



The first direct measurement of the 65 keV resonance strength of the $^{17}\text{O}(p,\gamma)^{18}\text{F}$ reaction at LUNA



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Astrophysical motivations

- A precise determination of proton capture reaction rates on oxygen is mandatory to **predict the abundance ratios of the oxygen isotopes** in stellar environments where hydrogen burning is active.
- The $^{17}\text{O}(p,\gamma)^{18}\text{F}$ reaction ($Q = 5607$ keV) plays a crucial role in **AGB nucleosynthesis** and in **explosive hydrogen burning** occurring in type Ia novae.
- In the the AGB scenario ($20 \text{ MK} < T < 80 \text{ MK}$) the main contribution to the **reaction rate** comes from the $E_{\text{cm}} = 65$ keV resonance.
- At novae temperatures ($100 \text{ MK} < T < 400 \text{ MK}$) the $E_{\text{cm}} = 183$ keV resonance dominates together with the **direct capture** (DC) component (Fig. 1).

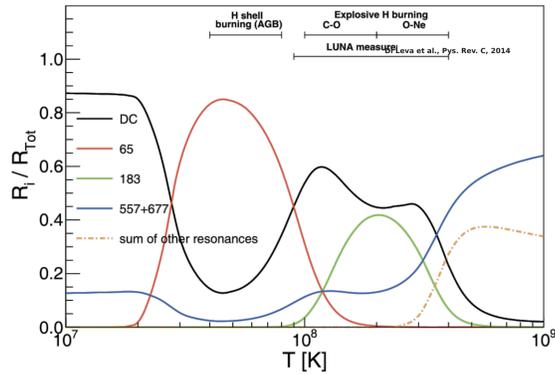


Fig.1: Fractional contribution of the reaction rate of the $^{17}\text{O}(p,\gamma)^{18}\text{F}$ [1].

State of the art

- Recently LUNA performed **precise determination of the $E_{\text{cm}} = 183$ keV resonance strength** of the $^{17}\text{O}(p,\gamma)^{18}\text{F}$ reaction [1].
- The strength of the $E_{\text{cm}} = 65$ keV resonance is presently determined only through **indirect measurements** [2-5], with the **adopted value of $\omega\gamma = (16 \pm 3)$ peV** [5].
- Current adopted value of proton capture partial width $\Gamma_p = (35.0 \pm 6.8) \times 10^{-9}$ eV [6].
- The **branching ratio** of the de-excitation of the resonance is **well known** [7].

Critical points of the measurement

- Low expected counting rate: $N = 0.31$ reactions/C.
- The environmental and beam induced **background must be reduced**.
- The knowledge of the **beam induced background** (BIB) is crucial for the evaluation of the signal.
- The **target** must be **well characterized** in thickness and stoichiometry.
- The **detection efficiency** must be **maximized**.
- Data analysis focused on a regime with $S/N \approx 1$.

Target characterization

- Anodisation of **tantalum backings** in water enriched in ^{17}O and ^{18}O .
- Runs on top of $^{17}\text{O}(p,\gamma)^{18}\text{F}$ $E_{\text{cm}} = 183$ keV resonance: \rightarrow **^{17}O abundance**.
- Periodic scans of the $^{18}\text{O}(p,\gamma)^{18}\text{F}$ $E_{\text{cm}} = 144$ keV resonance: \rightarrow **^{18}O abundance**; \rightarrow **target degradation**.
- At the energy of interest the **plateau remains unchanged** within the uncertainties after $Q > 20\text{C}$ cumulated on target (Fig.4).
- Produced targets $\sim 20\text{keV}$ and $\sim 50\text{keV}$ thick (@ $E_p=80$ keV, Fig.5).

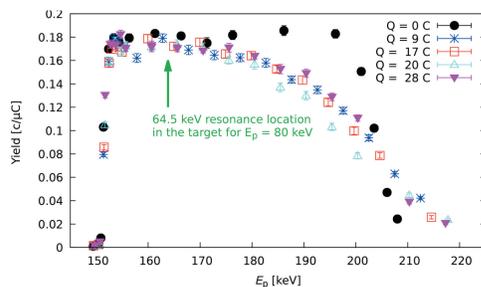


Fig.4: Example of a target scan.



Fig.5: Left: $\sim 20\text{keV}$ thick target. Right: $\sim 50\text{keV}$ thick target.

Experimental setup

- Situated at **LNGS** below a 1400m thick overburden of rock:
 - \rightarrow **Muon-induced background reduced by 6 orders of magnitude**;
 - \rightarrow **Neutron background reduced by 3 orders of magnitude**.
- LUNA **400kV electrostatic accelerator** can provide stable and intense ($<I>=200 \mu\text{A}$), proton or alpha beams with high energy resolution (30 eV) [8].
- 4π **BGO detector** segmented in 6 crystals, with high efficiency (74% @ 661 keV, Fig. 2).
- Aluminum target chamber and target holder** to reduce absorption.
- Lead + borated polyethylene shielding** for further background reduction of a factor 4.27 ± 0.09 in the ROI ($5200\text{keV} - 6200\text{keV}$) with respect to only lead shielding (Fig.3)[9].



Fig.2: Detail of the shielding.

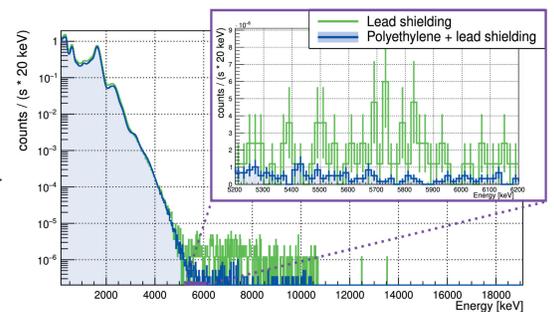


Fig.3: Environmental background reduction [9].

Measurement campaigns

- 4 campaigns** in 2021-2022.
- 420C** on top of 65keV resonance on ^{17}O targets.
- 300C** on top of 65keV resonance on **UPW targets to monitor BIB**.

Simulation of the setup

- Geant4 simulations **optimized** on well-known spectra of ^{137}Cs , ^{60}Co , $^{14}\text{N}+p@270\text{keV}$ (Fig.6).
- Difference with experiment $\lesssim 3\%$ in all three cases.
- Simulations used to **determine detector efficiency**.
- DC contribution** under the resonance determined simulating branching ratios of [1].

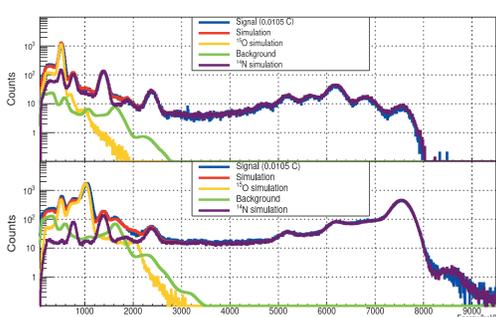


Fig.6: $^{14}\text{N}+p@270\text{keV}$ simulation [9].

Data analysis

- Measurement objective are **the net counts of the resonance**, but **p+D sum peak falls in the same ROI** \rightarrow **BIB contribution due to D contamination** (Fig.7).
- Use the detector **segmentation** and knowledge on ^{18}F **branching ratio** to **isolate resonance events** (Fig.8):
 - \rightarrow Select events in **sum peak**;
 - \rightarrow Among these select events with **primary energy reading in one crystal**.
- Apply the same gates on runs on **UPW targets** to subtract random coincidences.
- Apply the same gates on **DC simulation** to subtract its contribution under the resonance.

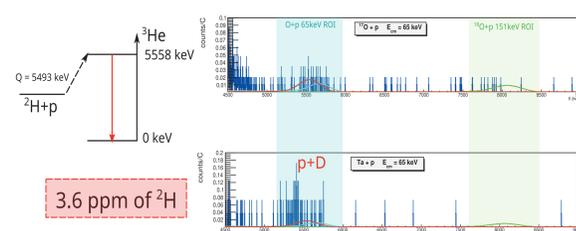


Fig.7: Deuterium contamination.

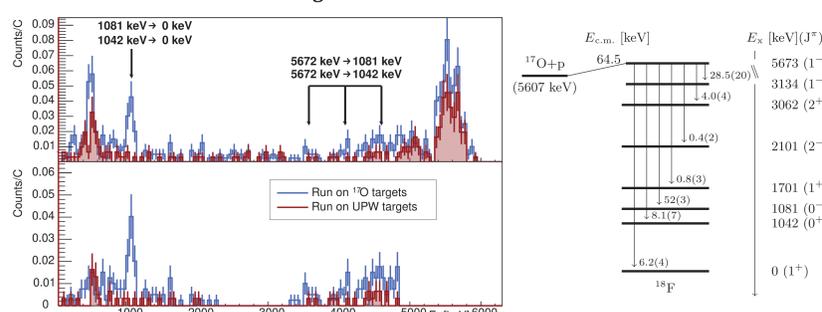


Fig.8: Right: gate on sum peak of $^{17}\text{O}+p$ $E_{\text{cm}}=65\text{keV}$ resonance [10]. Left: level scheme of ^{18}F with primary branching ratios, as reported in [7].

Results

Reference	$\omega\gamma_{p,\gamma}$ [peV]	$\omega\gamma_{p,\gamma}^{\text{bare}}$ [peV]	Γ_p [neV]	Γ_p^{bare} [neV]
Previously adopted value	16(3)[5]	40(7)[6]	35(6)[6]	
Present work [10]	34(8)	30(6)	39(9)	34(8)

- First direct measurement** of the resonance strength, about a **factor 2 higher** than the values reported in literature.
- The Γ_p was calculated, in excellent **agreement with previous LUNA result** [6].
- Lowest strength** value ever measured directly.
- Technical paper on setup **published** [9].
- Results paper **published** on PRL vol.133 issue 5 [10].

References

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