## Weak rates determining the production of the <sup>205</sup>Pb cosmochronometer in AGB stars

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- <sup>205</sup>Pb is an *s*-process element produced in AGB stars the exact production rate depends on:
  - the destruction of <sup>205</sup>Pb in the <sup>13</sup>C pocket at T~90 MK by electron capture from the thermally excited 1/2<sup>-</sup> state
  - the production of  $^{205}\text{Pb}$  by bound  $\beta$  decay of  $^{205}\text{Tl}$  at T~250 MK in He flashes
- The transition between the 1/2<sup>+</sup> ground state of <sup>205</sup>Tl to the first excited 1/2<sup>-</sup> state at 2.3 keV in <sup>205</sup>Pb was measured for the first time by the E121 Collaboration with the ESR at GSI Darmstadt [1].



- <sup>205</sup>Pb bound capture rate rapidly increases with thermal population of 1/2<sup>-</sup> state at 2.3 keV.
- Bound capture rate starts to drop when atomic L- and K-orbits become depopulated.
- <sup>205</sup>Pb continuum capture rate depends on population of 1/2state and increases rapidly with electron density.
- <sup>205</sup>Tl bound β decay requires vacancies in the atomic K-shell, only found at high temperatures, also sensitive to electron

## Method

- Following the basic approach of Takahashi and Yokoi [2] the  $^{205}\text{Pb}$  electron capture and  $^{205}\text{Tl}$  bound  $\beta$  decay rates depend on:
  - the nuclear matrix elements (fixed by experiment consistent with shell model calculations of first-forbidden transition using KHH interaction)
  - the thermal population of the nuclear states, especially the 1/2<sup>-</sup> state in <sup>205</sup>Pb at 2.3 keV
  - continuum and bound electron wave functions of ions (calculated with DHF code)
  - the energies and occupations of atomic configurations distribution of charge and excitation states of ions as a function of temperature and electron density determined by Saha equation
  - the interaction of the ions with the surrounding plasma changes the atomic ionization potentials

## density.

## **Comparison with literature rates**





on calculations by Takahashi and Yokoi [3] with different assumptions about the Q-value and are extrapolated to low temperatures.

• Our results [1] improve upon Takahashi and Yokoi mainly due to the experimentally determined nuclear matrix element and an improved Q-value but also because of a more detailed treatment of atomic states and wave functions and improvements in the ion-plasma interaction.



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[2] Takahashi, Yokoi, *Nucl. Phys. A* **404**, 578 (1983)

[3] Takahashi, Yokoi, Atom. Data and Nucl. Data Tables 36, 375 (1987)

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