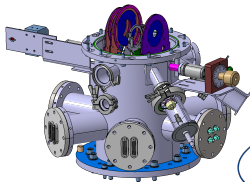
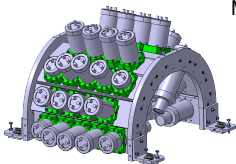


Solving the Carbon Fusion Riddle at Deep Sub-barrier Energy

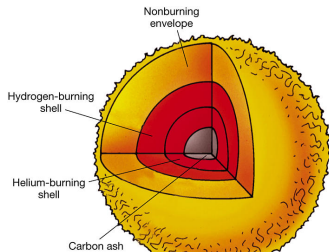
Marcel Heine for the  STELLA Collaboration

IPHC/CNRS Strasbourg

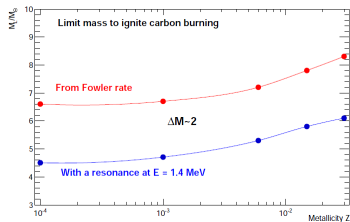
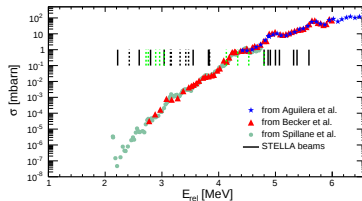
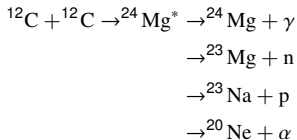
September 17, 2025



Carbon Fusion in Massive Stars



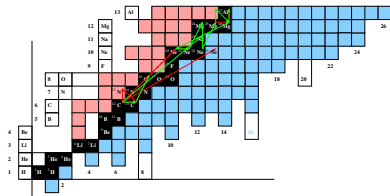
Copyright © 2005 Pearson Prentice Hall, Inc.



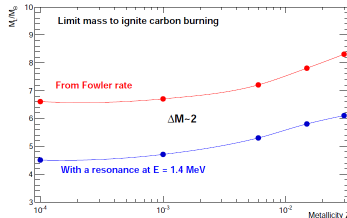
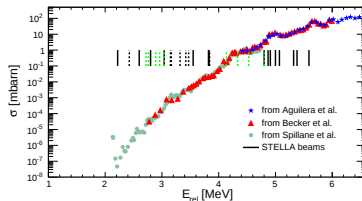
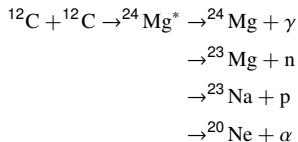
from: O. Straniero et al., JPC Ser. 665 (2016), 012008



Carbon Fusion in Massive Stars



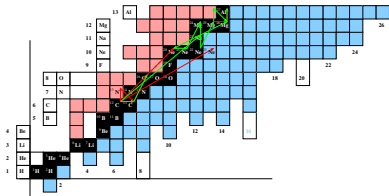
from A. Chieffi *et al.*, *APJ* 502 (1998), 737



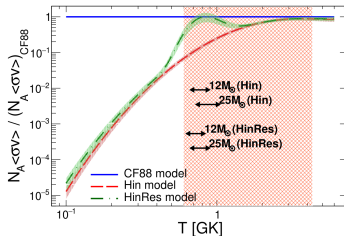
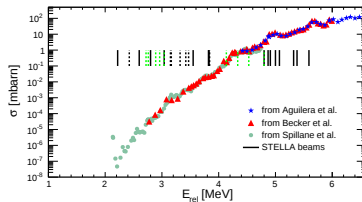
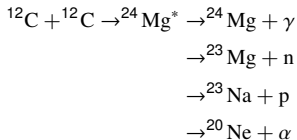
from: O. Straniero *et al.*, *JPC Ser.* 665 (2016), 012008



Carbon Fusion in Massive Stars



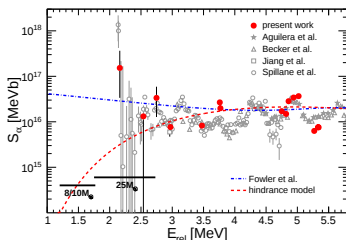
from A. Chieffi et al., APJ 502 (1998), 737



E. Monpriat et al., A&A 660, (2022) A47

Previous Experiment Related $^{12}\text{C}+^{12}\text{C}$ S-factors (Possibly not Complete)

- ▶ oscillating excitation function
- ▶ region systematically below CF88
- ▶ vanishing cross sections ($\leq \text{nbarn}$)



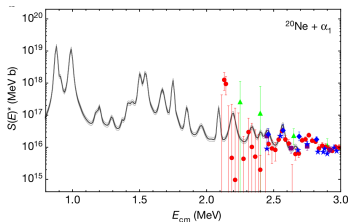
G. Fruet *et al.*, *PRL* **124**, (2020) 192701

- ▶ direct/indirect measurements
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- ▶ L. Barrón-Palos *et al.*, *NP A* **779**, 318, (2006)
- ▶ D. Jenkins *et al.*, *PRC* **76**, 044310, (2007)
- ▶ C.L. Jiang *et al.*, *PRC* **75**, 015803, (2007)
- ▶ T. Spillane *et al.*, *PRL* **98**, 122501, (2007)
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- ▶ A. Tumino *et al.*, *Nature Letter* (23.05.2018)
- ▶ J. Zickefoose *et al.*, *PRC* **97**, 065806 (2018)
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- ▶ C. Beck *et al.*, *EPJA* **56**, 87 (2020)
- ▶ G. Fruet *et al.*, *PRL* **124**, 192701 (2020)
- ▶ Y.J. Li *et al.*, *CPC* **44**, 115001 (2020)
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- ▶ N.T. Zhang *et al.*, *PLB* **801**, 135170 (2020)
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A. Tumino *et al.*, *Nature* **557**, (2018) 687

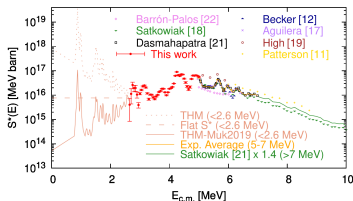
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- ▶ P.L. Marley *et al.*, **JMPE** 17, 2040, (2008)
- ▶ C.L. Jiang *et al.*, **NIM A** 682, 12, (2012)
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- ▶ M. Heine *et al.*, **NIM A** 903, 1 (2018)
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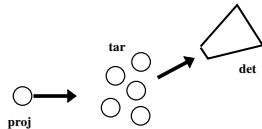


W.P. Tan *et al.*, *PRC* **110**, (2024) 035808

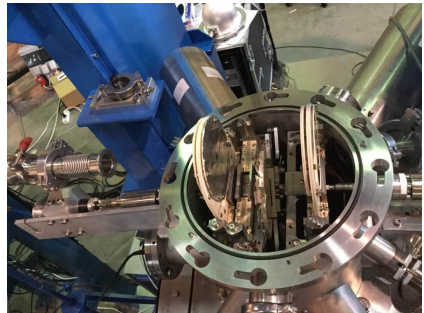
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- ▶ L. Morales-Gallegos *et al.*, *EPJA* **60**, 11 (2024)
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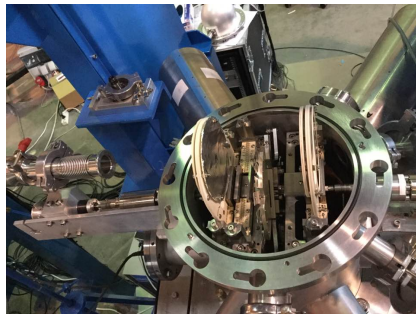
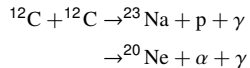
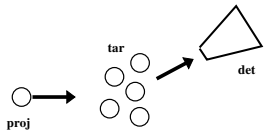
The STELLar Laboratory Mobile Station



1. beam intensities of a few pμA
2. data taking for weeks
3. γ -particle coincidences, granularity
4. low counting statistics analysis



The STELLar Laboratory Mobile Station

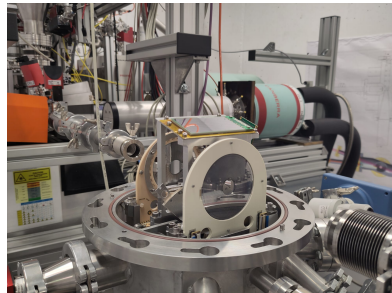


The STELLar Laboratory Mobile Station



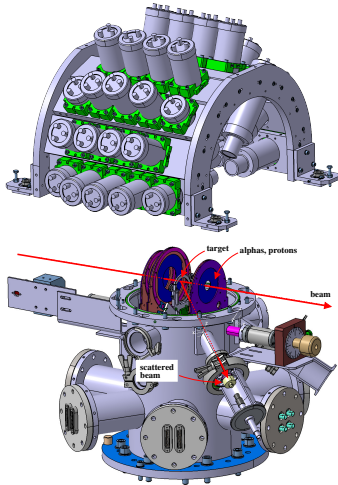
setup at the Felsenkeller laboratory
provided:

- ▶ Faraday cup
- ▶ pumps/venting
- ▶ DAQ



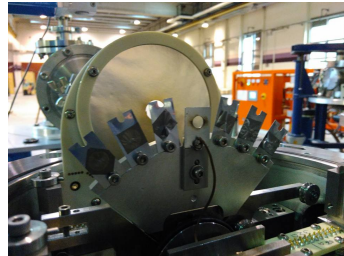


The STELLar Laboratory Mobile Station



monitoring of target state:

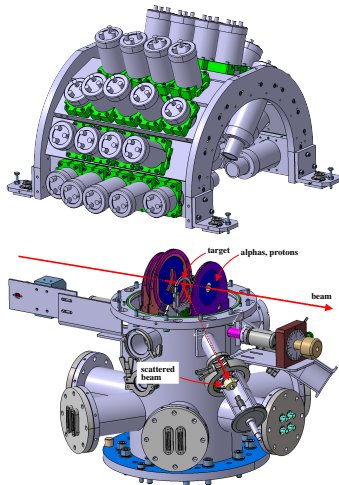
- ▶ elastic scattering
- ▶ beam integral
- ▶ alpha source runs



M. Heine *et al.*, NIM A **903**, (2018) 1–7
J. Nippert *et al.*, APPB **17**, (2024) 3-A33

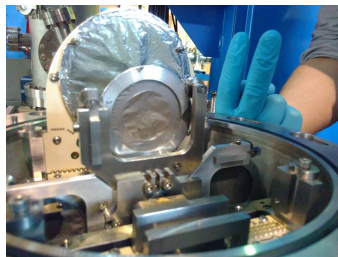


The STELLar Laboratory Mobile Station



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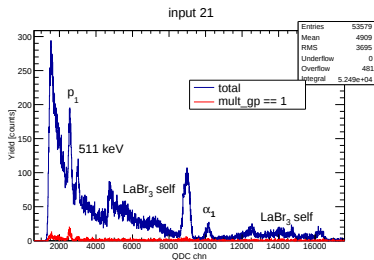
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M. Heine *et al.*, NIM A **903**, (2018) 1–7
J. Nippert *et al.*, APPB **17**, (2024) 3-A33



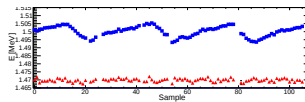
Analysis Technique



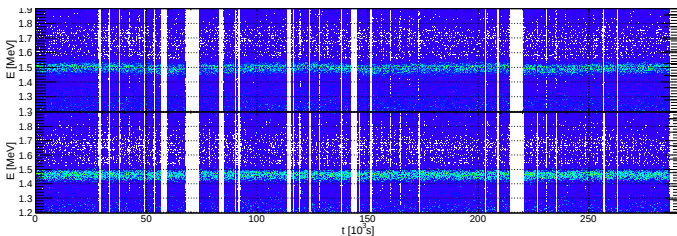
~ 3 days of data; 45 min blocks

▶ drift of 1.47 MeV line: 1.5 keV

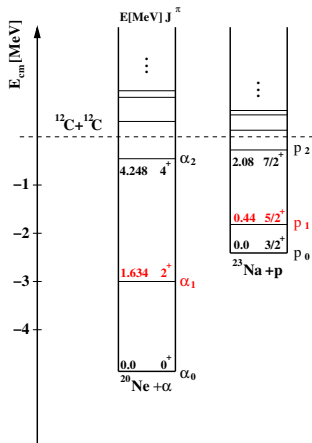
▶ since calibration: 30 keV



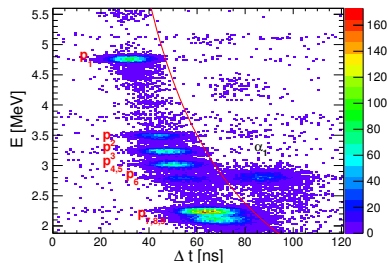
M. Heine *et al.*, NIM A **903** (2018), 1–7



- ▶ synchronization of 1 GHz gamma DAQ and 125 MHz particle DAQ
- ▶ energy-deposition in silicon substrate: triggering on pulse shape



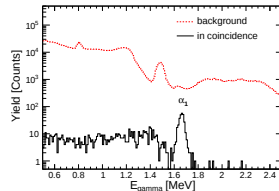
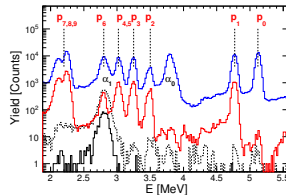
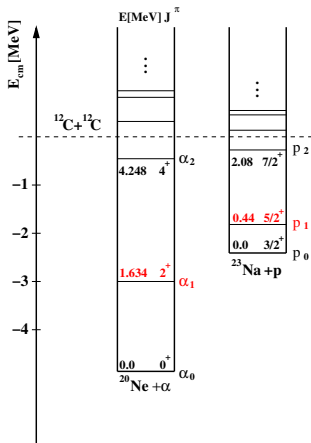
- ▶ timing gates $\sigma \sim 15$ ns
- ▶ proton- α separation



M. Heine *et al.*, NIM A 903 (2018), 1

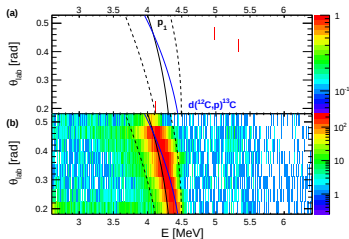
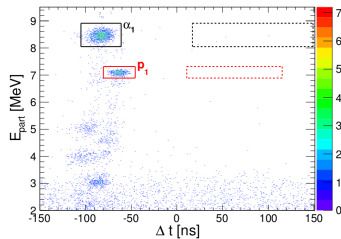


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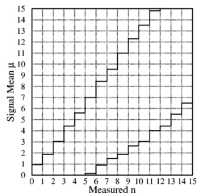


G. Fruet et al., PRL 124, (2020) 192701

Analysis Technique

G. Fruet *et al.*, PRL 124 (2020), 192701M. Heine *et al.*, EPJ Web Conf 260 (2022), 01004

► avoid negative physics counts:



G.F. Feldman & R.D. Cousins, PRD 57 (1998), 3873

$$n = 1 \text{ (observed)}, b = 1 \text{ (background)}, \mu = n - b$$

$$\frac{\Delta\mu}{\mu} = \sqrt{\left(\frac{\Delta n}{n}\right)^2 + \left(\frac{\Delta b}{b}\right)^2} \text{ N.A.N.}$$

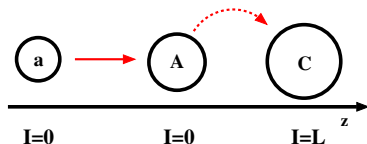
$$n = 1, b = 1 \rightarrow \mu \in [0.00, 1.75]$$

$$n = 2, b = 2 \rightarrow \mu \in [0.00, 2.25]$$



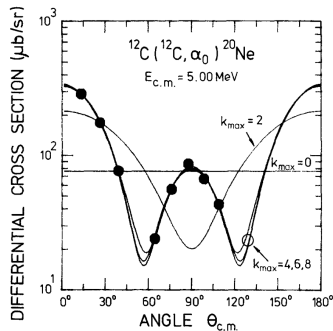
N. Bohr hypothesis of independence: compound formation and decay are independent
but angular momentum conservation

fusion of spinless particles:



- ▶ L perpendicular to z
- ▶ equal probability $\pm z$

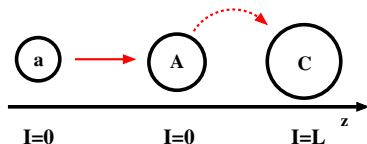
$$\left(\frac{d\sigma}{d\Omega}\right) = \sum_{k=0}^{k_{\max}} a_k P_k(\cos(\theta)), \quad k = 0, 2, 4, \dots$$



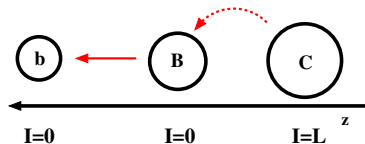


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- compound state spin from fusion measurements
- ▶ normalisation of cross sections:
 $P_0(\cos(\theta))$

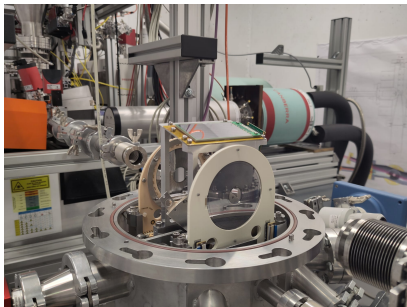
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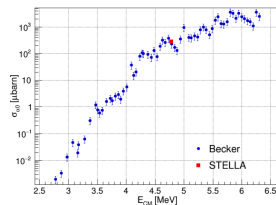
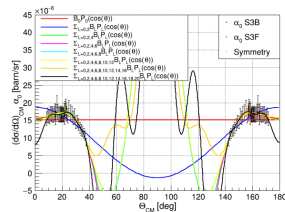
Angular-Differential Fusion Cross-Sections

E/p conversation: angular distribution α , p:

- ▶ accurate normalisation
- ▶ spin parity of compound nucleus ^{24}Mg
- ▶ effective beam energy



- ▶ χ^2 and uncertainty analysis
- ▶ spin-parity conservation

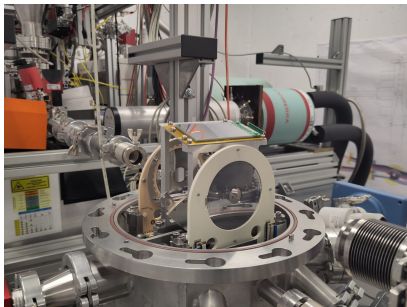


courtesy: G. Harmant

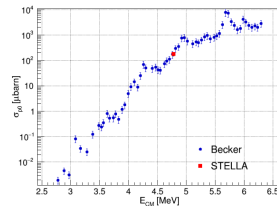
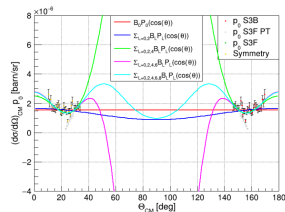
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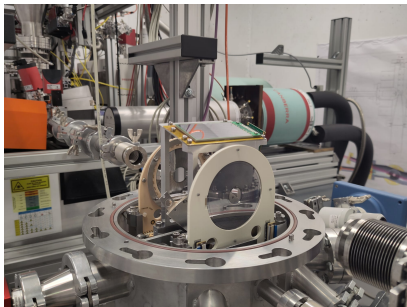
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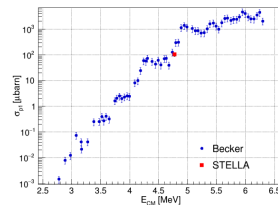
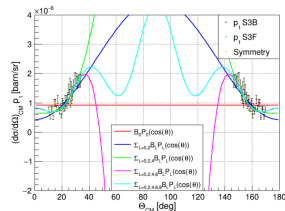
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courtesy: G. Harmant

Thank You Very Much for Your Attention

- ▶ carbon burning in massive stars
 - ▶ cross sections below nano barns
 - ▶ oscillating excitation function
- ▶ direct measurements of carbon fusion with STELLA
 - ▶ high duty experimentation
 - ▶ long term stability
 - ▶ low count statistics
- ▶ resonance determination with STELLA
 - ▶ scan of known resonances: 0^+ , 2^+ , 4^+
 - ▶ alpha, proton detection at steep angles