

Universe Through Archival Data

Understanding and Creating Astronomical Images

Pavlo Plotko

2nd ChETEC-INFRA Observational School (ChINOS) 2025

21.07.2025

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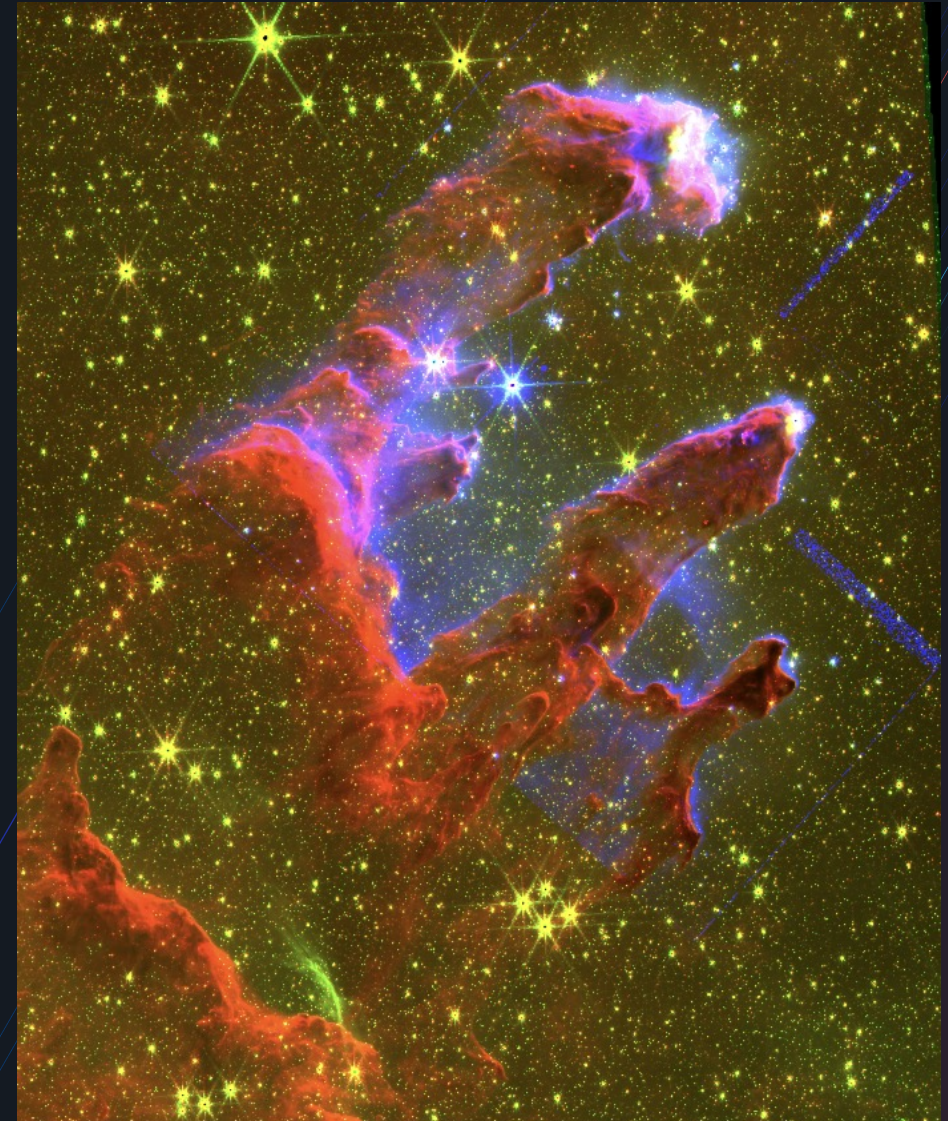
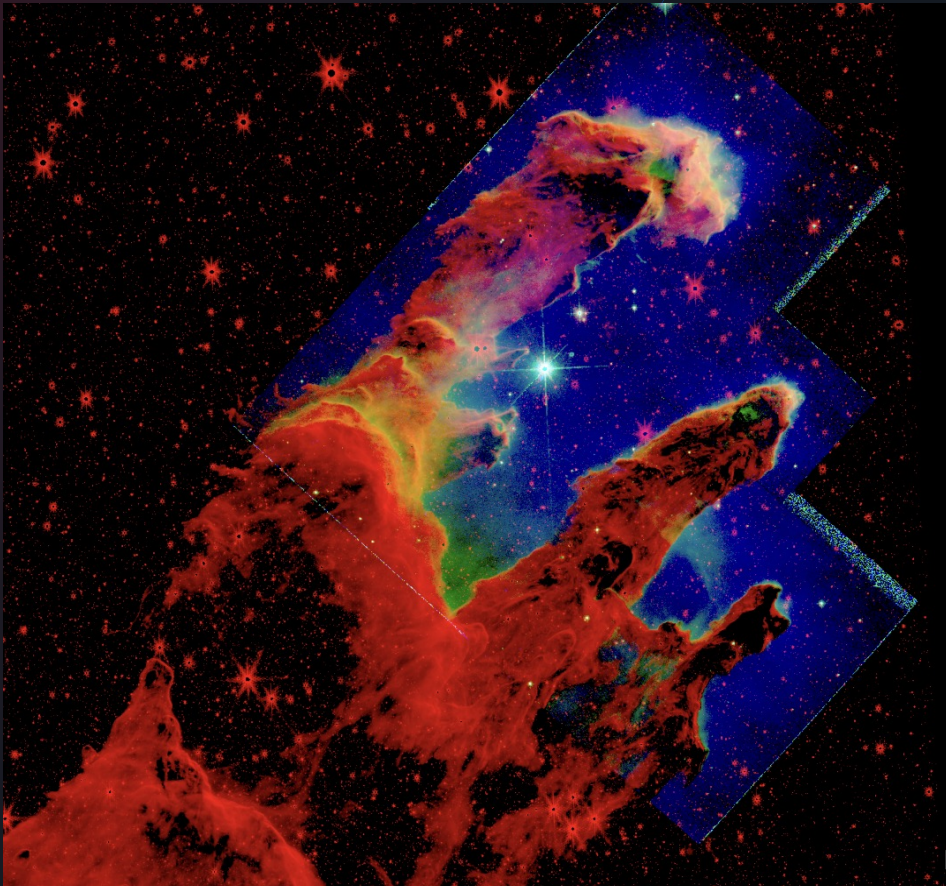
plotkopavlo.com



The goal: create an image of M16

Learn about archives and images

/2



M16 – Pillars of Creation



Who am I?

Building and coordinating of all activities with Eastern Europe/Baltic at DZA research ("part-time")



Organizing:
Astronomy Tournaments
Astrophysics Olympiads
Astrophysics Conferences

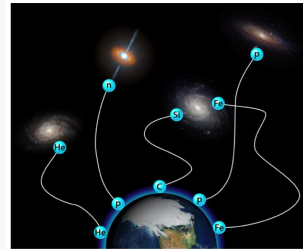


Summer schools
Hackathon



Coordinator of the recovery of Ukrainian Astronomy

PhD: theoretical astrophysics; UHECRs; (DESY Germany)



Master's and Bachelor's: Astronomy;
Thesis: Acceleration and non-thermal emission of cosmic rays in SNR (Kyiv, Ukraine)



Before that: Photometric studies of asteroids, stars, and GRB afterglow (Kharkiv, Ukraine)



Web developer and Software developer in the past

Addressing the Elephant in the Room: AI and Chatbots

Asking the right questions and understanding the ideas is more important than ever — that's what I aim to do here.

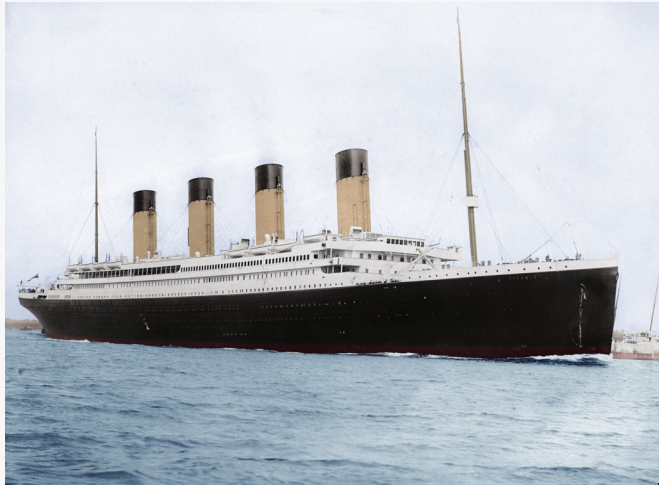
Since you can simply ask ChatGPT (please not the new version of Grok) for step-by-step DS9 instructions — and it works.”



<https://zapier.com/blog/best-ai-chatbot/>

Images and Archives Are Like Icebergs

The beauty is just the tip; the complexity is hidden below.



https://commons.wikimedia.org/wiki/File:Titanic_in_color.png

Archive data

How
long
to
store

Who
pays



What
type
of
data

<https://www.pinterest.com/pin/tip-of-the-iceberg-poster--189151253084507208/>

Image

Png
Jpg
Fitts
fiff

CIE 1931
color
space



Vector
Raster

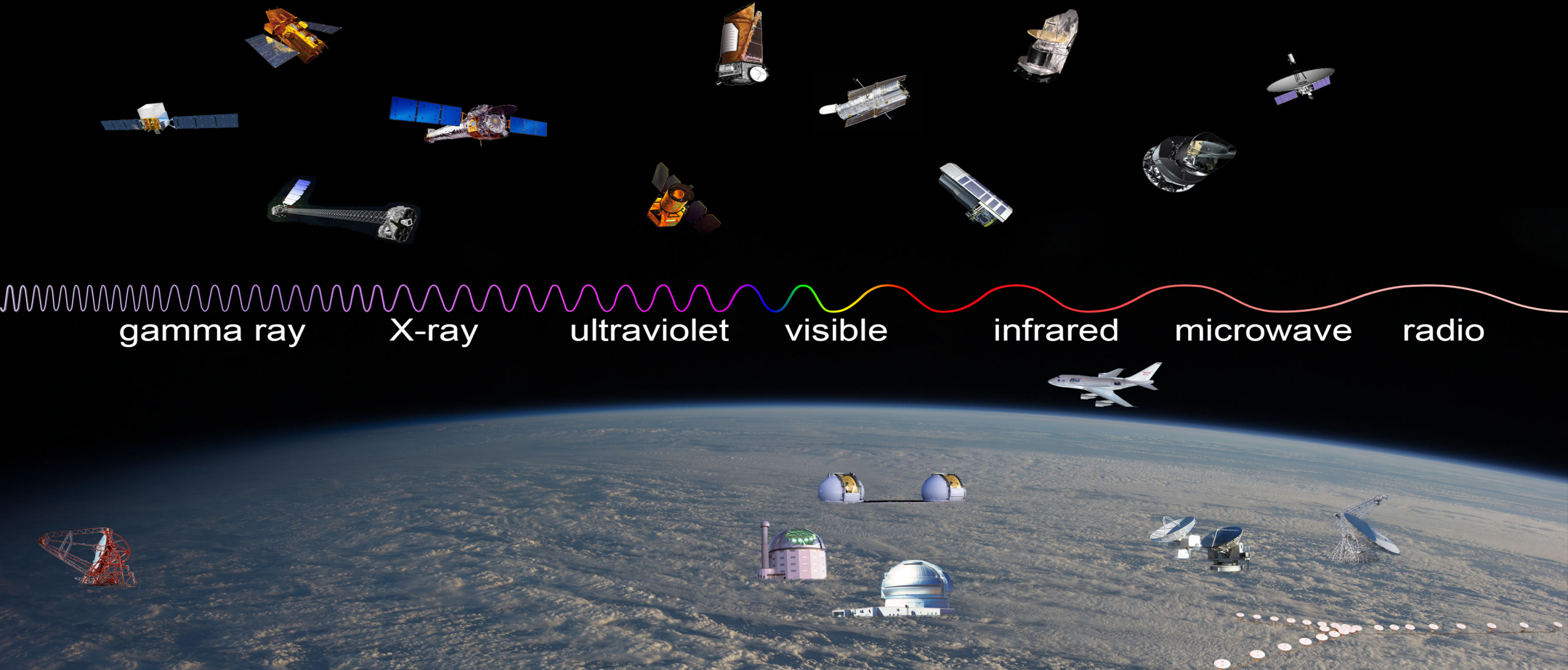
GRB
CMYK

Gamut

OLED
LCD
LED

Huge volumes of astronomical data from multiple missions

https://imagine.gsfc.nasa.gov/science/toolbox/emspectrum_observatories1.html



Each "Detector" has its own archive

The multi-messenger observatories are distributed in many places around the world, and in satellites around it. They necessarily demand a global collaboration, early distribution of any detected transient signal, full accessibility and transparency of the scientific data, and of the methods used to analyse them.

THE MAP OF DETECTORS



LIGO
KAGRA
VIRGO
LIGO INDIA
EINSTEIN TELESCOPE/COSMIC EXPLORER



GRAVITATIONAL WAVES

ALMA
ASKAP
ATCA
LOFAR
MEERKAT
VLA/NGVLA
SKA
E-ELT
GEMINI
GMT
KECK
LSST
SUBARU
TMT
VLT
ZTF
CTA
HAWC
HESS
MAGIC



ELECTROMAGNETIC

ANITA
GVD
ICECUBE
KM3NET
ANTARES
BOREXINO
DUNE
JUNO
SNO
SUPER K/HYPER K



NEUTRINOS

AUGER
LHAASO
TA



COSMIC RAYS

ANDES
CUPP
DUSEL
KAMIOKA
LNGS
LSC
LSM
SNOLAB



UNDERGROUND LABORATORIES

EPTA
NANOGRV
PPTA
PTA



RADIO PULSAR TIMING ARRAYS

No One-Size-Fits-All

Archives are shaped by:

- Mission goals
- Instruments
- Wavelength coverage

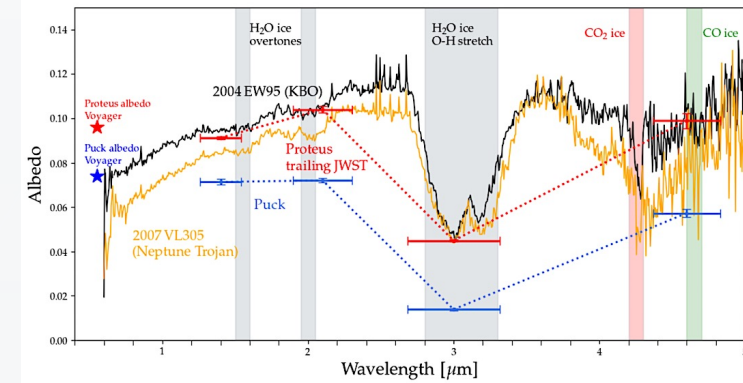
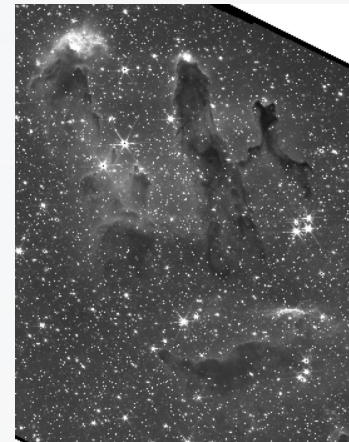
Examples:

Gaia: astrometry & photometry

JWST: IR imaging/spectroscopy

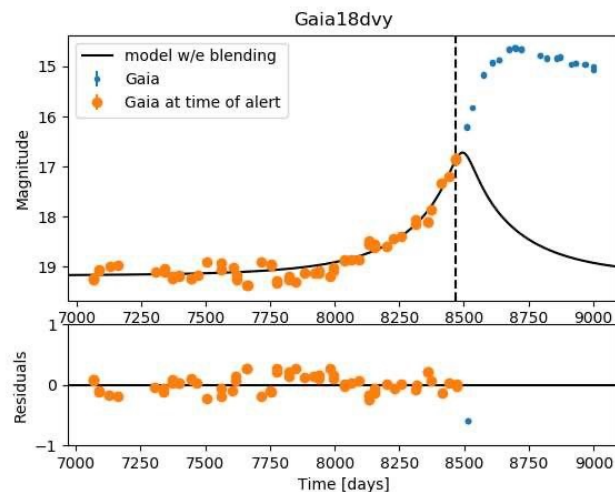
Planck: microwave background

XMM-Newton: X-ray spectra

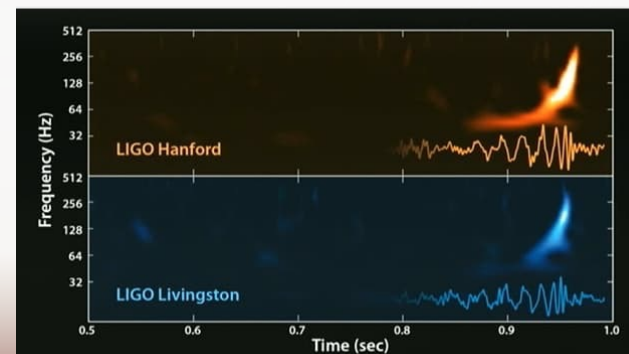


<https://sci.esa.int/web/planck/-/51555-planck-power-spectrum-of-temperature-fluctuations-in-the-cosmic-microwave-background>

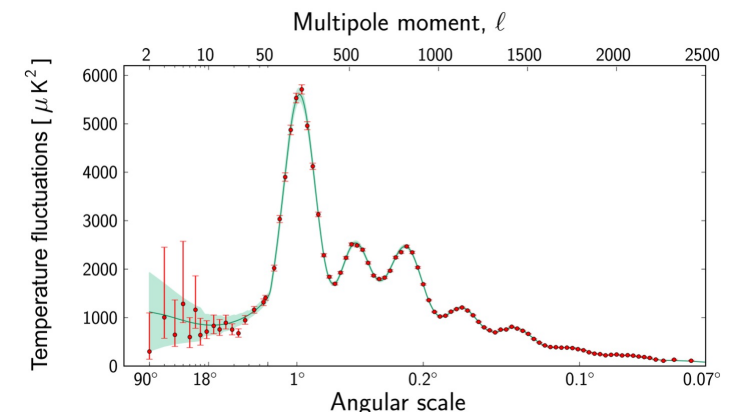
Archives are unique, science-driven, and often incompatible without translation tools

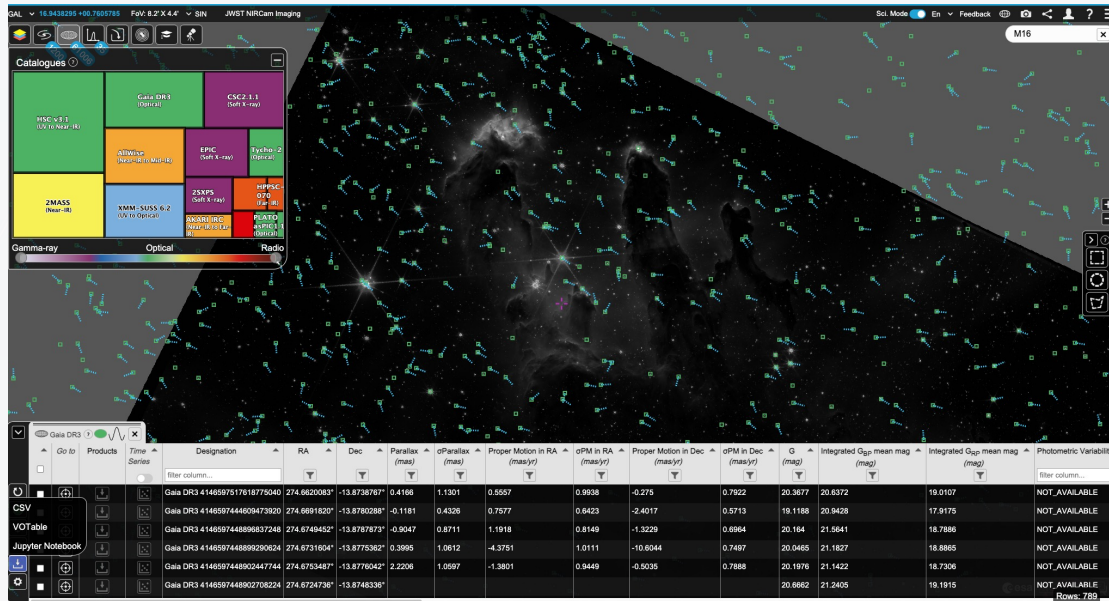


[10.48550/arXiv.2201.12209](https://arxiv.org/abs/10.48550/arXiv.2201.12209)



[LIGO/Phys. Rev. Lett. 116 061102](https://arxiv.org/abs/10.1103/PhysRevLett.116.061102)





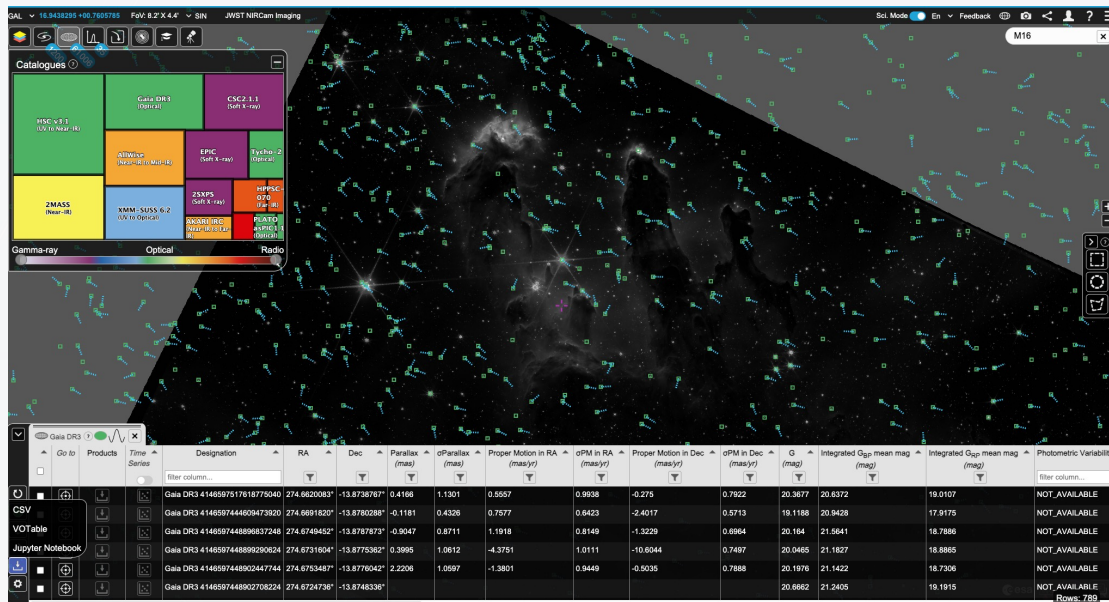
Access Modes

Web interfaces (ESA Sky, MAST, IRSA, etc.)

APIs and programmatic access (e.g., astroquery)

Table downloads, cutout services, cone searches

Different missions & archives may support different modes
— choose what fits your workflow!



PositionFile

Name

Equatorial

Target in

Circle

Box

Name

Radius

5

arc sec

Search in:

gaiadr3.gaia_source

Extra conditions

Display columns

Reset Form

Show Query

Submit Query

Output is limited to 2,000 sources

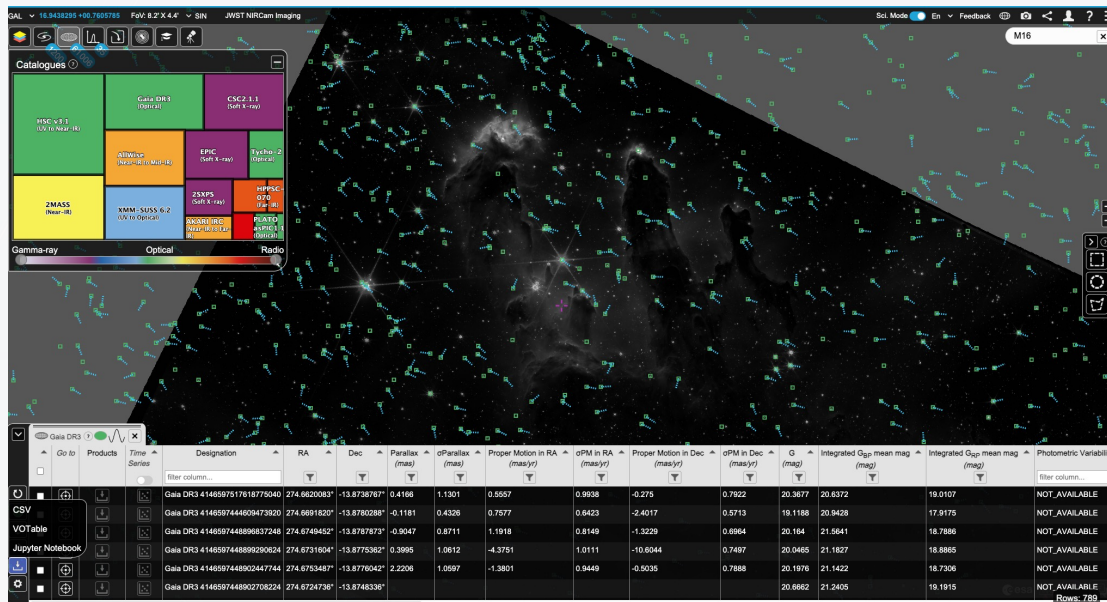
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```
In [7]: queryGJ473 = """SELECT source_id, phot_g_mean_mag, parallax, pmra, pmdec, ruwe
FROM gaiadr3.gaia_source
WHERE
CONTAINS(
POINT('ICRS',ra,dec),
CIRCLE(
'ICRS',
COORD1(EPOCH_PROPOS(188.3223395,9.0210534,227.9000,-1730.0000,203.0000,1.0000,2000,2016.0)),
COORD2(EPOCH_PROPOS(188.3223395,9.0210534,227.9000,-1730.0000,203.0000,1.0000,2000,2016.0)),
5./3600.)
)=1"""

In [8]: job = Gaia.launch_job(queryGJ473)
r = job.get_results()
r
```

No One-Size-Fits-All

Archives are shaped by: Examples:

- Mission goals
 - Instruments
 - Wavelength coverage
- Gaia:** astrometry & photometry
JWST: IR imaging/spectroscopy
Planck: microwave background
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Open vs Restricted Archives

Open archives: Freely accessible data (e.g., Hubble, Gaia, JWST)

Restricted (proprietary): Data under embargo for 6–12 months (or years...)

Key point: Always check data rights before using or publishing.

Developing and Maintaining

Developing and Maintaining

- Who should build and maintain archives?
**mission team; Dedicated institutions;
Community-driven or VO (Virtual
Observatory)**
- How long should data stay online?

Who Should Build and Maintain Archives?



Mission team



Knows instruments & data



Limited by mission budget & timeline

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- ✓ Long-term expertise & infrastructure
- ✗ Needs sustained funding

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Community/Virtual Observatory

- ✓ Shared standards & effort
- ✗ Harder to coordinate

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Why keep data?

- Legacy science, reproducibility, cross-mission studies
- New discoveries from old data

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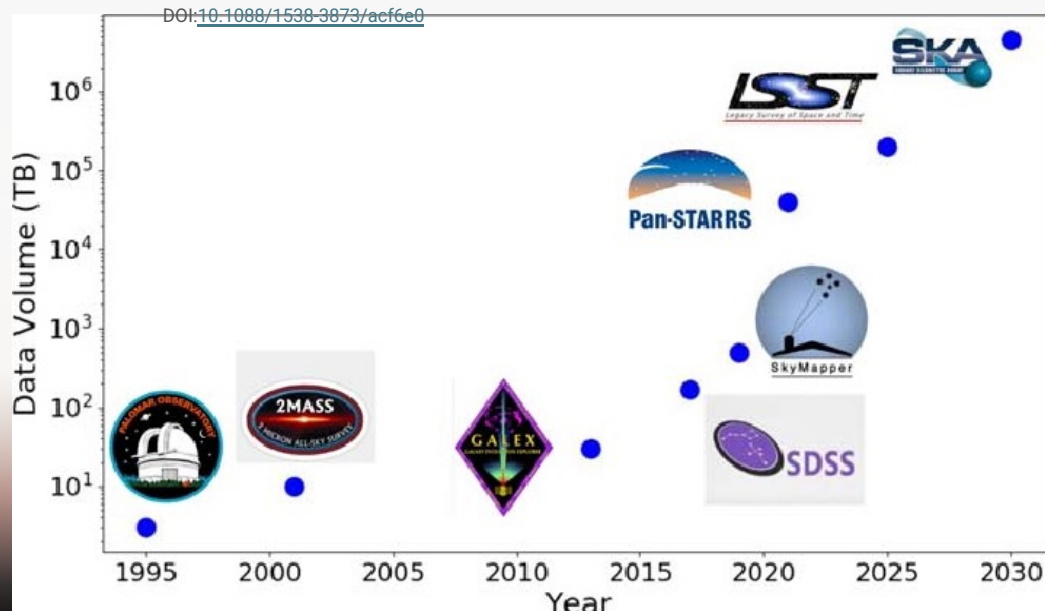
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Why not?

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DOI: [10.1088/1538-3873/acf6e0](https://doi.org/10.1088/1538-3873/acf6e0)



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If we had unlimited money...

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Could we store all the data?

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Yes.

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But should we?

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Making everything more complex: Carbon footprint (ALMA example)

Storage of 1TB of data -> 2000kg/yr of CO₂

Transfer of 1 GB of data -> 3kg



@Wiki

https://www.mur.gov.it/sites/default/files/2024-03/Astronomical_data_SIGRE.pdf

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A 100 GB dataset in ALMA Archive in 5 yr generates 13000kg of CO₂
corresponding to CO₂ generated by 3 cars driven continuously per 1yr
CO₂ absorbed by 215 trees in 10 yr



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the ALMA archive we have more than 60000 datasets (10GB-1TB size) =
= CO₂ absorbed by trees covering the whole area of Florence for 10yr



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Centre de Données astronomiques de Strasbourg (CDS)



Founded in 1972, located in Strasbourg, France

A multidisciplinary team of astronomers, software engineers, and data librarians

Curates and distributes astronomical data from publications, catalogs, and reference images

Provides three major services:

SIMBAD – object database with cross-identifiers and bibliography (~7.5 million objects)

VizieR – access to catalogs, observation logs, tables, images, spectra, time series

Aladin – interactive sky atlas and Virtual Observatory (VO) portal

Xmatch – a new tool allowing astronomers to efficiently cross-identify



ESA Sky Interface



Unified data access portal for ESA missions (HST, JWST, Gaia, XMM-Newton, Planck, etc.)

Allows sky exploration, image previews, and FITS downloads

Combines professional-grade tools with an accessible, visual interface

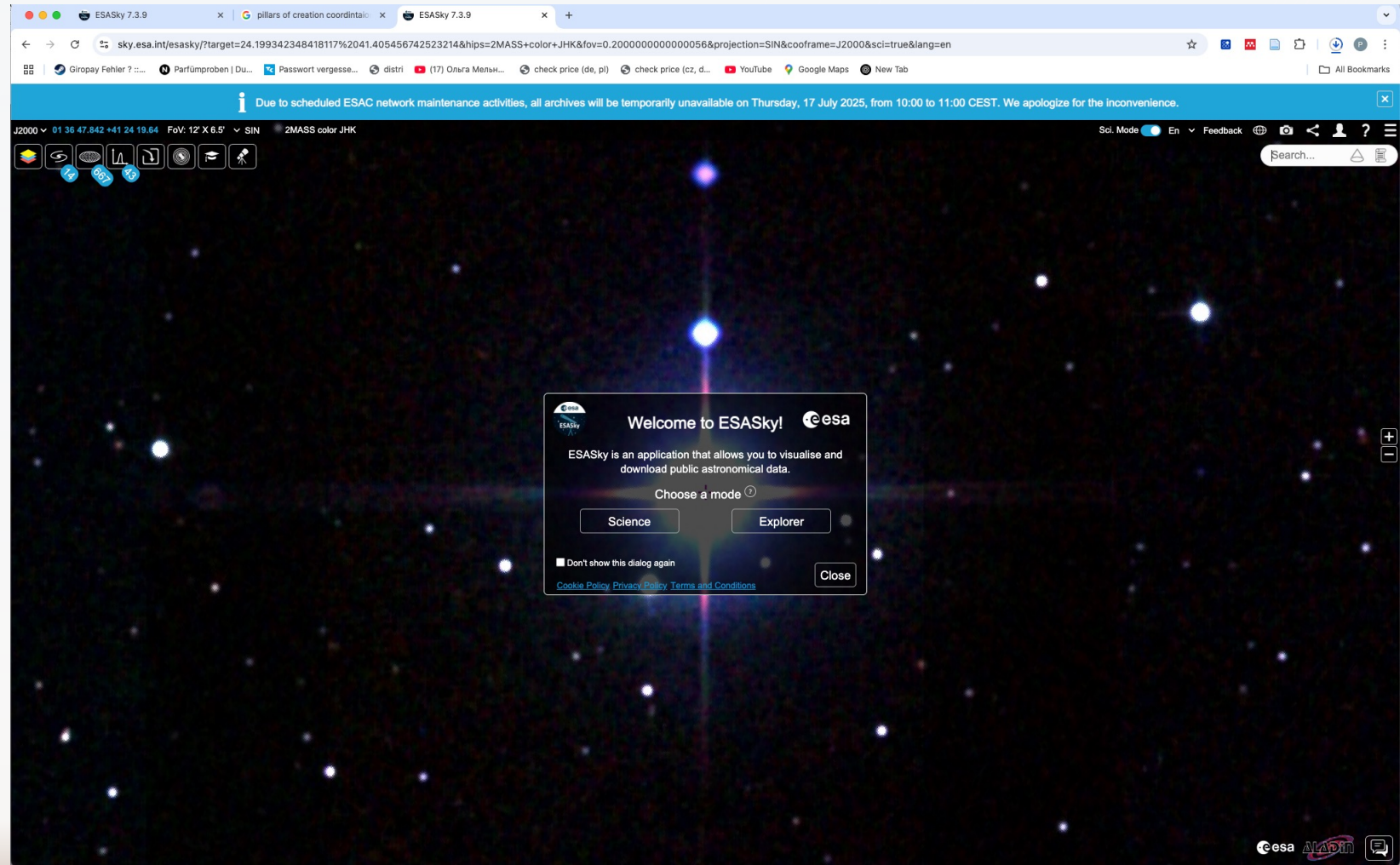
Focusing on ESA SKY

Why ESA? Unified interface, high usability, access to major missions
Overview of <https://sky.esa.int>
Supports Hubble, JWST, Gaia, XMM-Newton, Planck, and more

URL: <https://sky.esa.int>

Main features:

- Sky navigation
- Object search
- Filter selection
- FITS download
- Multi-mission layering



Websites with Professional FITS Image Archives

Professional astronomers use these websites to store and distribute their data, including FITS images. Many of these websites expect people to search for images of specific objects; a list of example objects is presented after this list of websites.

ESASky <http://sky.esa.int/>

GALEX View (requires Adobe flash)
<https://galex.stsci.edu/GalexView/>

Herschel Database in Marseille <http://hedam.lam.fr/>

Herschel User Provided Data Products
<http://www.cosmos.esa.int/web/herschel/user-provided-data-products>

Mikulski Archive for Space Telescopes (MAST) Portal
<https://mast.stsci.edu/portal/Mashup/Clients/Mast/Portal.html>

Sloan Digitized Sky Survey <https://dr12.sdss.org/fields>

NASA/IPAC Infrared Science Archive
<http://irsa.ipac.caltech.edu/images.html>

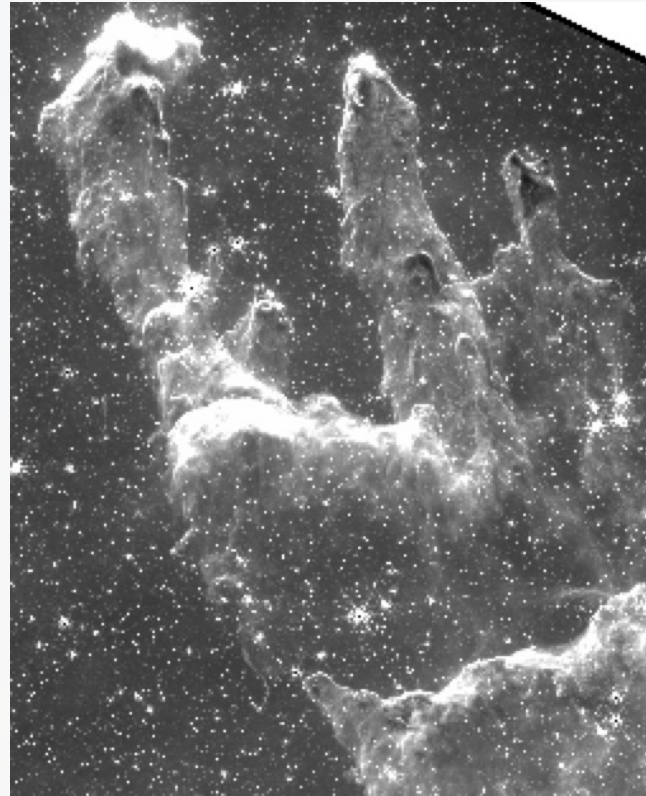
NASA/IPAC Extragalactic Database
<http://ned.ipac.caltech.edu/>

PanSTARRS-1 Image Access
<https://ps1images.stsci.edu/cgi-bin/ps1cutouts>

Spitzer 24-160 Micron Data for the Herschel-SPIRE Local Galaxies Guaranteed Time Programs
http://www.jb.man.ac.uk/~gbendo/exchange/SpitzerData/spitzerdata_main.html

UKIRT InfraRed Deep Sky Surveys
http://wsa.roe.ac.uk:8080/wsa/getImage_form.jsp

Use JWST and Hubble data of the Pillars of Creation.



How do we go from grayscale to color?

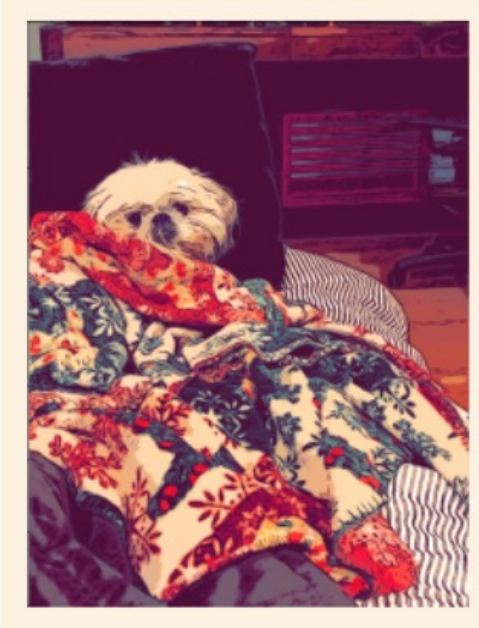


What is an image?*



* Raster

What is an image?



https://www.brandonrohrer.com/images_to_numbers

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What is an image?



https://www.brandonrohrer.com/images_to_numbers

Color system

Pillars of Creation (NIRCam Image)

[VIEW ALL IMAGES >](#)



 Zoom Image

Download Options

Full Res (For Print), 8423 X 14589, TIF (163.43 MB) 

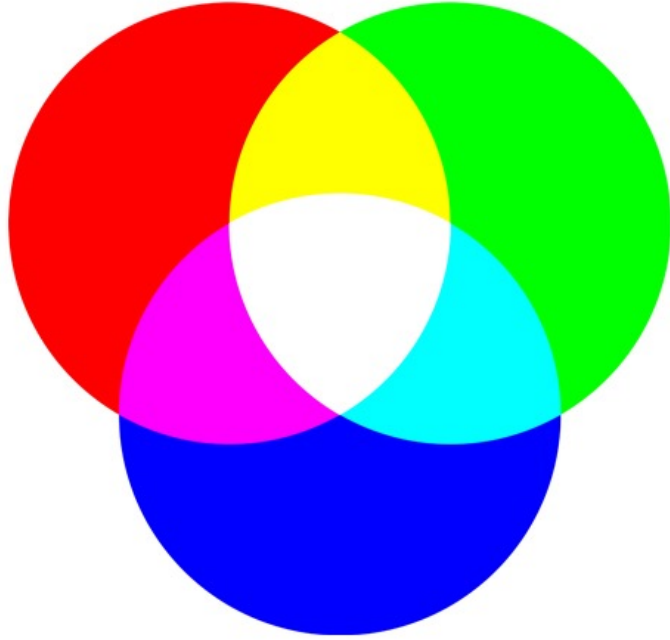
Full Res (For Display), 8423 X 14589, PNG (152.33 MB) 

1155 X 2000, PNG (4.73 MB) 

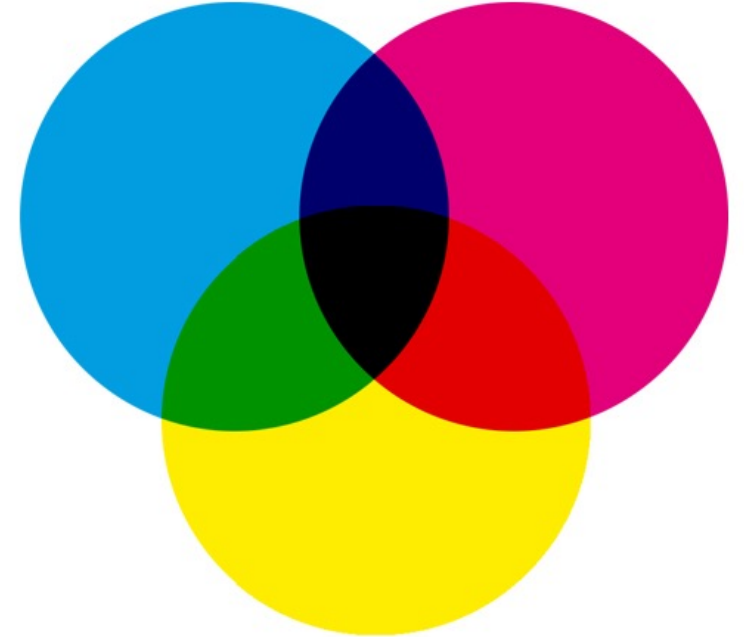
Why “For Display” and “For Print”?

Color system

What is the difference?



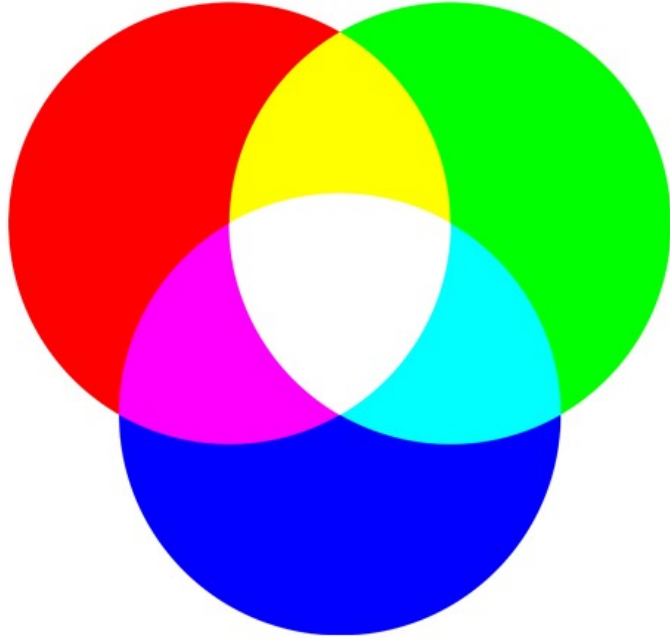
RGB



CMYK

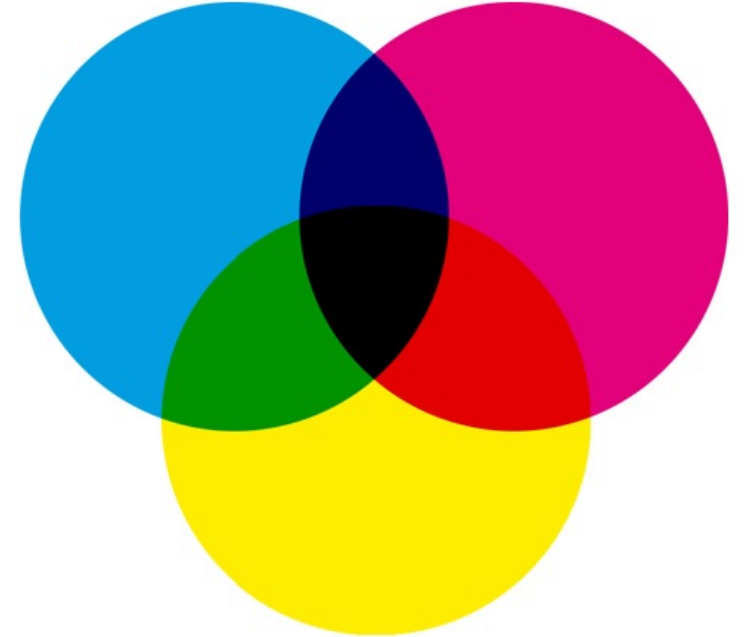
Color system

What is the difference?



RGB

Screens; Additive



CMYK

Printing; Subtractive

<https://www.epackprinting.com/cmyk-subtractive-color-mixing/>

Color system

What is the difference?



Screens; Additive



Printing; Subtractive

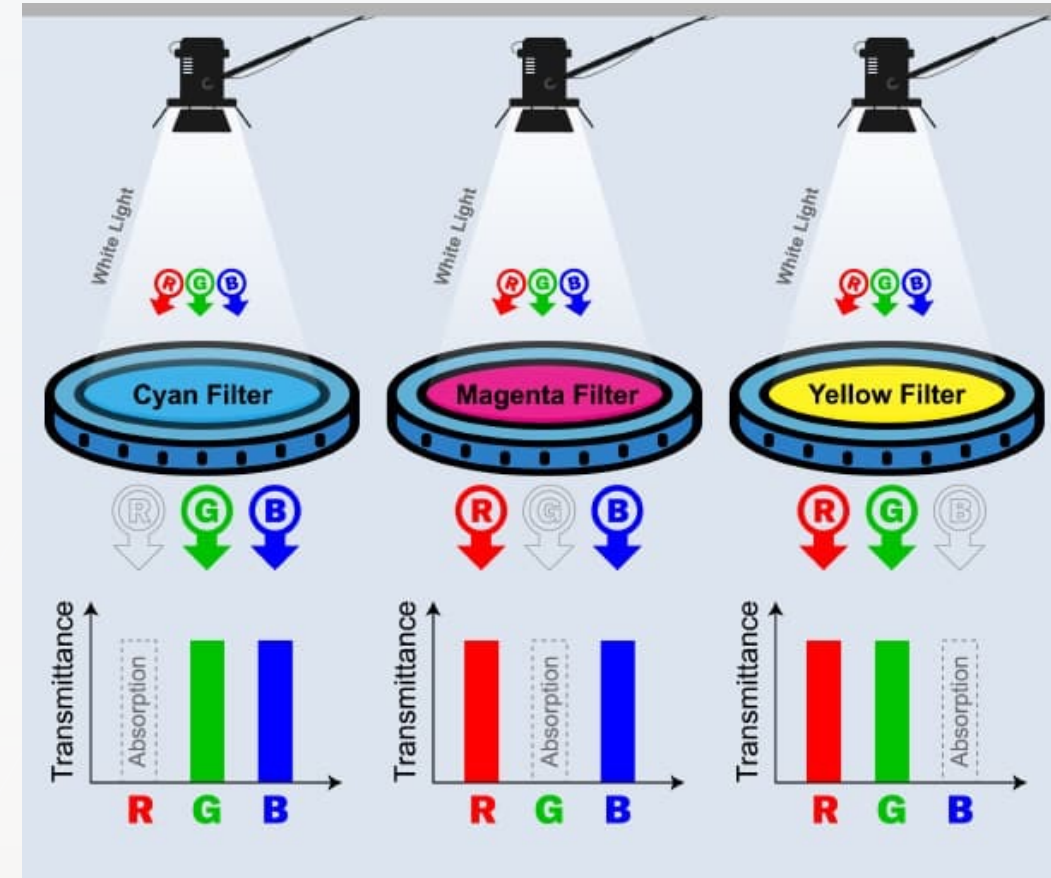
<https://www.epackprinting.com/cmyk-subtractive-color-mixing/>

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<https://www.epackprinting.com/cmyk-subtractive-color-mixing/>