### New measurement of the ${}^{2}H(p,\gamma){}^{3}He$ reaction at Felsenkeller underground lab Alexander von HUMBOLDT STIFTUNG

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# Abstract

The primordial deuterium abundance can be used to probe the cosmological baryon density. The  $^{2}H(p,\gamma)^{3}He$  reaction is responsible for deuterium destruction during Big Bang Nucleosynthesis (BBN) and plays a crucial role in determining its abundance. This reaction was previously measured at BBN energies by the LUNA collaboration using a windowless gas target system, covering a proton beam energy range of 50–400 keV. However, measurements at higher energies using solid targets indicate an extrapolated cross-section that is 10 % higher. To reduce the tension between low- and high-energy data, a new measurement of the  $^{2}H(p,\gamma)^{3}He$  reaction was performed at the Felsenkeller shallow-underground laboratory in Dresden, covering a proton beam energy range of 300–1200 keV, with partial overlap with both data sets. The reaction was studied using a solid target setup and multiple HPGe detectors positioned around the target, providing both cross-section and angular distribution data. The latter provides additional constraints for improving ab initio calculations.

## **Motivation**

# State of the art

# **Big Bang Nucleosynthesis (BBN)**





The promordial abundance depens only on:

- Baryon density:  $\Omega_b h^2$
- Particle Physics: *N*<sub>eff</sub>...,
- Nuclear astrophysics: Cross sections of relevant processes at BBN energies

For many years  ${}^{2}H(p,\gamma){}^{3}He$  reaction has governed the uncertainty on the primordial deuterium abundance.

- The BBN energy range measured by LUNA, <3% uncertainty
  - $\Omega_{\rm B} h^2 |_{\rm CMB} = (2236 \pm 15) \times 10^{-5}$
- $\Omega_{\rm B}h^2|_{\rm BBN, \ LUNA[10]} = (2233 \pm 36) \times 10^{-5}$
- High energy data at  $E_p = 400-800$  keV at HZDR

# **Felsenkeller GOAL**



HELMHOLTZ ZENTRUM

**DRESDEN** ROSSENDORF

• Angular distribution measurements at  $E_p = 300 - 1200 \text{ keV}$ 

• Cross-section data at  $E_p = 300 - 1200$  keV (overlap LUNA-HZDR data)

# **Experimental setup**

#### **5 MV Felsenkeller accelerator**



#### **ChETEC-INFRA** supported

<sup>●</sup> <sup>2</sup>H, <sup>4</sup>He RF beams <sup>12</sup>C, <sup>16</sup>O, and other beams Beam current up to 30 uA

15.5

143.5

# Data taking and analysis

- Detection efficiency using standard sourses,  ${}^{27}AI(p,\gamma){}^{28}Si$  reaction at 992 keV resonance and Monte Carlo simulations
- High statistics run at  $E_p = 300 1200$  keV with 50 keV step
- Reference run at  $E_p = 608$  keV before and after each energy



# Preliminary results and outlook



### **Target analysis Before irradiation** After irradiation







#### **ERDA** EBS **NRA** Energy (MeV) (p,p) 2MeV D(³He,p)⁴He

### **Angular distributions**

![](_page_0_Figure_43.jpeg)

![](_page_0_Figure_44.jpeg)

![](_page_0_Figure_45.jpeg)

# **Preliminary S-factor**

![](_page_0_Figure_47.jpeg)

![](_page_0_Picture_48.jpeg)