

# Constraining the $\gamma$ process

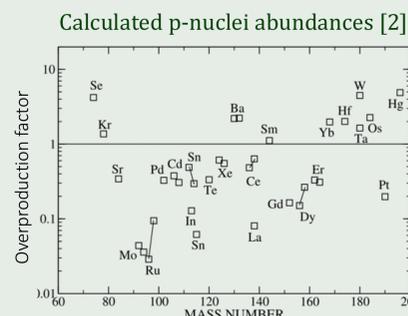
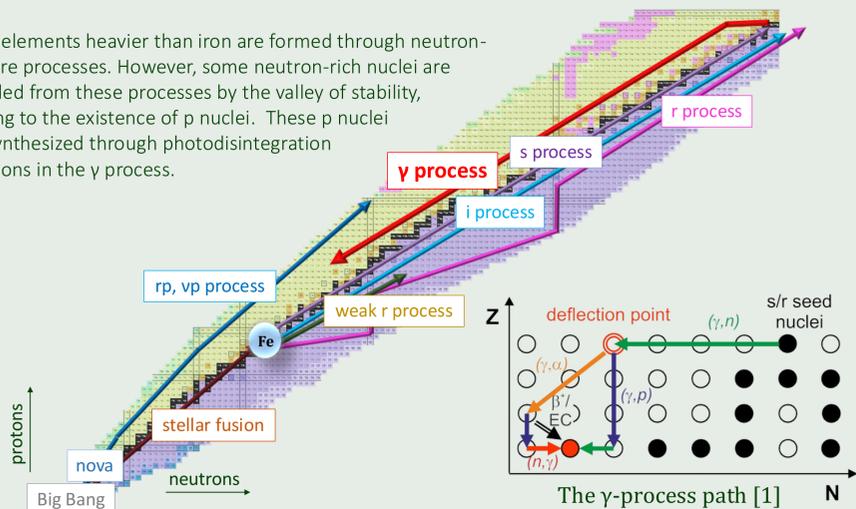
## Cross section measurements of $(p,\gamma)$ reactions in inverse kinematics

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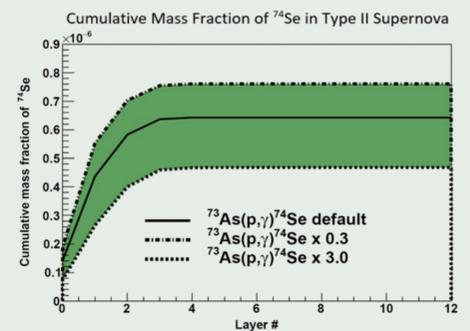
### INTRODUCTION

Most elements heavier than iron are formed through neutron-capture processes. However, some neutron-rich nuclei are shielded from these processes by the valley of stability, leading to the existence of p nuclei. These p nuclei are synthesized through photodisintegration reactions in the  $\gamma$  process.

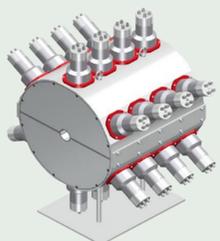


The path of the  $\gamma$  process begins by a series of  $(\gamma,n)$  reactions, followed by a  $(\gamma,p)$  or  $(\gamma,\alpha)$  reaction. This network of reactions involves more than 2000 nuclei, most of which are radioactive. Therefore the study of those  $(\gamma,p)$  and  $(\gamma,\alpha)$  reactions on radioactive elements is important for the accurate reproduction of the p nuclei.

The  $^{74}\text{Se}(\gamma,p)^{73}\text{As}$  reaction has been identified as one with high impact on the final abundance of the lightest p-nucleus  $^{74}\text{Se}$ , as it is one of its main destruction mechanisms. In this work we focus on the measurement of the inverse reaction  $^{73}\text{As}(p,\gamma)^{74}\text{Se}$ , with a radioactive  $^{73}\text{As}$  beam.



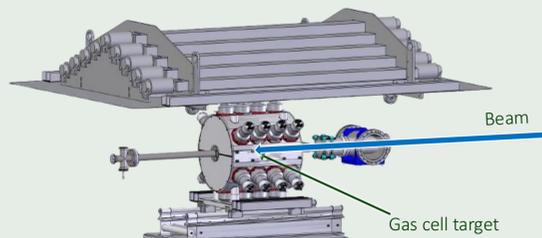
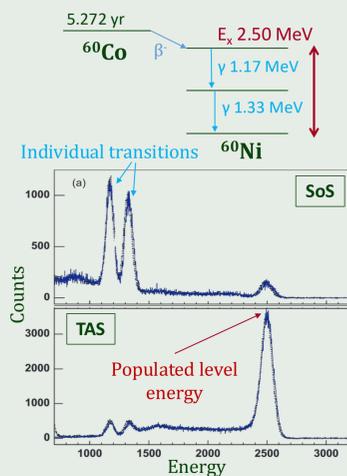
### METHODS



#### Summing NaI(Tl) SuN

- Large size, high efficiency  $\gamma$ -ray calorimeter
- 8 optically isolated segments, 24 PMTs
- Sum of Segments (SoS)  $\rightarrow$  Information about individual  $\gamma$ -rays
- Total Absorption Spectrum (TAS)  $\rightarrow$  Information about total excitation energy Ex

#### Total Absorption Spectroscopy

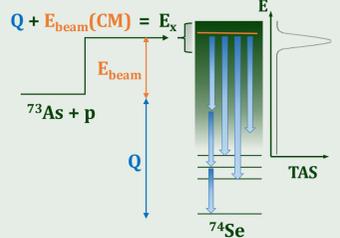


#### Experiments

- Cross section measurements in inverse kinematics at the Facility for Rare Isotope Beams
  - Proof-of-principle experiment with a stable beam for the  $^{82}\text{Kr}(p,\gamma)^{83}\text{Rb}$  reaction (2017)
  - Radioactive beam experiment for the  $^{73}\text{As}(p,\gamma)^{74}\text{Se}$  reaction (2023)
- Hydrogen gas cell target located in the center of SuN
- SuN + SuNSCREEN [3] detectors for  $\gamma$ -ray detection and cosmic background rejection

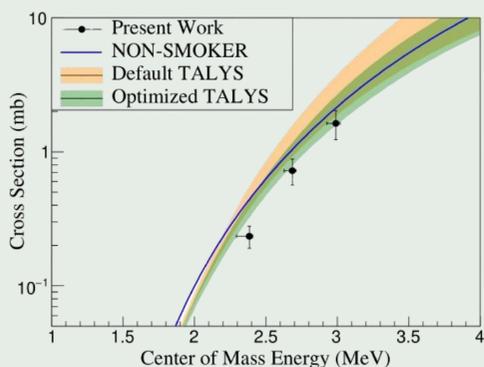
#### Analysis Overview

- After background subtraction, the measured yield corresponds to the integral of the total absorption peak (sum peak)
- Due to energy straggling within the gas target and Doppler shift, the sum peak was widened significantly  $\rightarrow$  The detection efficiency needs to be calculated as a function of all contributing energies within the sum peak
- The de-excitation of the compound nucleus was simulated using Hauser-Feshbach theory in order to calculate the function of energies that contribute in the sum peak  $\rightarrow$  Obtained constraints on the statistical properties of the compound nucleus

$$\sigma = \frac{\text{Yield}}{\Phi \cdot N_t \cdot \epsilon_{eff}(E)}$$


### RESULTS

#### $^{82}\text{Kr}(p,\gamma)^{83}\text{Rb}$

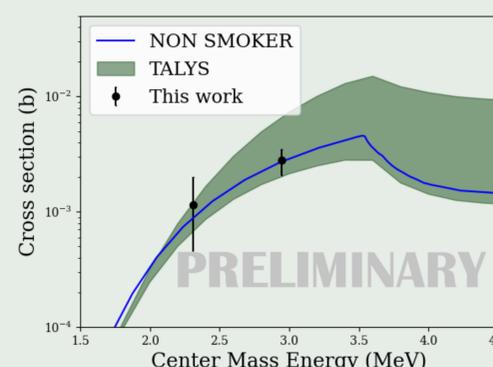


The  $^{82}\text{Kr}(p,\gamma)^{83}\text{Rb}$  cross section was measured in three energies within the Gamow window for the  $\gamma$  process [4].

The results indicate that standard statistical model calculations using NON-SMOKER and TALYS tend to overestimate the cross section. Based on experimental data in neighboring nuclei [5,6], there appears to be a consistent trend in this mass region.

The constraints on the statistical properties of the  $^{83}\text{Rb}$  nucleus allow for a better description of the experimental data with TALYS, as well as a constrain on the cross section on broader energy range

#### $^{73}\text{As}(p,\gamma)^{74}\text{Se}$



The  $^{73}\text{As}(p,\gamma)^{74}\text{Se}$  cross section was measured for the first time within the Gamow window for the  $\gamma$  process.

There is good agreement between the measured cross section and statistical model calculations using NON-SMOKER. This may suggest that the overproduction of Se in network calculations is not due to uncertainty in this reaction.

#### Future Work

1. Finalize analysis of the  $^{73}\text{As}(p,\gamma)^{74}\text{Se}$  data
2. Provide broader cross section constraint from statistical properties
3. Study the effect of the extracted  $^{73}\text{As}(p,\gamma)^{74}\text{Se}$  cross section on the  $^{74}\text{Se}$  final abundance for a SNIa scenario

### REFERENCES

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