Development of the Charge-Exchange Oslo Method and its First Application to Constrain (n,γ) Cross Sections

N.D. Pathirana^{1,2}, H. Berg^{1,2}, B. Gao³, M. Guttormsen⁴, A.C. Larsen⁴, C. Maher^{1,2}, S. Noji¹, A. Riley⁵, A. Schiller⁴, A. Spyrou^{1,2}, S. Uthayakumaar¹, R.G.T. Zegers^{1,2}, and members of the e15112 experiment collaboration ¹Facility for Rare Isotope Beams, ²Michigan State University, ³Institute of Modern Physics, ⁴University of Oslo, ⁵Ursinus College



Introduction

- The development of the Charge-Exchange Oslo (CE-Oslo) method is important for constraining multiple inputs for nucleosynthesis simultaneously;
 - Nuclear level density (NLD)
 - γ-ray strength function (γSF)
 - Gamow-Teller strengths
 - β -delayed neutron decay probabilities
- The long-term goal is to pursue (p,n+γ) experiments in inverse kinematics to extract such constraints for unstable nuclei.

Motivation

1. Charge-Exchange (CE) Reactions

• CE reactions are characterized by the exchange of a proton and a neutron between the target nucleus and the projectile nucleus.

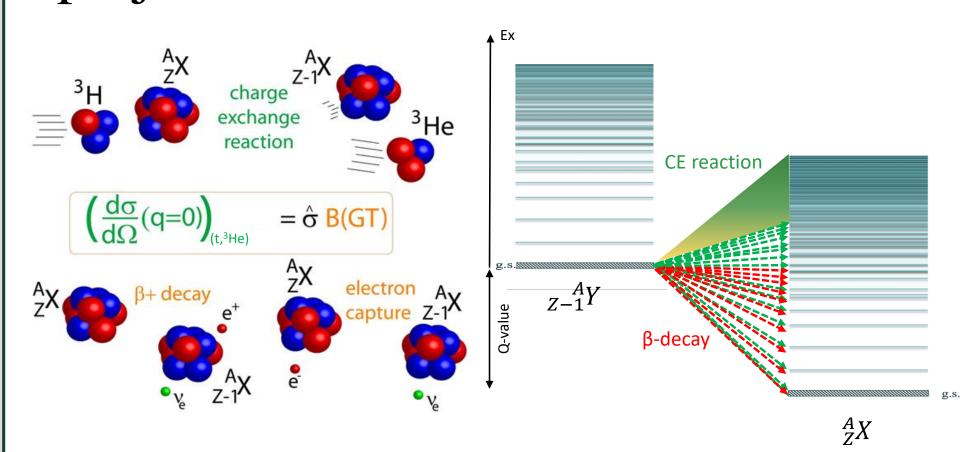


Fig. 1 Charge-Exchange reactions and comparison with EC and β -decay

2. CE-Oslo Method

• Similarly to Oslo Method and β -Oslo Method, we now develop the CE-Oslo Method for CE reactions.

Experimental Details

- To test the CE-Oslo method, data from a previous CE experiment ⁹³Nb(t, ³He+γ)⁹³Zr is taken, which has been run at 115 MeV/u in coincidence with S800 and GRETINA at NSCL.³
- Using the particle- γ coincidence data, the NLD of 93 Zr, γ SF of 93 Zr, and (n,γ) cross sections of 92 Zr are extracted with the Oslo method package and the TALYS reaction code. 1,4,5,6,7



Fig. 2 ⁹³Nb(t, ³He+γ)⁹³Zr experimental setup

Conclusion

- This is the first time that:
- ⁹²Zr(n,γ)⁹³Zr cross sections are measured using the Oslo Method.
- The CE-Oslo method is applied to constrain the (n,γ) cross sections.
- The Oslo method is applied to constrain (n,γ) cross sections using the GRETINA γ -coincidence data with S800.

Results for ⁹³Nb(t, ³He+γ)⁹³Zr

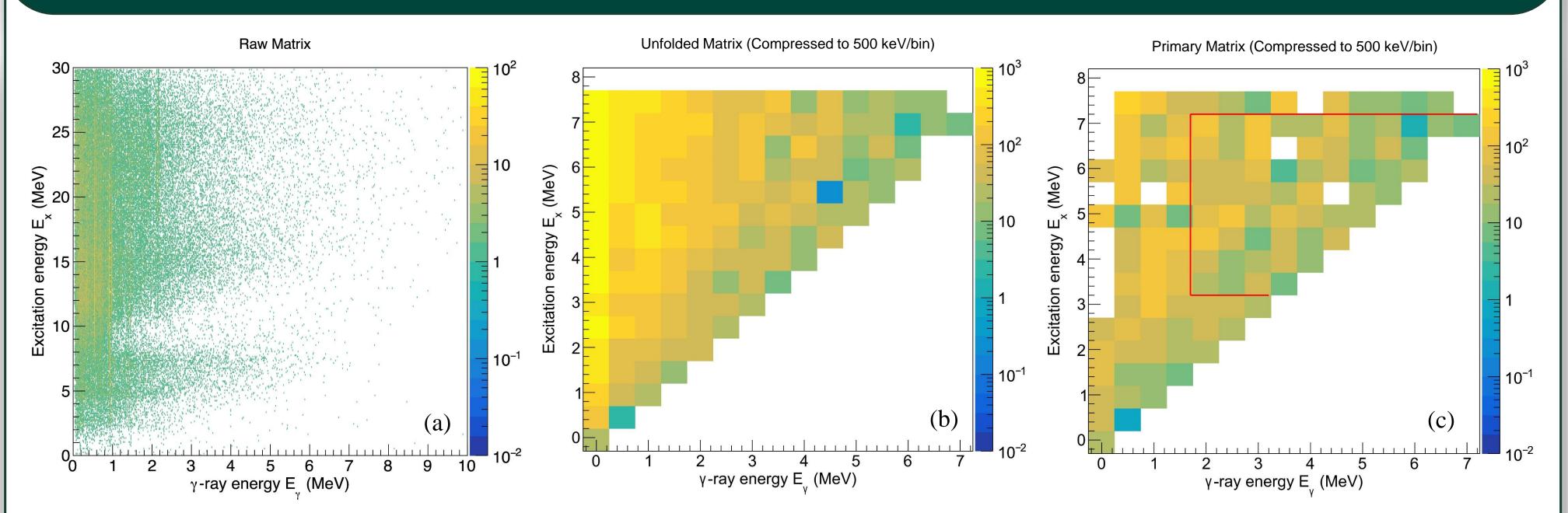


Fig. 3 The (a) raw matrix (b) unfolded matrix and, (c) primary matrix

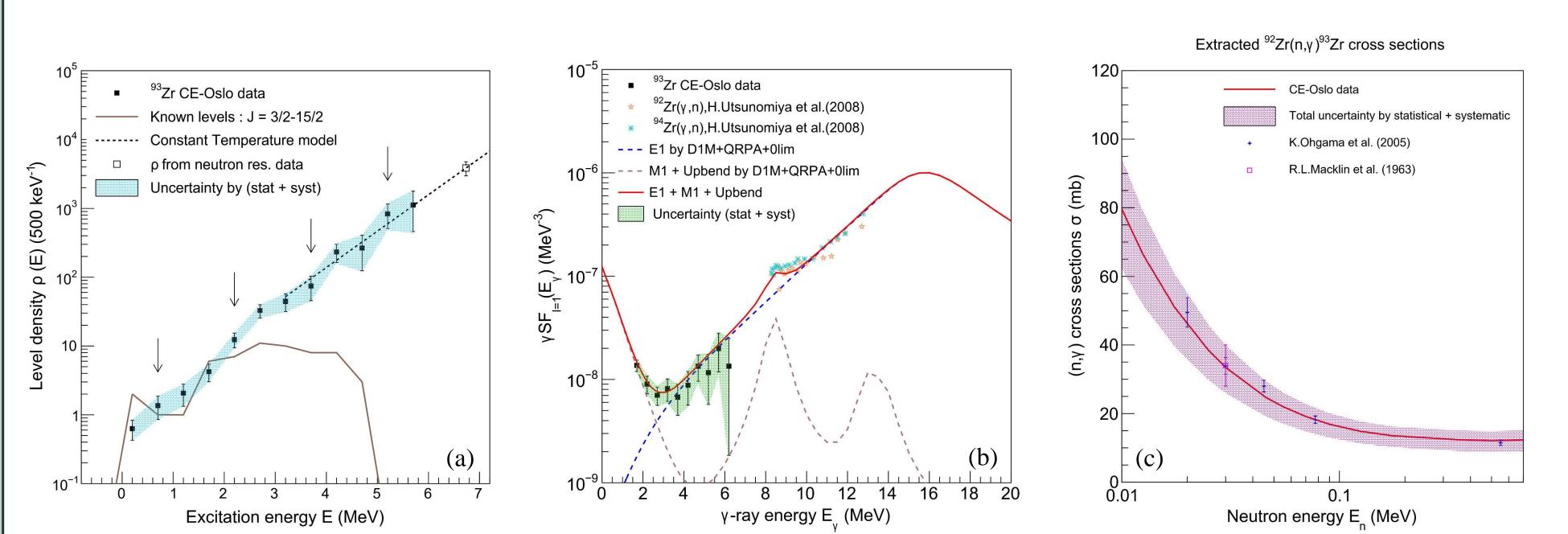


Fig. 4 The (a) NLD of 93 Zr (b) γ SF of 93 Zr and, (c) 92 Zr(n, γ) 93 Zr cross sections

Acknowledgement

I would like to thank my Ph.D. advisor Prof. Remco Zegers for his guidance and our Charge-Exchange group for their support, Prof. Artemis Spyrou and her SuN group for providing details and the NSF for providing funding to make this possible. This work is supported by the National Science Foundation PHY-2209429, "Windows on the Universe: Nuclear Astrophysics at the FRIB".

References

- [1] M.Guttormsen, Nucl. Instrum. Methods Phys. Res.A 255 (1987)
- [2] A. Spyrou et al., Phys. Rev. Lett. 113, 232502 (2014)
- [3] B.Gao et al., PhysRevC.101.014308 (2020)
- [4] M.Guttormsen, Nucl. Instrum. Methods Phys. Res. A 374 (1996)
- [5] A Schiller et al., Instrum. Methods Phys. Res.A 447 (2000)
- [6] A. C. Larsen et al., Phys. Rev. C 83, 034315 (2018)
- [7] Arjan Koning et al., Eur. Phys. J. A (2023) 59:131





