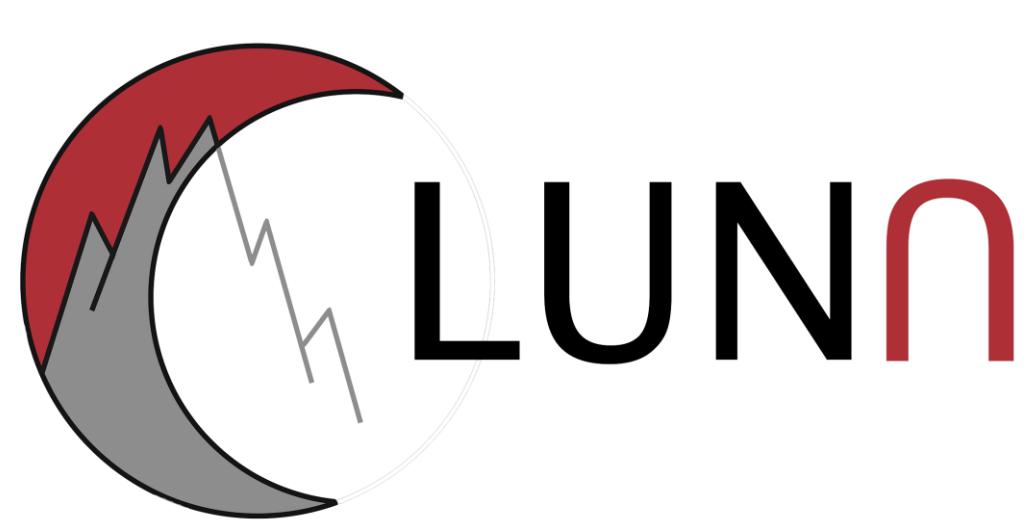


The SHADES Project:

Underground Measurement of the Low Energy $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$ Cross Section



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For the LUNA Collaboration



1. Motivation

- **Massive stars** ($M > 8\text{M}_\odot$) and Asymptotic Giant Branch (AGB) stars synthesise neutron-rich isotopes [1,2] across $A \sim 60 - 90$ and $90 - 209$, respectively.
- Path proceeds via slow neutron-capture process (**s-process**), where reactions $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$ and $^{13}\text{C}(\alpha, n)^{16}\text{O}$ are the dominant neutron sources.
- $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$ reaction rate is **weakly constrained at astrophysical temperatures** $100 - 300$ MK.

2. Goals

- Measure $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$ across $E_\alpha = 600 - 886$ keV ($E_{\text{cm}} = 507 - 750$ keV)
- Remeasure $E_\alpha = 832$ keV resonance, search for $E_\alpha = 635$ keV resonance

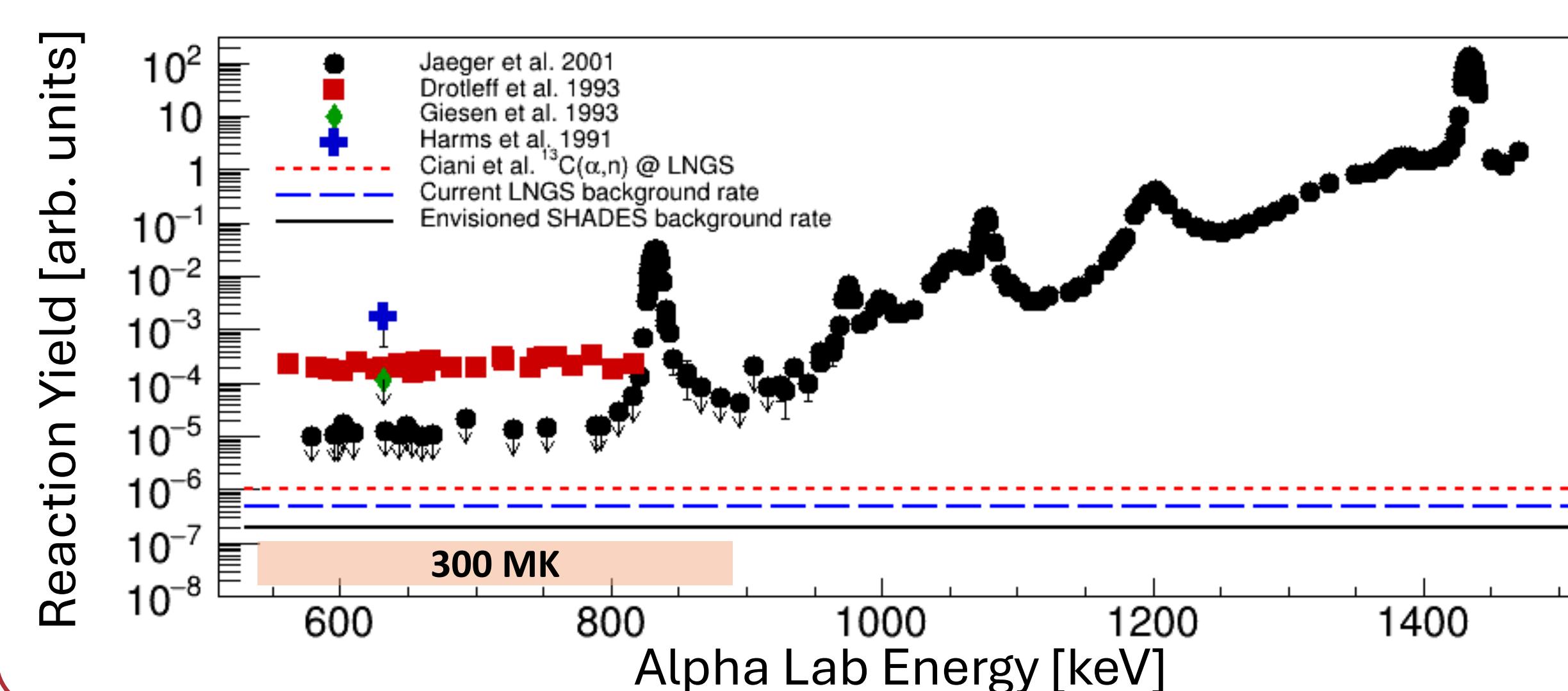
 $(E_{\text{thresh}} = 561$ keV)

Fig 1. State-of-art of $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$ cross-section data. Data points from [3-6]. Black line shows envisioned sensitivity at SHADES + Belotti Facility.

3. Experimental Setup

- INFN – Laboratori Nazionali del Gran Sasso
 - **Ultra-low** background rates
- Accelerator: **Belotti Ion Beam Facility** [7], 3.5 MV
- Beam: $^{4}\text{He}^{1+}$, $I \sim 300 - 500$ e μ A
- Target: **Enriched ^{22}Ne gas** ~ 2 mbar, 20 cm, 10^{18} atom/cm 2
- DAQ: COMPASS & Caen V1725SB/D digitisers 250 MSample/s
- Detector array: **SHADES** [8]
 - Scintillator- ^{3}He Array for Deep-underground Experiments on the **S**-process
 - 12x EJ-309 liquid scintillators \rightarrow **neutron/gamma separation & n-moderation**
 - 18x ^{3}He proportional counters \rightarrow **reaction yield**
 - Borated polyethylene shielding

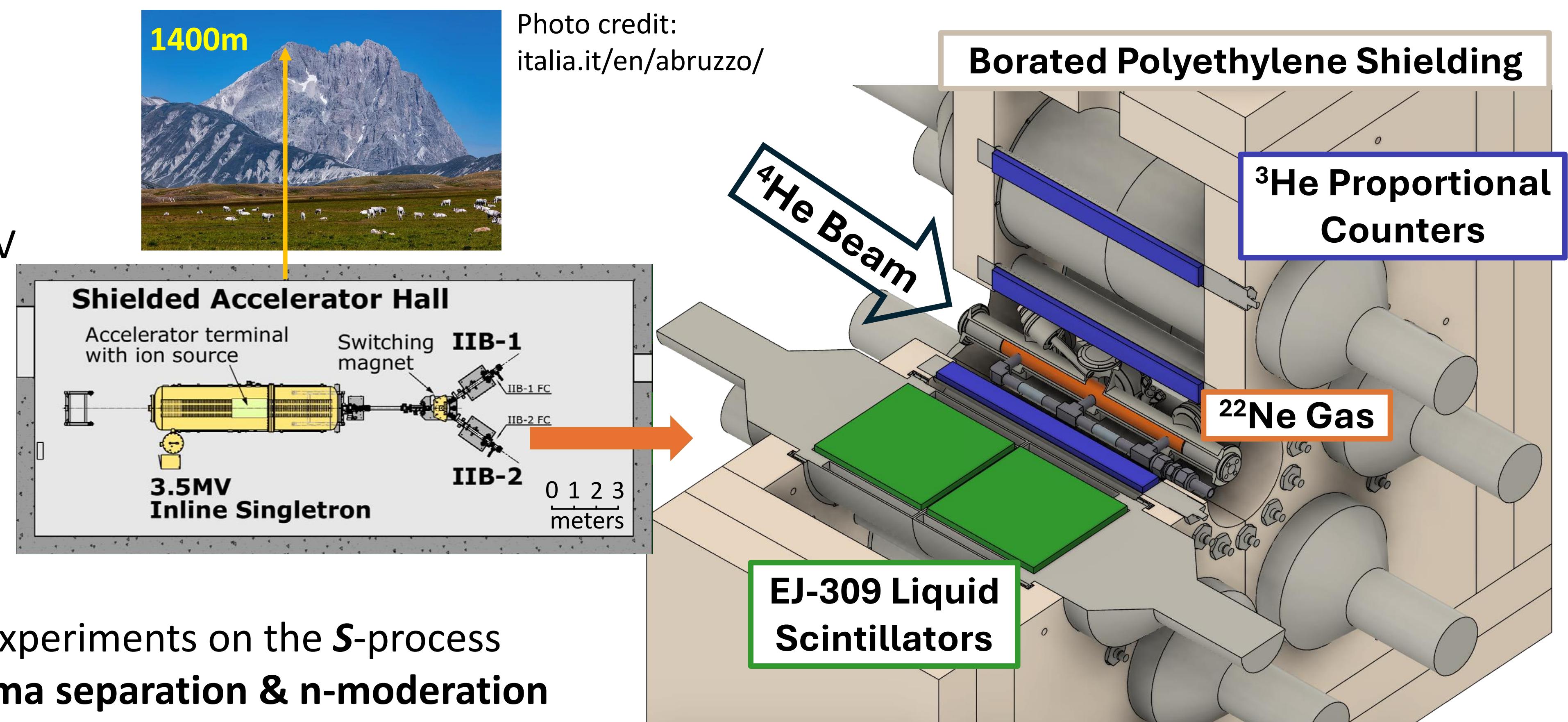
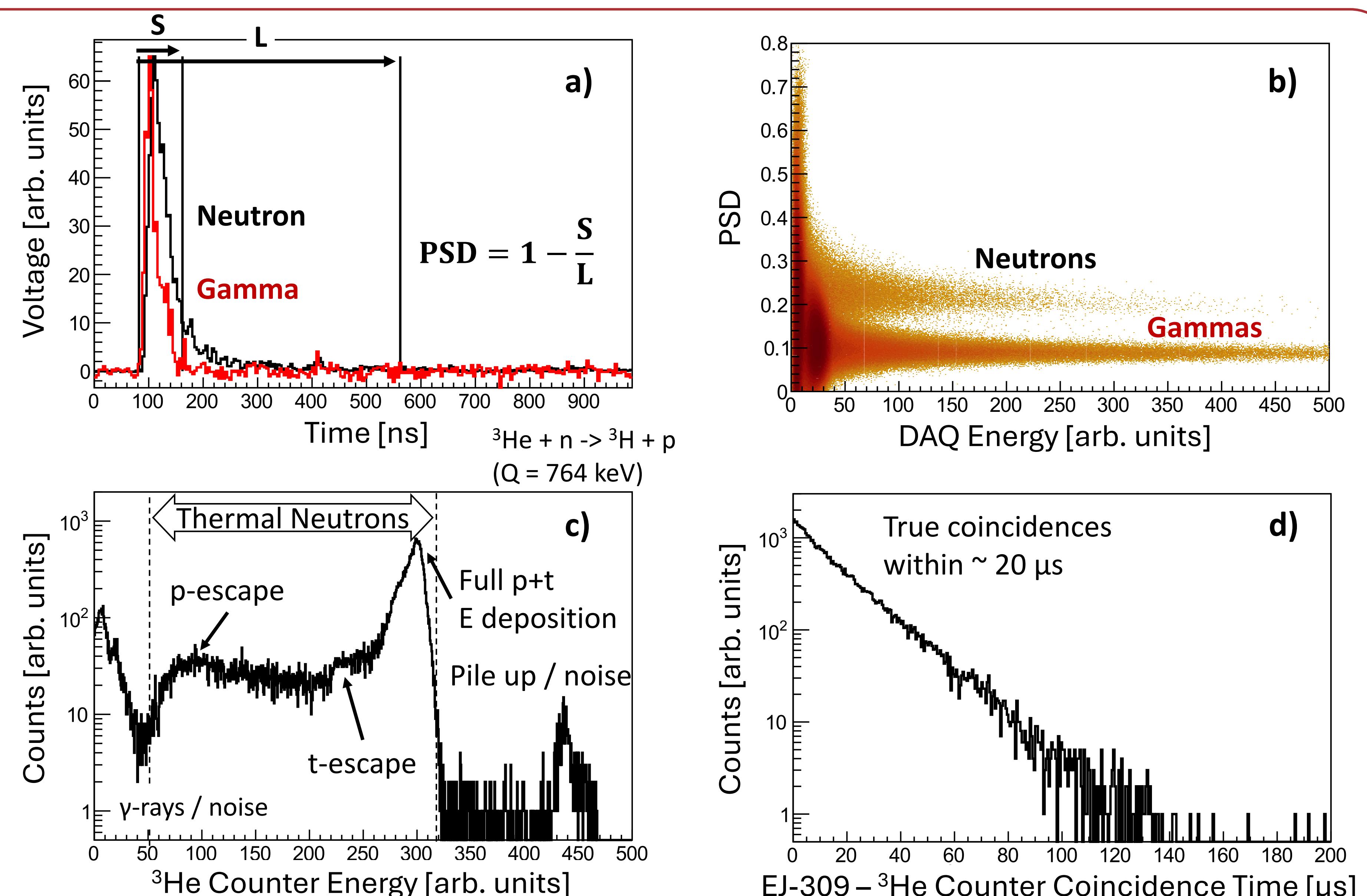


Fig 2. Computer-assisted drawing of SHADES detector array

4. Preliminary Data Analysis

- **Calibration data:** AmBe neutron source
 - a) **Typical waveforms** after baseline subtraction
 - b) **Scintillator PSD vs energy**
 - c) **Energy spectrum** from ^{3}He counters
 - d) **Coincidence timing** between Scintillators and ^{3}He counters
- March 2024: **First beam test** using ^{nat}Ne gas
- Encountered beam-induced background (BIB) from $^{13}\text{C}(\alpha, n)^{16}\text{O}$ and $^{17}\text{O}(\alpha, n)^{20}\text{Ne}$ (**preliminary**)
- Improvements made include **additional PE shielding** and covering beamline elements with Tantalum, Copper, and Gold



5. Future Work

- Autumn 2024: Additional beam tests to ensure low BIB
- Winter 2024: $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$ reaction cross-section measurements with enriched ^{22}Ne gas
- Paper in progress: EJ-309 – ^{3}He counter prototype measurements of $^7\text{Li}(p, n)^7\text{Be}$ at Frankfurt.

6. References

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