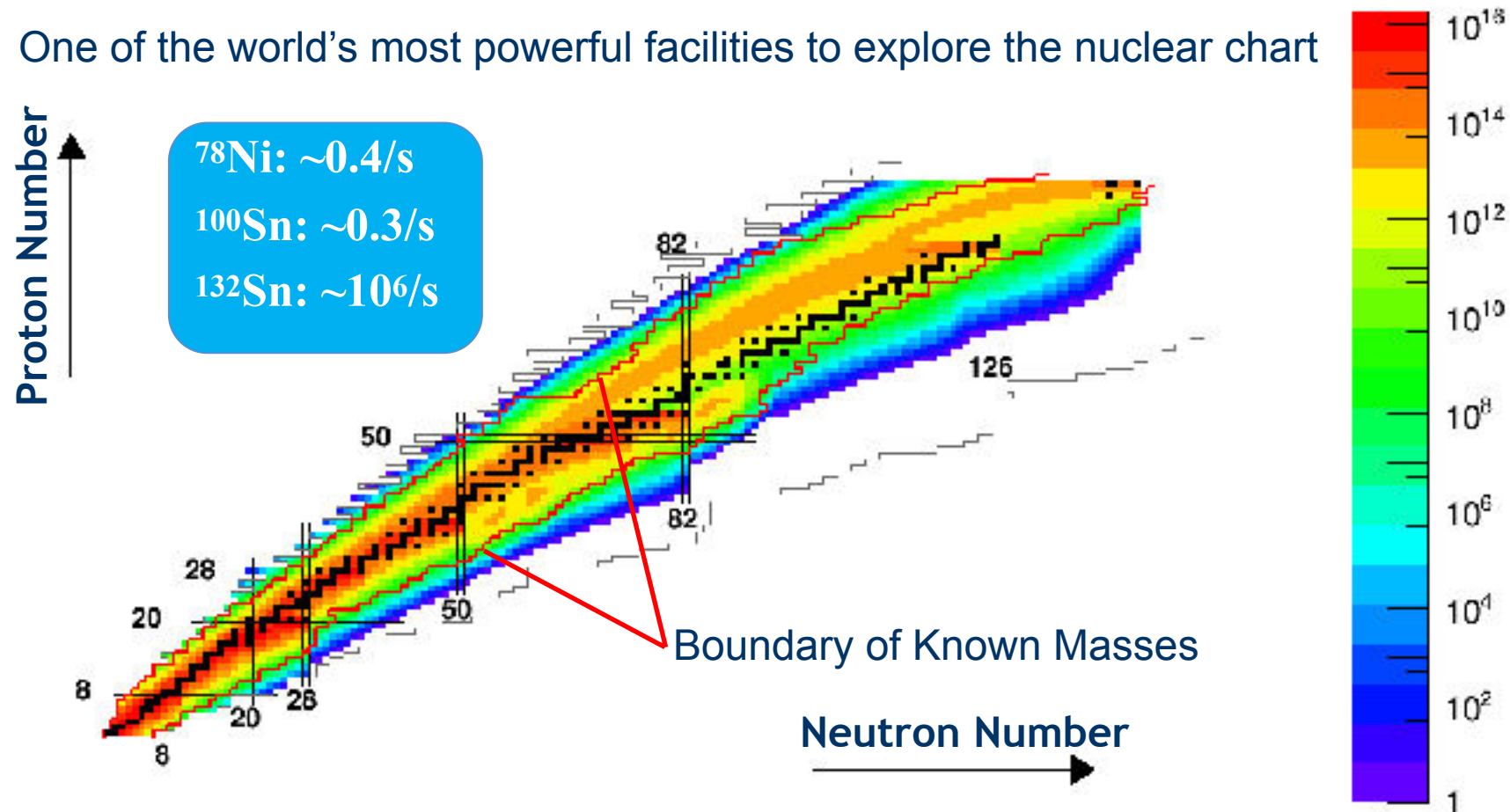
November 2023





Nuclides Available (Daily Production Yield) at HIAF

One of the world's most powerful facilities to explore the nuclear chart



Prolific sources of nuclides far away from the stability line will be provided using various reactions. The limits shown are the production rate of one nuclide per day, which enable the “discovery experiments”



Thanks

Shilun Jin
Institute of Modern Physics, CAS
jinshilun@impcas.ac.cn