

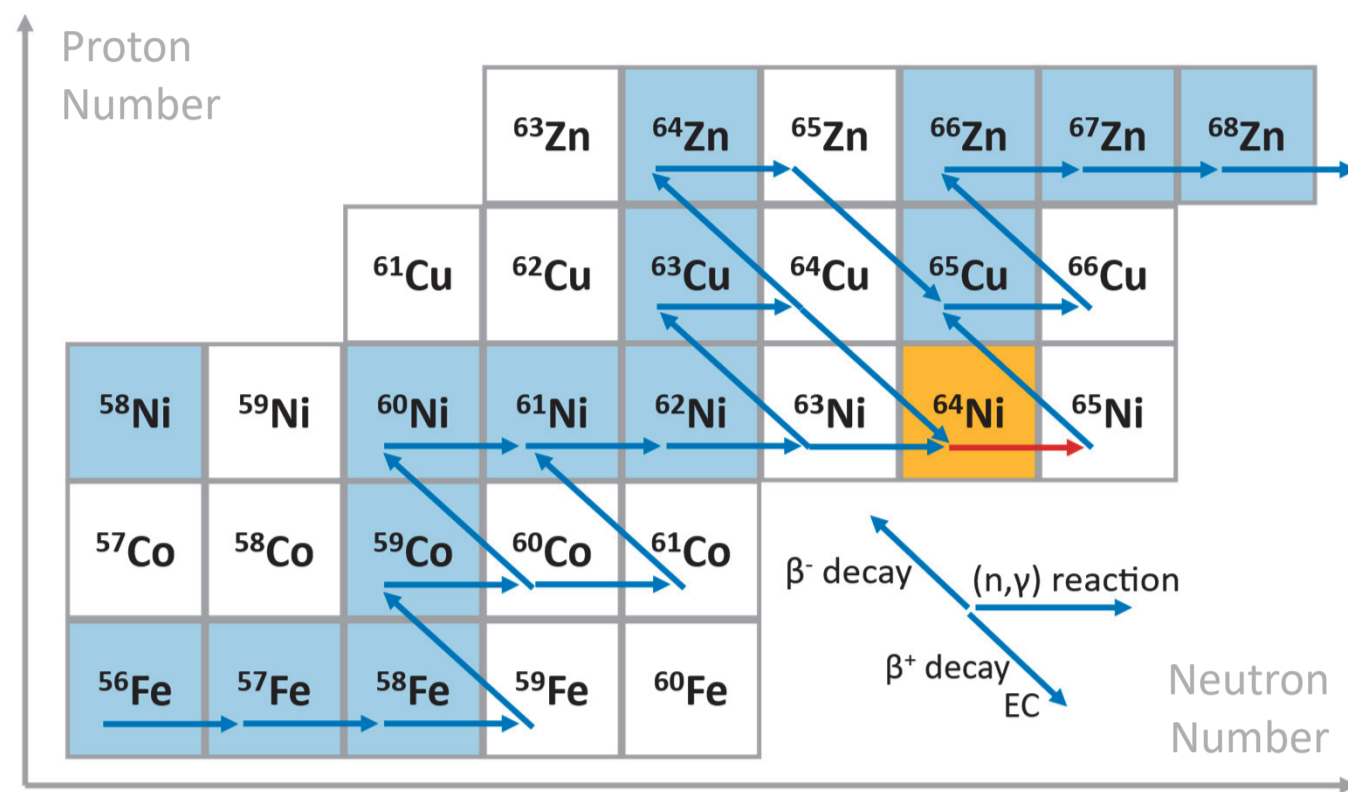
Michele Spelta<sup>1</sup>, Giuseppe Tagliente<sup>2</sup> and the n\_TOF collaboration [1]

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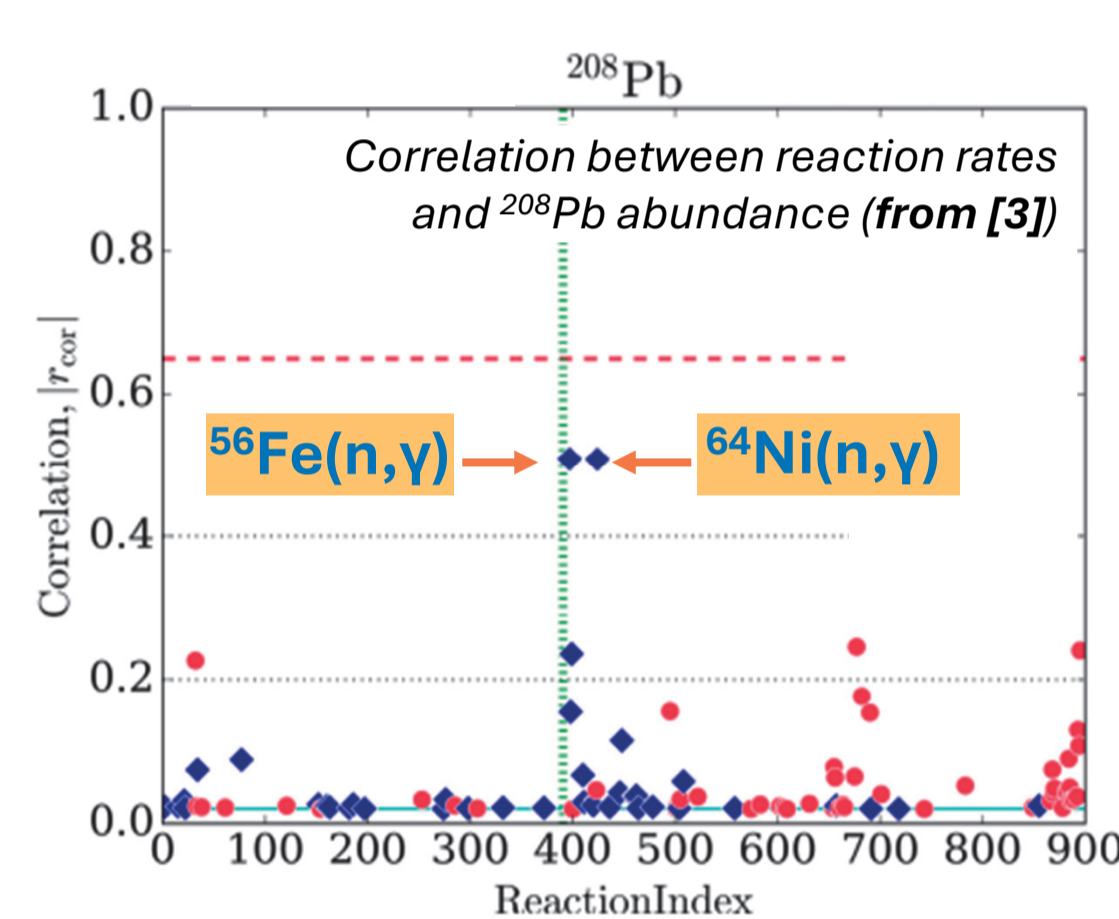
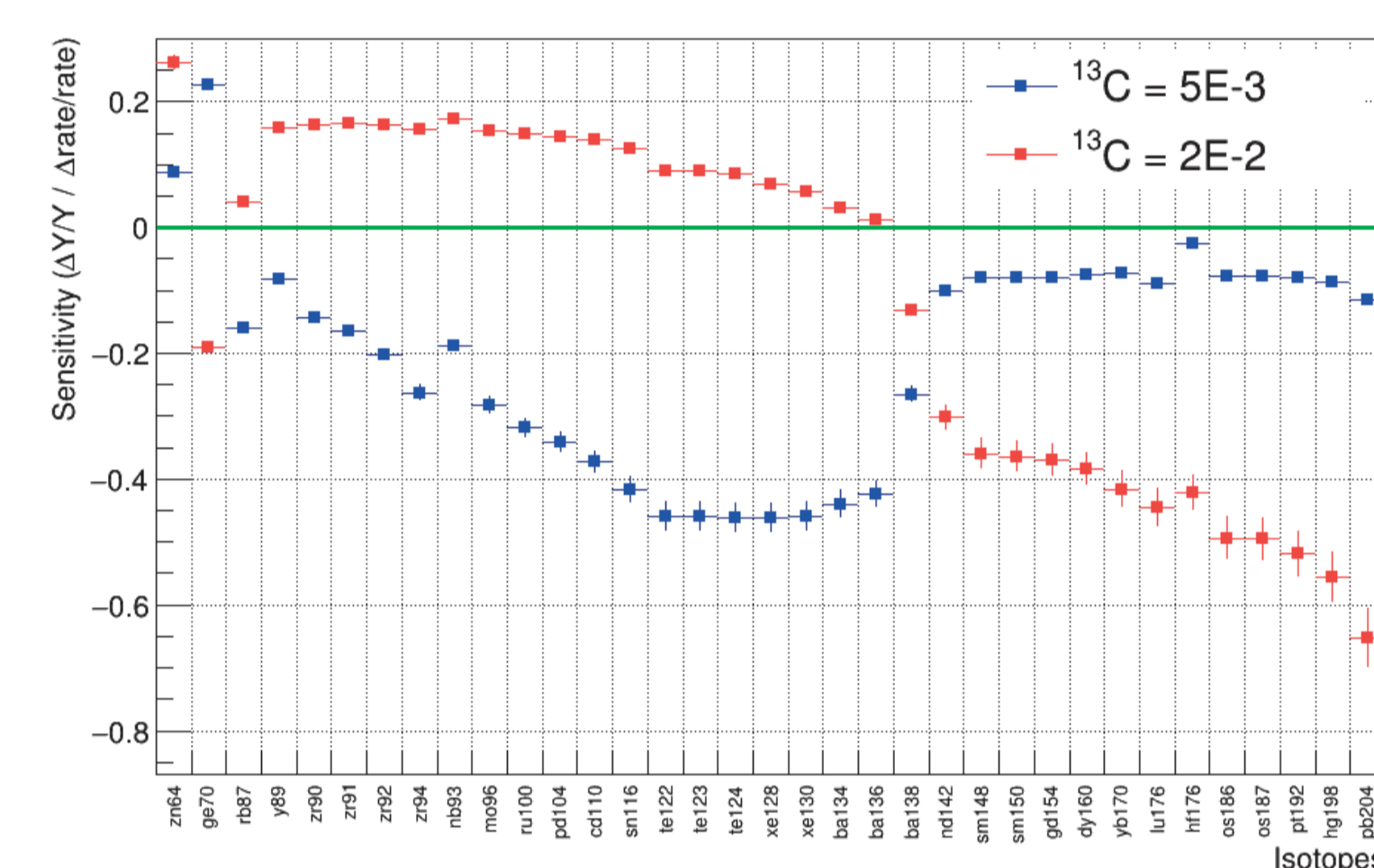
## Motivations

$^{64}\text{Ni}$  is one of seeds of the astrophysical **s-process**, a series of neutron capture reactions and beta decays responsible for the nucleosynthesis of half of the nuclei heavier than Iron. Its neutron capture cross section is therefore crucial:

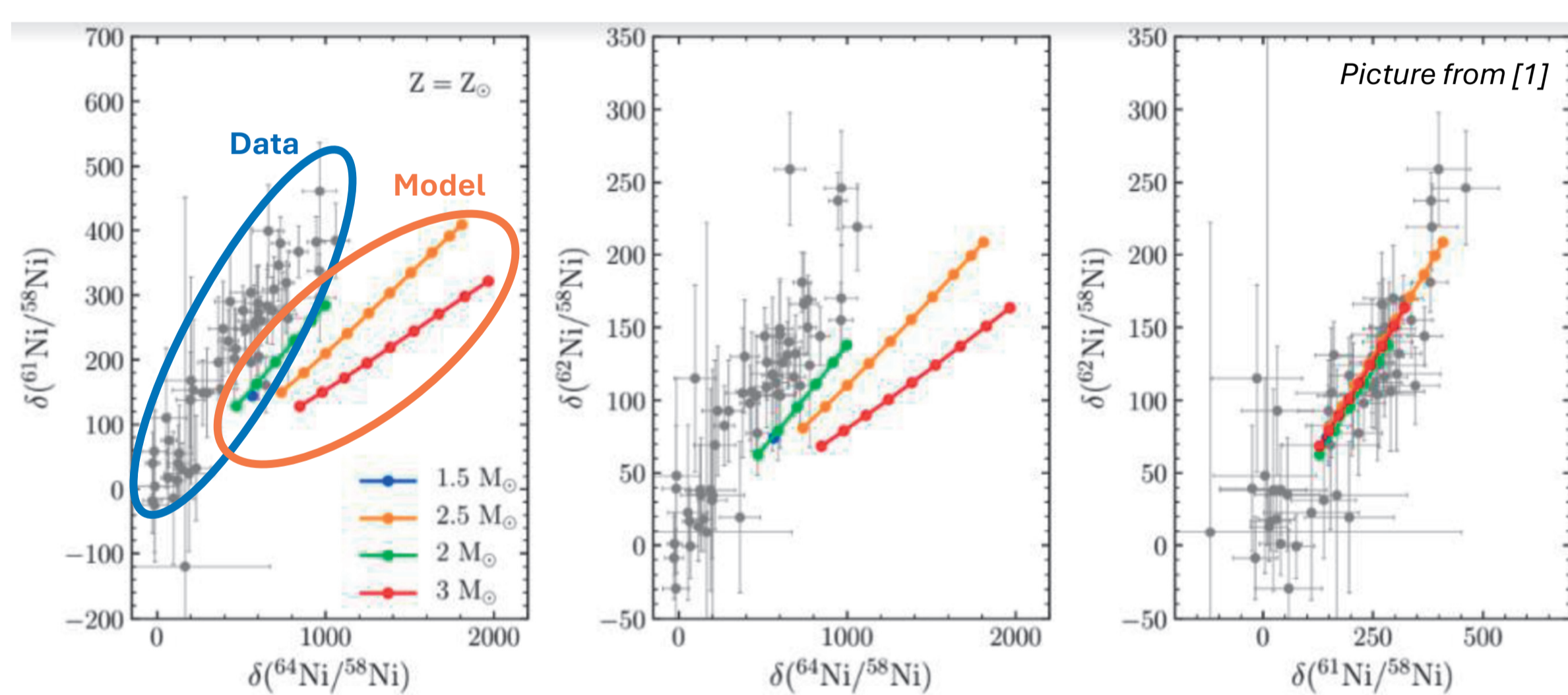


to accurately simulate the weak s-process in massive stars ( $A < 90$ ), where uncertainties on a cross section are known to propagate on the uncertainties of the abundances of all the heavier nuclei [2]

to accurately simulate the main s-process in AGB stars ( $A > 90$ ), where the neutron capture rate of  $^{64}\text{Ni}$  has been found to importantly affect the abundances of many isotopes later synthesized in the process [3]



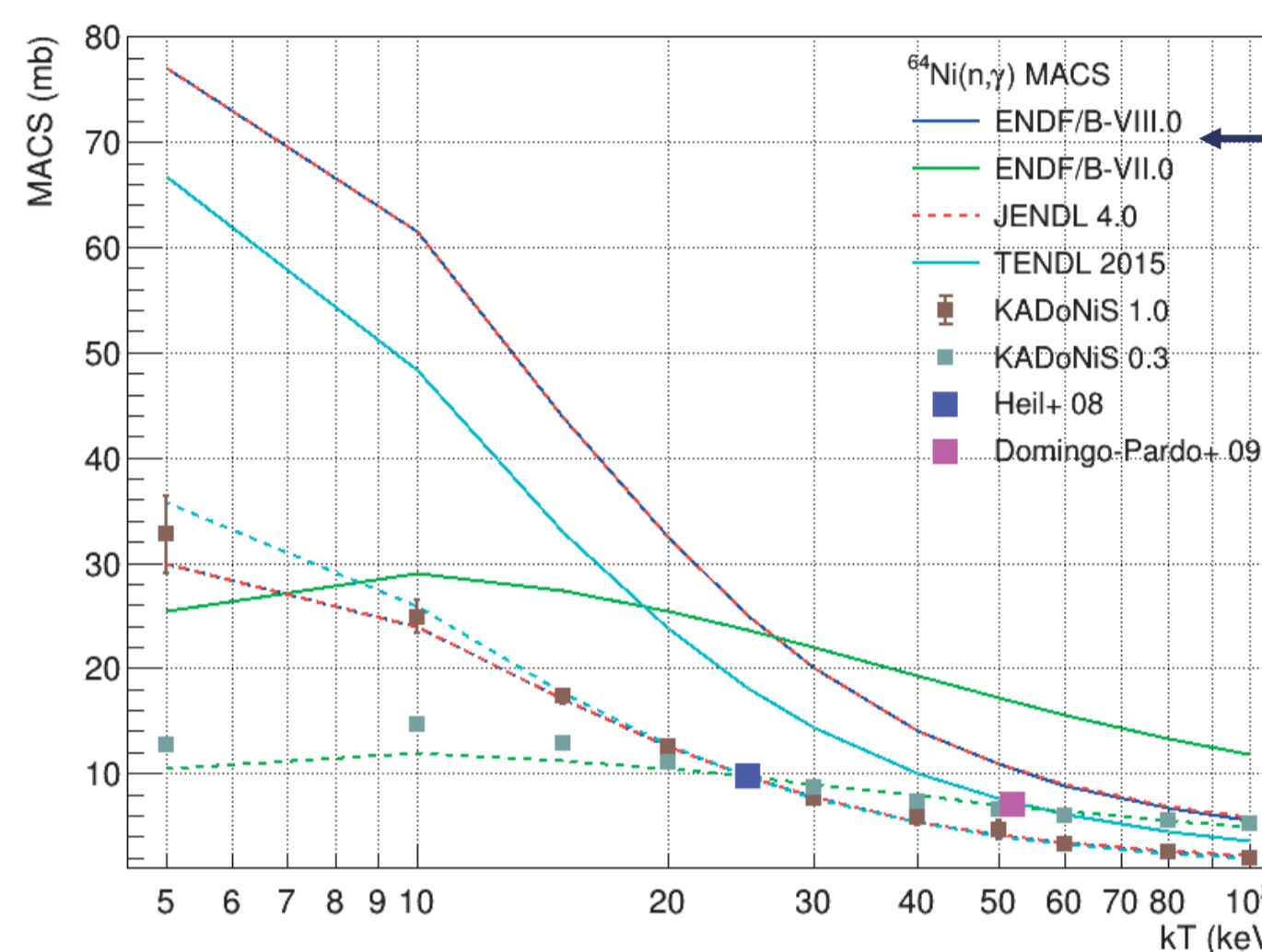
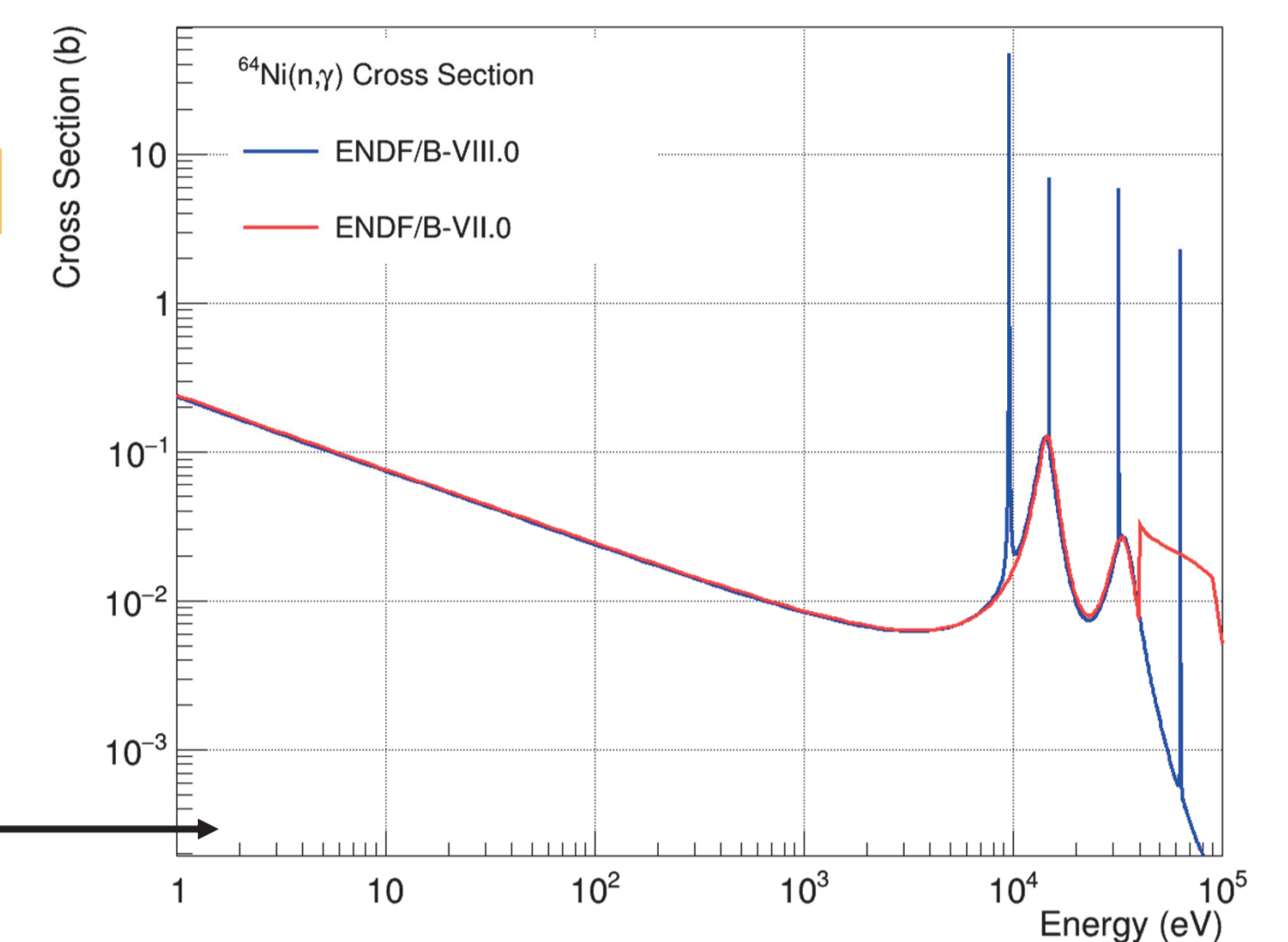
to possibly explain the discrepancy observed in presolar SiC grains between the measured abundance of  $^{64}\text{Ni}$  and the predictions from a recently introduced model for AGB stars [4]. Since the model is working for most of the isotopes, the discrepancy could be due to an incorrectness of the input neutron capture rate of  $^{64}\text{Ni}$  rather than an incompleteness of the model.



## Previous measurements

Time-of-flight measurements of the  $^{64}\text{Ni}$  capture cross section reported in literature are **few, discrepant and incomplete**.

The most recent measurement [5] was performed in the '80s and it found important discrepancies with respect to previous measurements [6], but without reporting p-wave resonances. Different nuclear data libraries also show huge discrepancies.



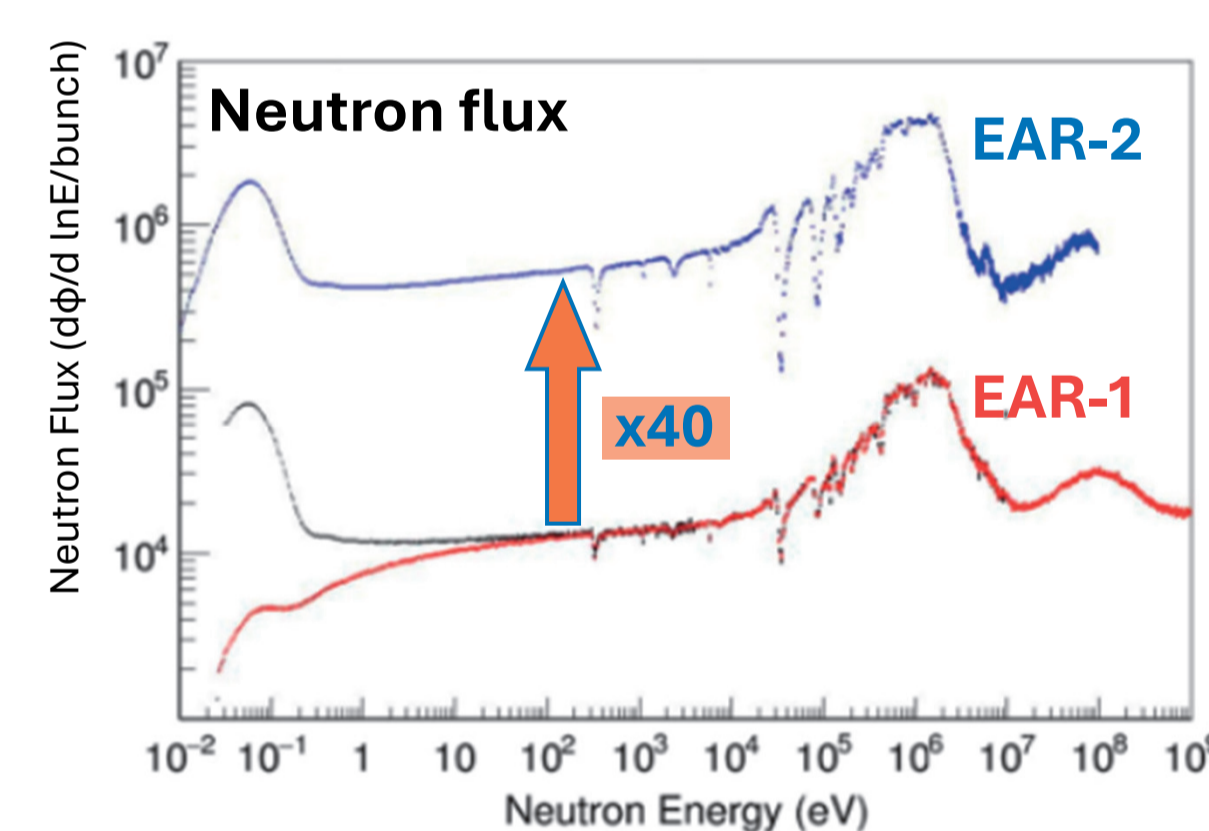
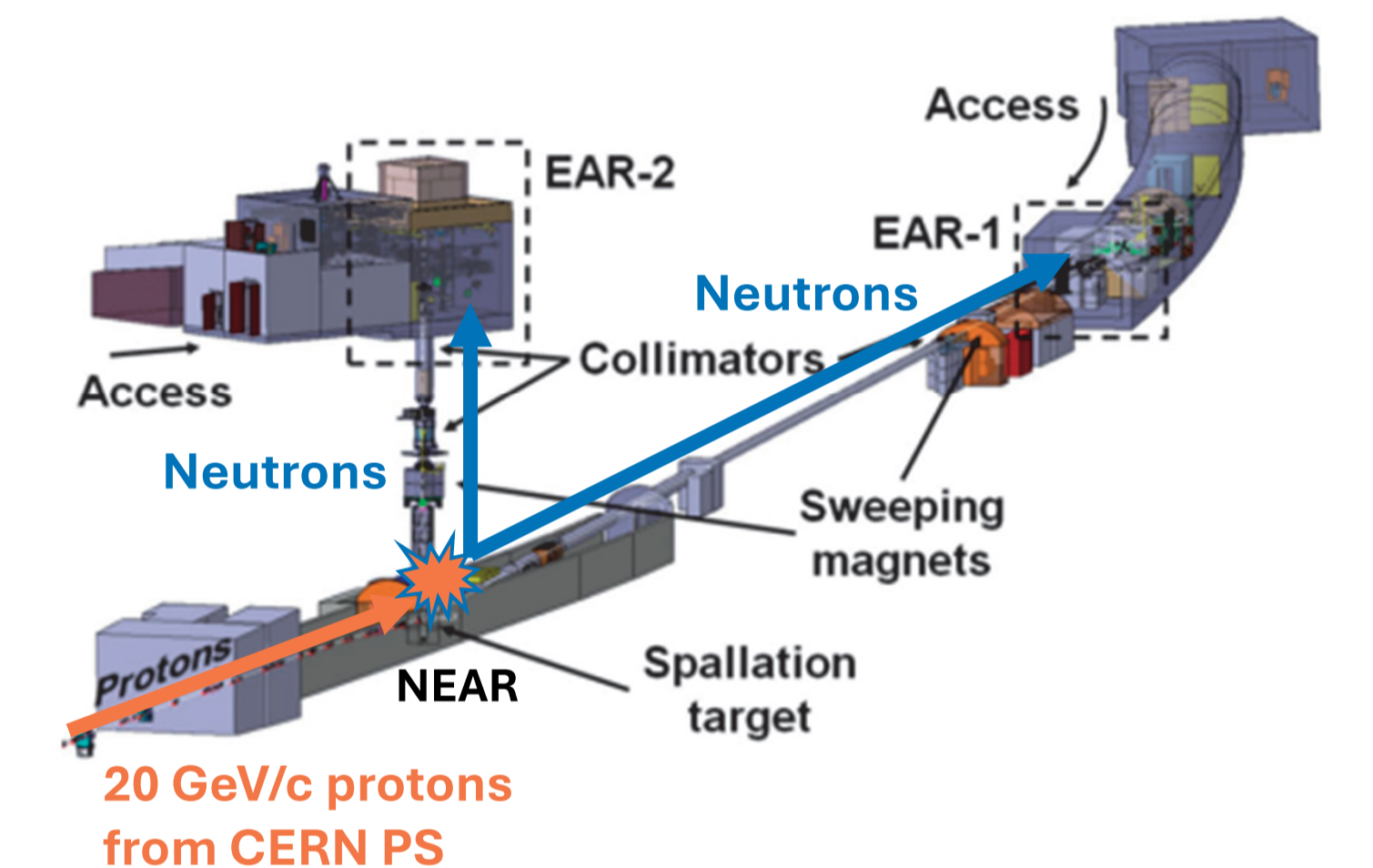
**Maxwellian Averaged Cross Sections** from activation measurements [7, 8] disagree with values computed from nuclear data libraries.

However, their energy extrapolation relies on the discrepant time-of-flight measurements that currently lead to discrepancies larger than a factor 2 at low energy.

## The n\_TOF facility

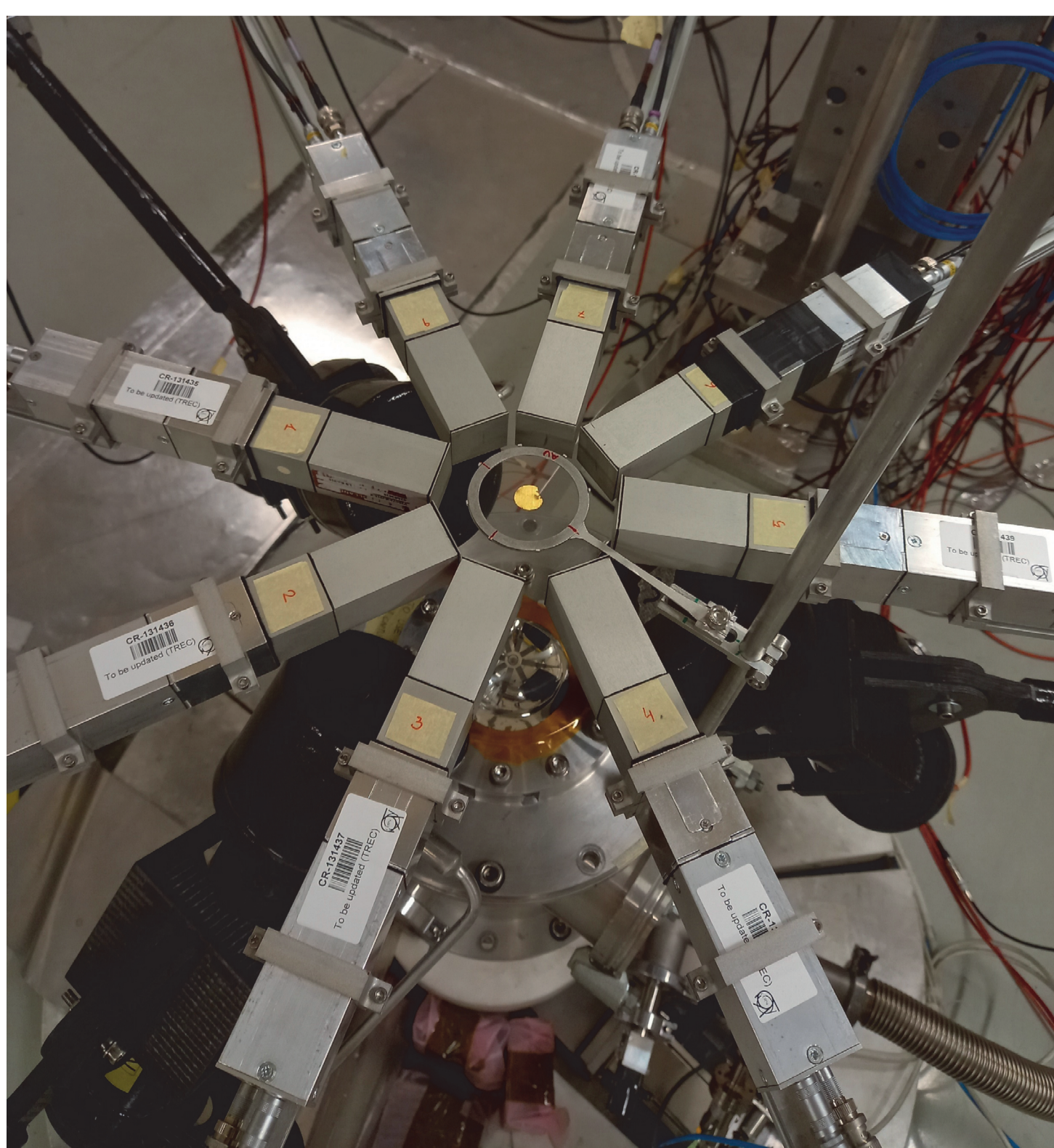
The **n\_TOF facility** is a pulsed white neutron spallation source at CERN for time-of-flight measurements of neutron-induced reaction cross sections. It is characterized by:

- High neutron flux
- Wide neutron energy range
- Excellent energy resolution



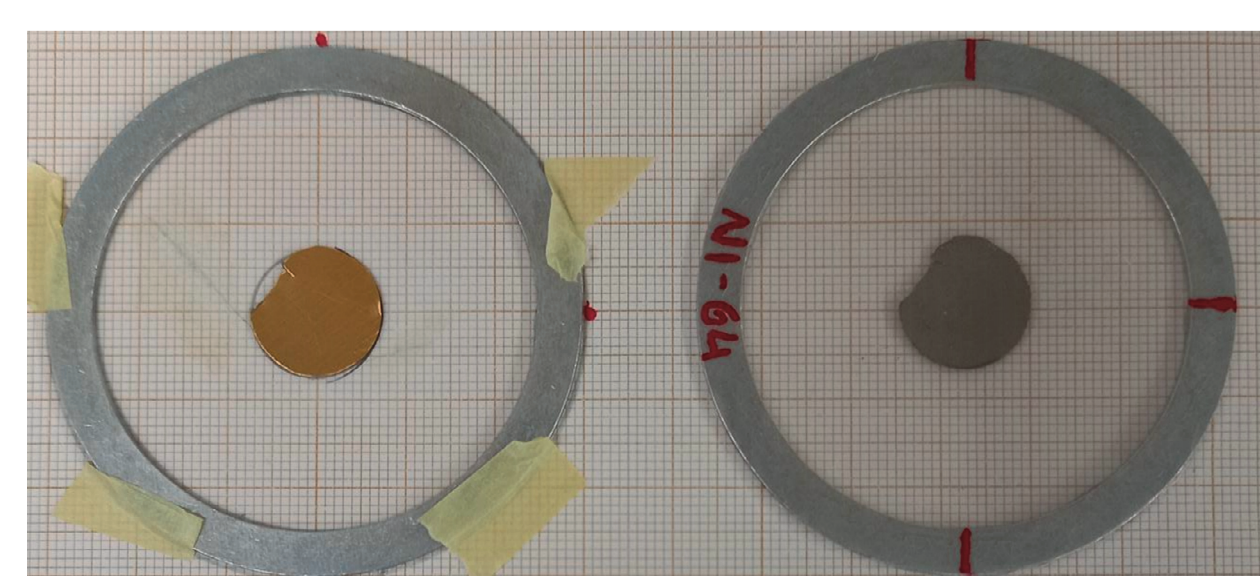
The construction of **EAR-2**, closer to the spallation target and characterized by a higher neutron flux, has enabled the measurement of samples with high radioactivity and/or small mass (down to  $\mu\text{g}$ ). As a result, the costly  $^{64}\text{Ni}$  could also be studied.

## Measurement and Data Analysis



The new time-of-flight measurement was performed at n\_TOF EAR-2 in 2023 using:

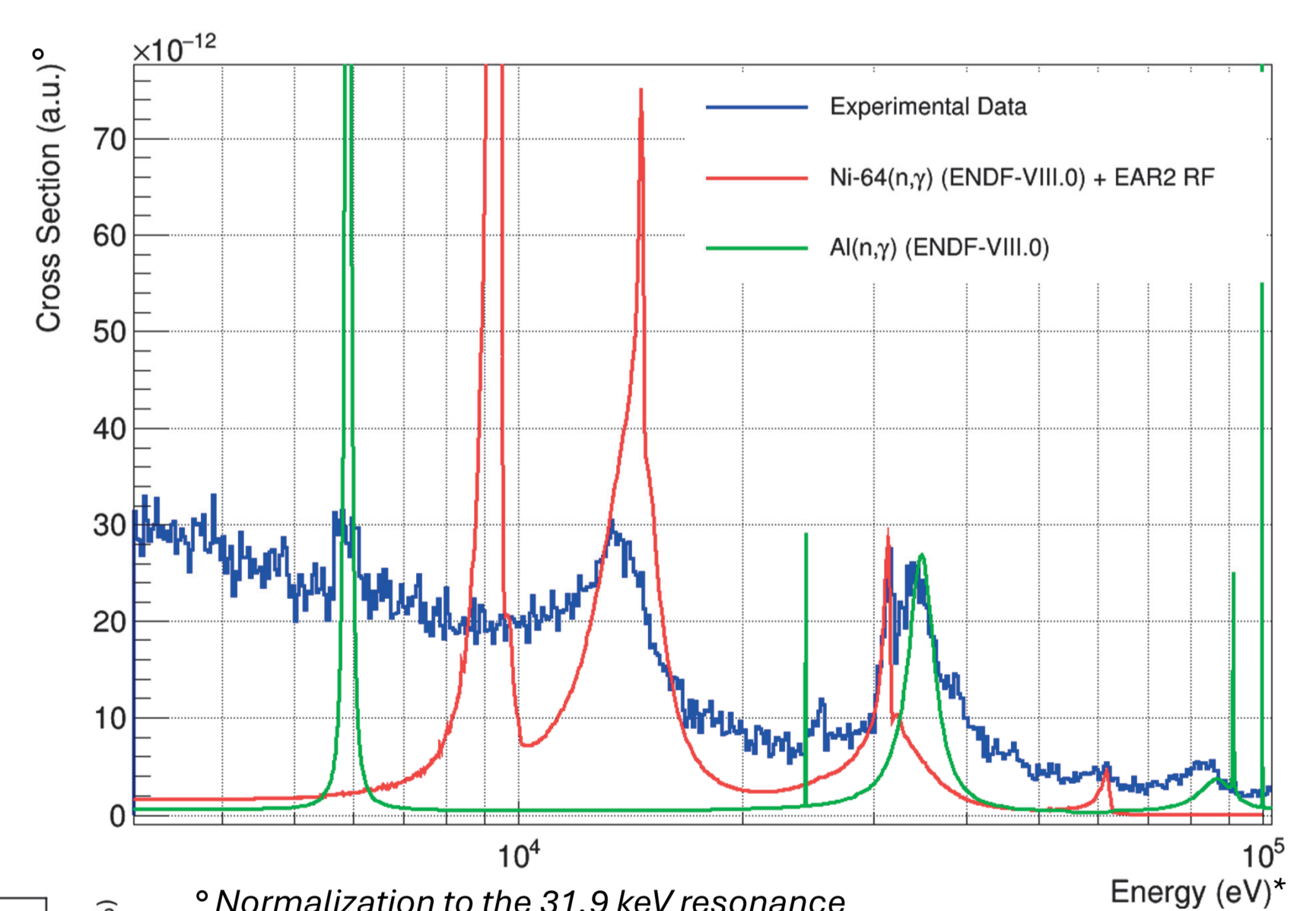
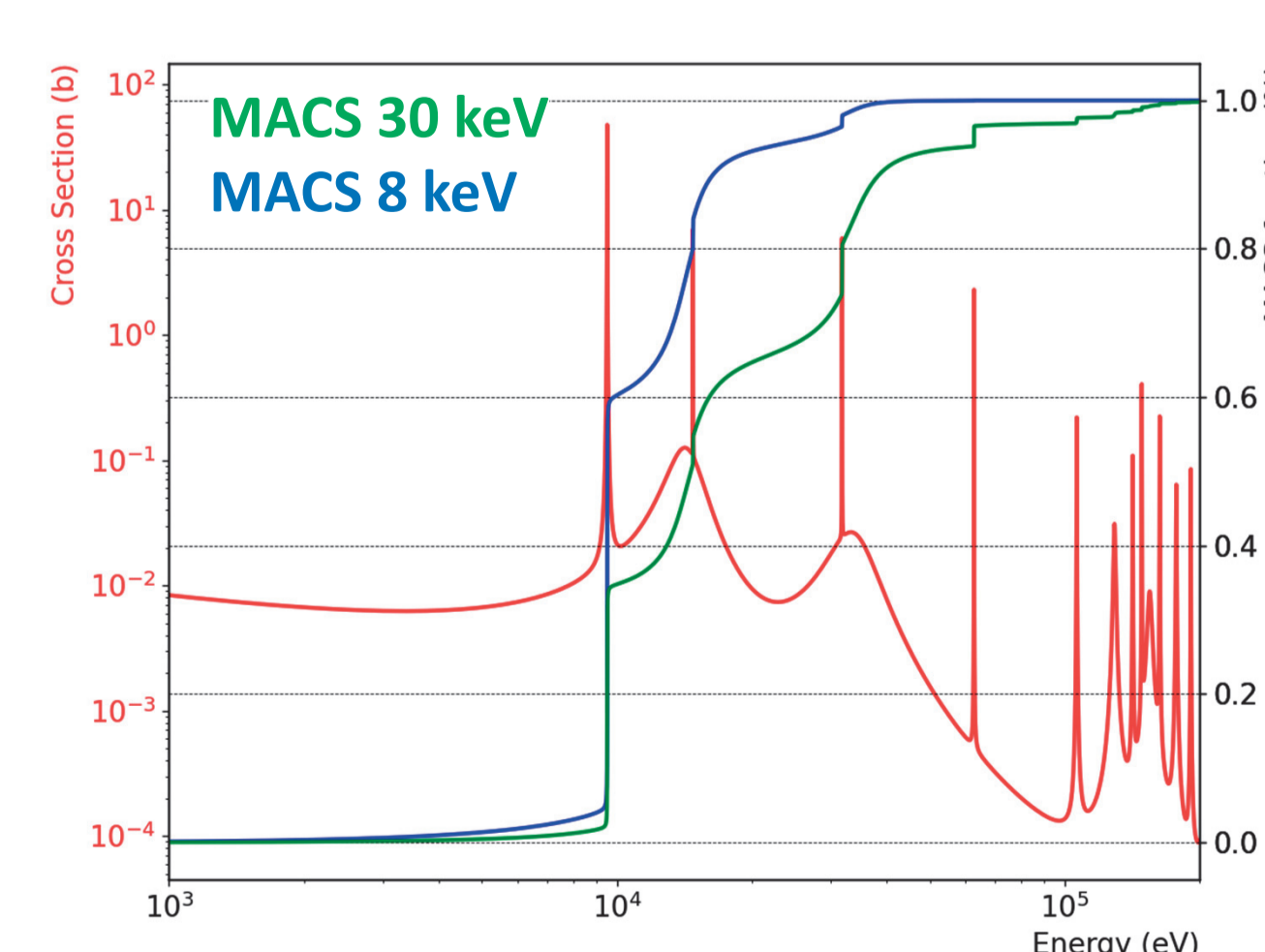
- State-of-the-art  **$\text{C}_6\text{D}_6$  segmented liquid scintillators (sTED)**, optimized for the higher counting rate of EAR-2
- $^{64}\text{Ni}$  enriched sintered sample (99.5% enrichment)



Data will be analyzed using the **Total Energy Detection** principle combined with the **Pulse Height Weighting Technique** and will be normalized with respect to the 4.9 eV **Gold** resonance.

## Preliminary Results

- No resonance at 9.52 keV**
- Evidence for only some of the expected p-wave resonances (31.9 keV, 62.8 keV\*)
- Effects due to neutron-capture reactions on **Aluminum** (contamination or neutron-scattering effect)



\* Normalization to the 31.9 keV resonance

\* ToF-to-energy conversion is preliminary

**MACS at low energy is expected to be smaller** than the value reported in the latest nuclear data library releases

**Analysis is currently ongoing...**

## References:

- [1] G. Tagliente et al., CERN-INTC-2022-033/INTC-P-208 (2022)
- [3] G. Cescutti et al., MNRAS 478, 4101-4127 (2018)
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- [2] M. Pignatari et al., ApJ 710, 1557 (2010)
- [4] D. Vescovi et al., ApJ Lett. 897, 25 (2020)
- [6] H. Beer et al., Nucl. Phys. A 240, 29 (1975)
- [8] C. Domingo-Pardo et al., AIP Conf. Proc. 230 (2009)