Fast GPU-powered and auto-differentiable forward modeling for cosmological hydrodynamical simulations

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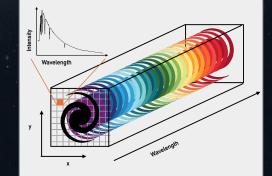




Observations of galaxies

GECKOS

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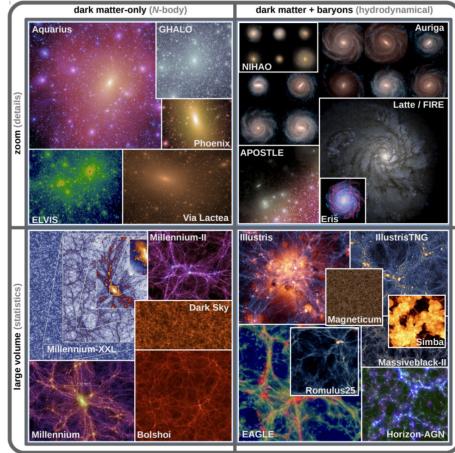
Cosmological Simulations

• Observations produce a lot of data

 \rightarrow interpretation difficult

- Simulate galaxy formation using computational methods
- Assume cosmology, initial conditions and physical processes, then let it run

Good overview: Vogelsberger et al. 2019



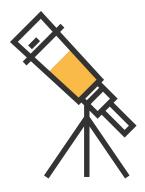
Vogelsberger et al. 2019

Methods of comparing simulations and observations

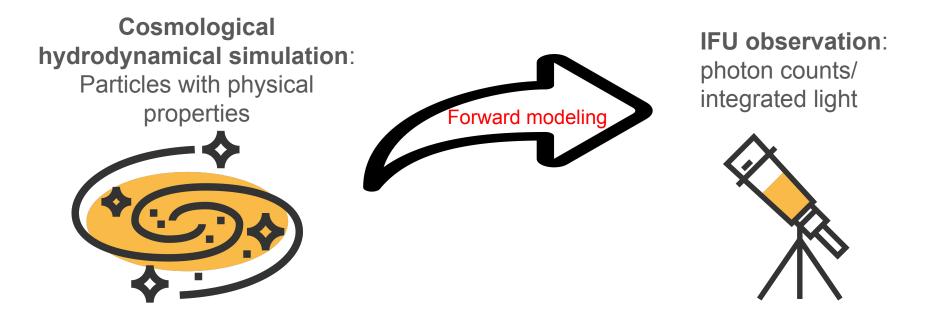
Cosmological hydrodynamical simulation: Particles with physical properties



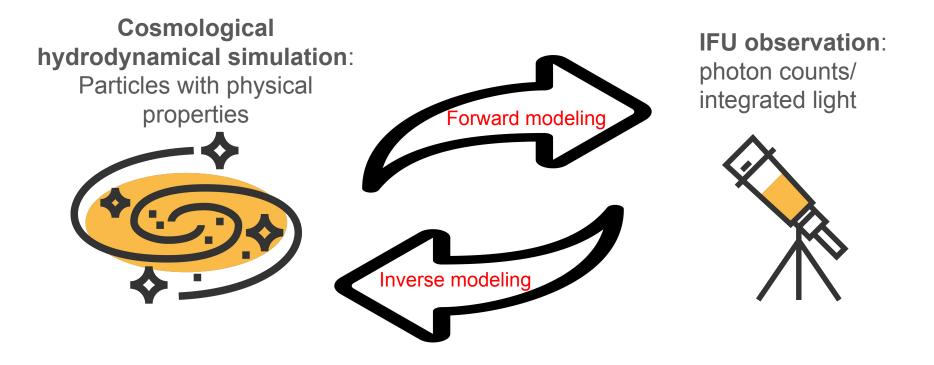
IFU observation: photon counts/ integrated light



Methods of comparing simulations and observations

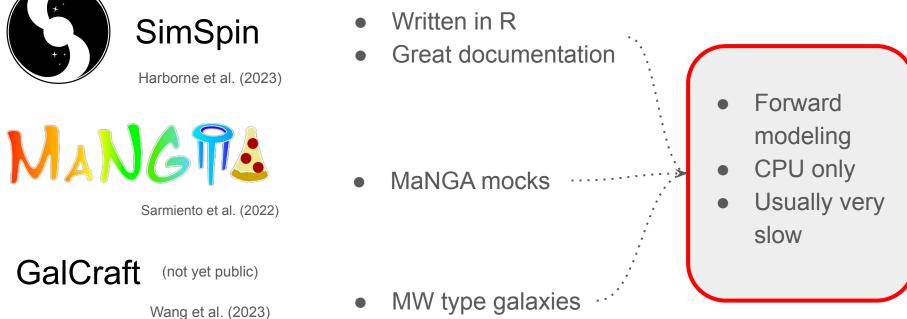


Methods of comparing simulations and observations



Other codes





Virtual telescope: RUBIX



What is different to already existing software

1. Computation time

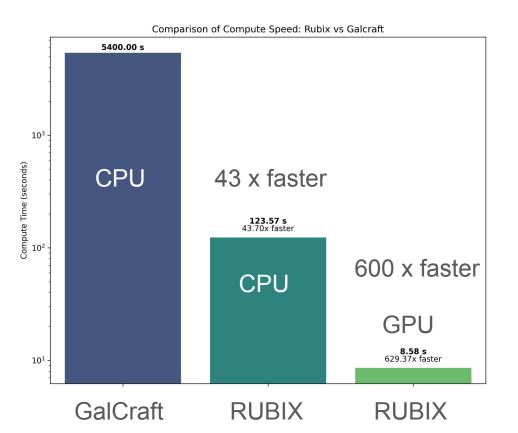
2. Code structure: research software practice for good open source software

3. Forward modeling and inverse modeling

4. Applications in machine learning

1. Computing time

Speed Comparison



GalCraft (Wang et al 2023):

- 6 × 10⁶ particles approx. 1.4 hours
- 24 Core CPU (2.5Ghz)

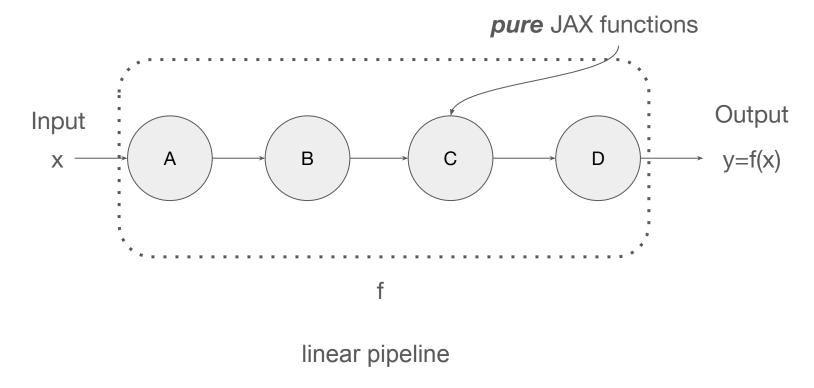
Rubix: 5.8 × 10⁶ particles

- 24-Core CPU: ~120s (40x)
- NVIDIA A100: ~10s (600x)

2. Code structure

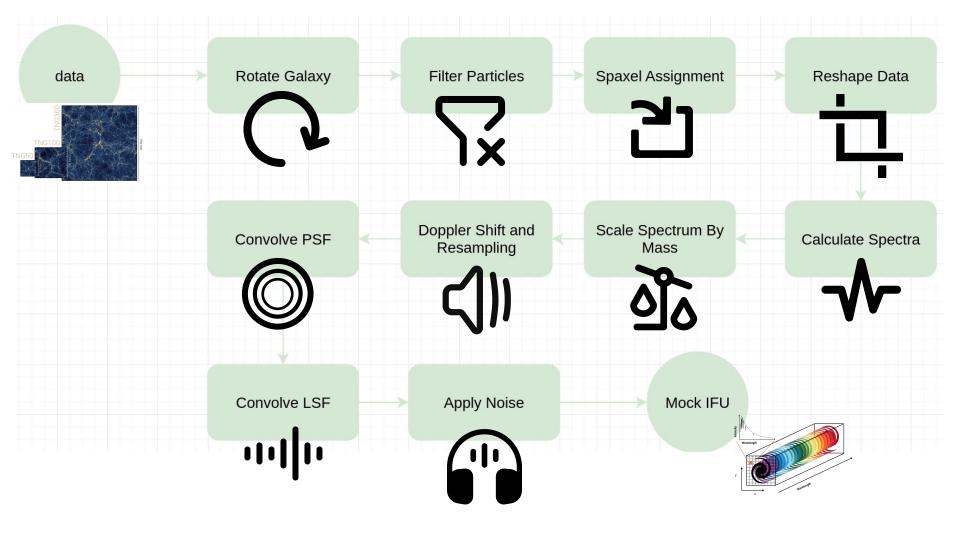


Code structure behind RUBIX



3.a Forward modeling





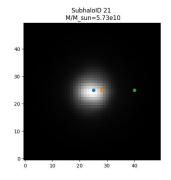
How to run the pipeline?

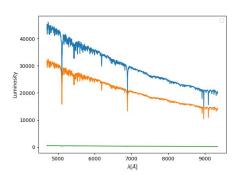
- 1. User configuration to specify e.g telescope, distance to galaxy, cosmology, SSP template
- 2. (Pipeline configuration)
- 3. Then simply run ...

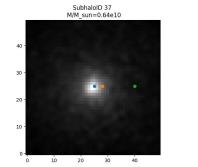
```
50 pipe = RubixPipeline(config) # Setup
51 data= pipe.run() # Run forward model
```

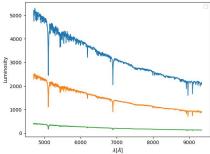
... and analyze the mock-data

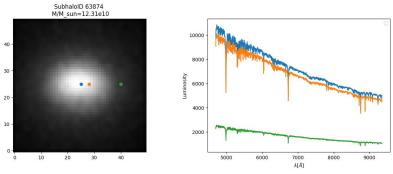
```
from rubix.core.pipeline import RubixPipeline
2 import os
 config = {
      "pipeline":{"name": "calc_ifu"},
      "logger": {
          "log_level": "INFO".
          "log_file_path": None,
          "format": "%(asctime)s - %(name)s - %(levelname)s -
     %(message)s".
      },
      "data": {
           "output_path": os.path.join(os.getcwd(), "output"),
          "save_name": "tng_14".
          "subset": {
               "use_subset": True.
              "subset_size": 10.
          }.
           "simulation": {
               "name": "IllustrisAPI".
               "args": {
20
21
                   "api_key": os.environ.get("ILLUSTRIS_API_KEY
      "),
22
                   "particle_type": ["stars"].
23
                   "simulation": "TNG50-1".
                   "snapshot": 99.
24
                   "galaxy_id": 14.
                   "reuse": True.
27
                           }.
29
30
      }.
      "telescope":
          {"name": "MUSE",
32
           "psf": {"name": "gaussian", "size": 5, "sigma":
33
     0.6}.
           "lsf": {"sigma": 0.5},
34
           "noise": {"signal_to_noise": 1, "noise_distribution"
      : "normal"}.}.
      "cosmology":
36
          {"name": "PLANCK15"},
37
38
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39
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-20
           "rotation": {"type": "edge-on"},
41
          },
42
43
3.1
      "ssp": {
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          }.
```



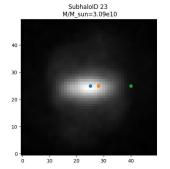


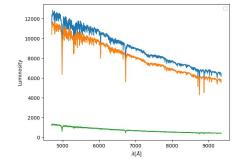












3.b Inverse modeling

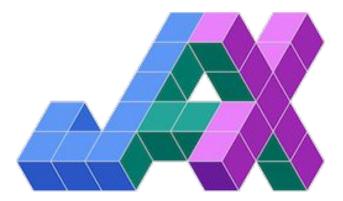


Inverse modeling

Use JAX auto differentiation

Toy example:

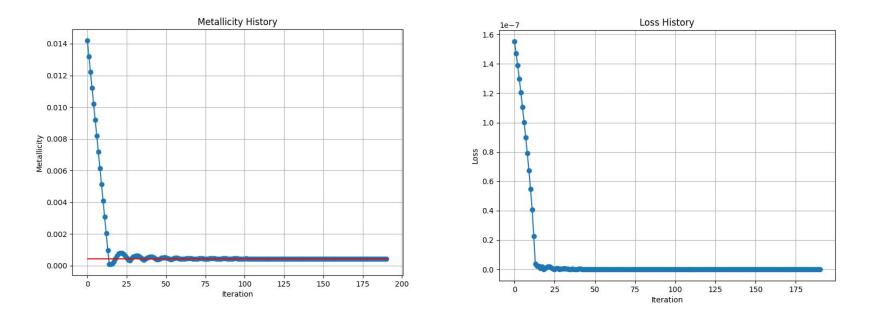
- one pixel
- one particle
- just vary the metallicity



Inverse modelling: Metallicity

At age 10 Gyr (100), start metallicity 1.4e-2 (9), target metallicity 4.5e-4 (3)

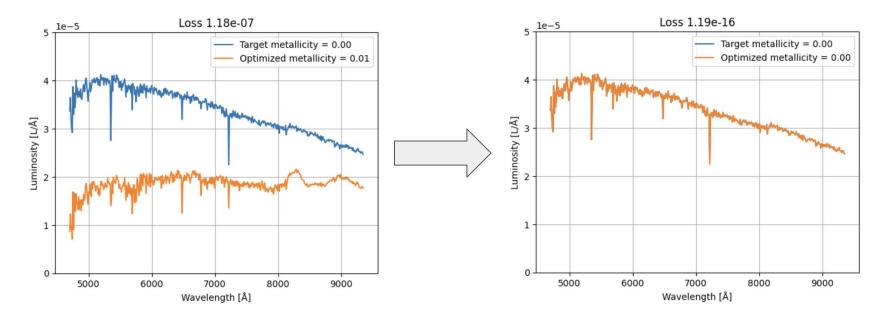
Learning rate 0.001, tol 1e-8



Inverse modelling: Metallicity

At age 10 Gyr (100), start metallicity 1.4e-2 (9), target metallicity 4.5e-4 (3)

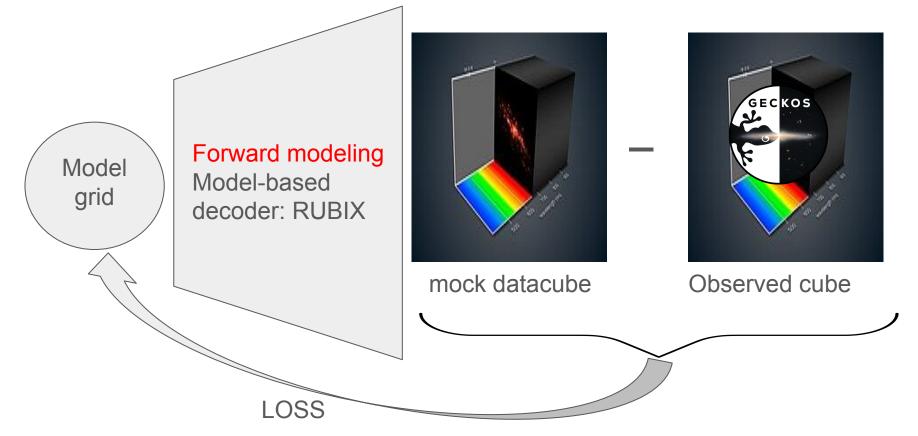
Learning rate 0.001, tol 1e-8



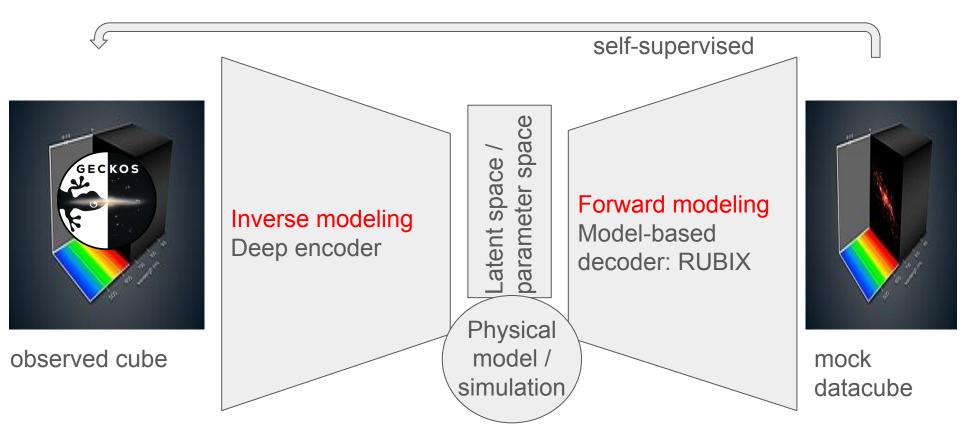
4. Machine learning applications



Gradient based optimization



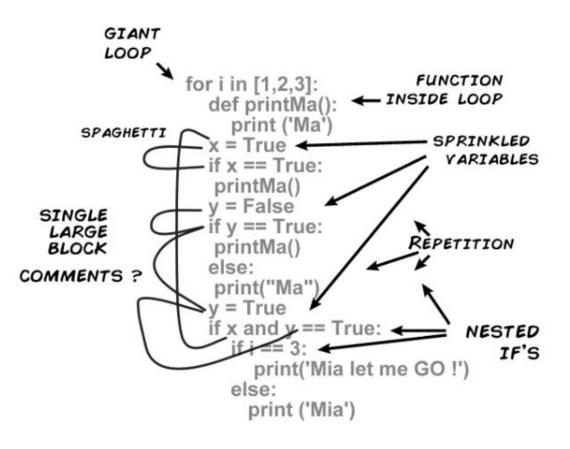
Self-supervised simulation based inference for GECKOS



Lessons learned from software development as PhD student

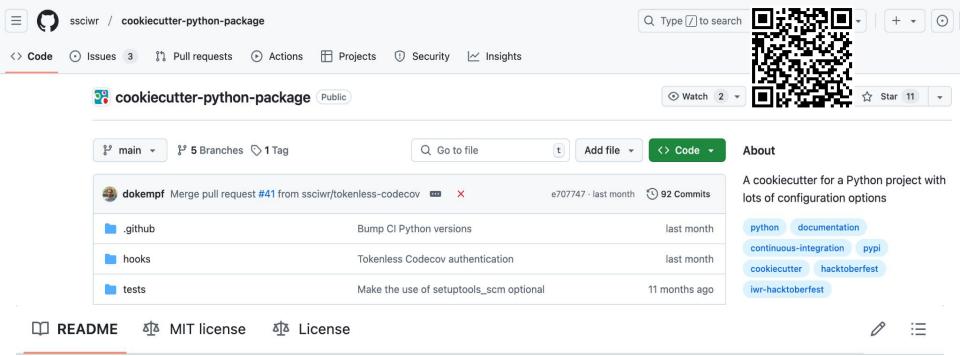


```
Spaghetti
Code
```





Konsultieren $ ightarrow$	Initieren Sie ein Projekt $ ightarrow$	Alle Projekte \rightarrow	Alle Kurse $ ightarrow$
PhD Program \rightarrow	Unser Team \rightarrow	Offene Stellen \rightarrow	Kontakt →



Welcome to Python Package Cookiecutter!

This repository is a template repository (a cookiecutter) that allows you to quickly set up new Python packages. It is geared towards scientific applications and implements the best practice guidelines of the <u>Scientific Software</u> <u>Center of Heidelberg University</u>.

SSC PROGRAM "SSC FELLOWS"

The Scientific Software Center is establishing a new mentoring program "SSC Fellows". This competitive program replaces the previous mentoring program "Reproducible Science".

The program

SSC fellows receive individual training in state-of-the-art software engineering and tools for improving their research software development skills. Each fellow is assigned an RSE mentor from the SSC with whom they will meet on a monthly basis, to track and support their research software engineering project. Fellows will further engage in community events and educational efforts organized by the SSC. In addition, fellows will receive a travel stipend to participate in a national RSE conference, with the opportunity to engage in the research software engineering community, and present their contribution.

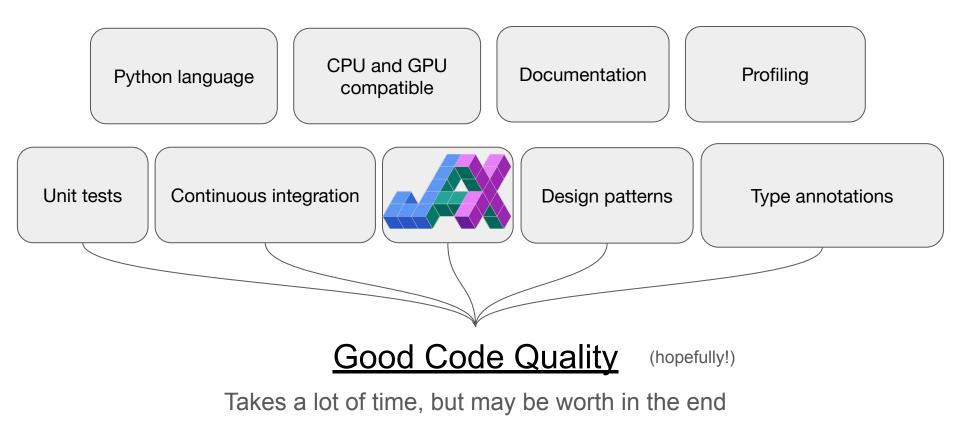
Eligibility and Selection Criteria

The SSC fellowship is limited to graduate students and postdoctoral researchers at Heidelberg University. Research software development should play a major role in the applicant's PhD /postdoctoral project. The SSC fellowships are an equal opportunity program open to all qualified persons without regard to race, gender, religion, age, physical disability, or national origin. SSC fellows will be selected by the SSC Advisory Board based on (1) the quality of the applicant's research software proposal and its relevance within the department/domain; (2) the applicant's research productivity, including previous



https://www.ssc.uni-heidelberg.de/en/learning/ssc-program-ssc-fellows

Guidelines we try to follow for RUBIX



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RUBIX documentation:

Welcome to RUBIX's documentation!

Installation

Code versions

Publications

License

Acknowledgments

Notebooks:

Create RUBIX data

Concept of the pipeline

Welcome to RUBIX's documentation!

RUBIX is a tested and modular Open Source tool developed in JAX, designed to forward model IFU cubes of galaxies from cosmological hydrodynamical simulations. The code automatically parallelizes computations across multiple GPUs, demonstrating performance improvements over state-of-the-art codes. For further details see the publications or the documentation of the individual functions.

Currently the following functionalities are provided:

- Generate mock IFU flux cubes for stars from IllustrisTNG50
- · Generate mock photometric images for stars for different filter curves
- Use different stellar population synthesis models
- · Use MUSE as telescope instrument (and some other instruments)

Currently the code is under development and is not yet all functionality is available. We are working on adding more features and improving the code, espectially we work on the following features:

- Adding support for more simulations
- Adding support for more telescopes
- · Adding gas emission lines and gas continuum
- Adding dust attenuation
- Adding support for gradient calculation

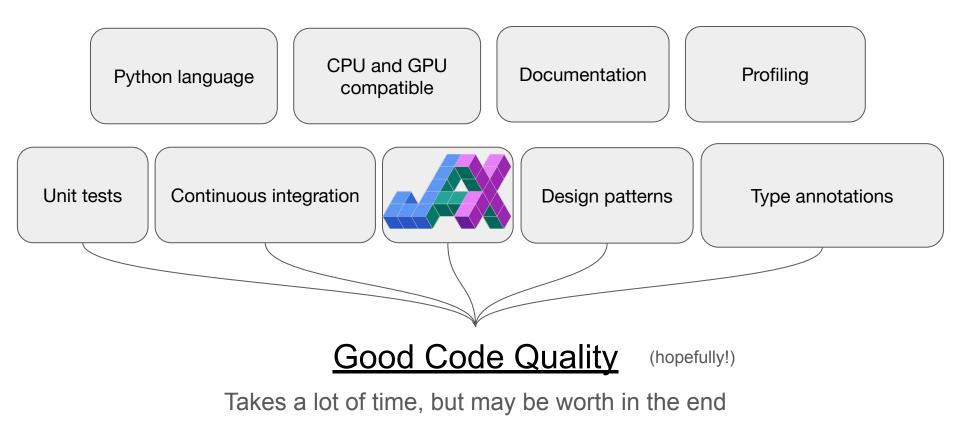
If you are interested in contributing to the code or have ideas for further features, please contact us via a github issue or via email. If you use the code in your research, please cite the following paper: <u>Publications</u>

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Welcome to RUBIX's documentation! Notebooks Code base documentation

Guidelines we try to follow for RUBIX



Pipeline run for 1000 stellar particles

Pipeline time 3.09 sec

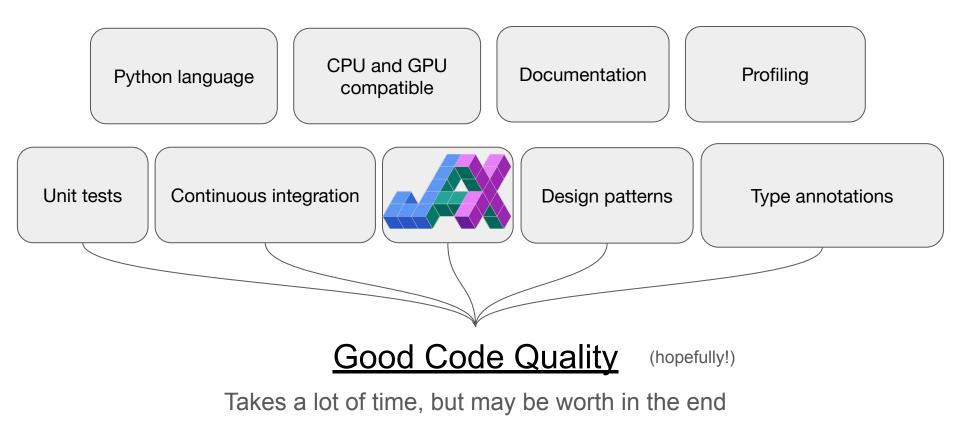
Data preparation in the beginning 1 sec

Calculate spectra function 463 millisec

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Guidelines we try to follow for RUBIX



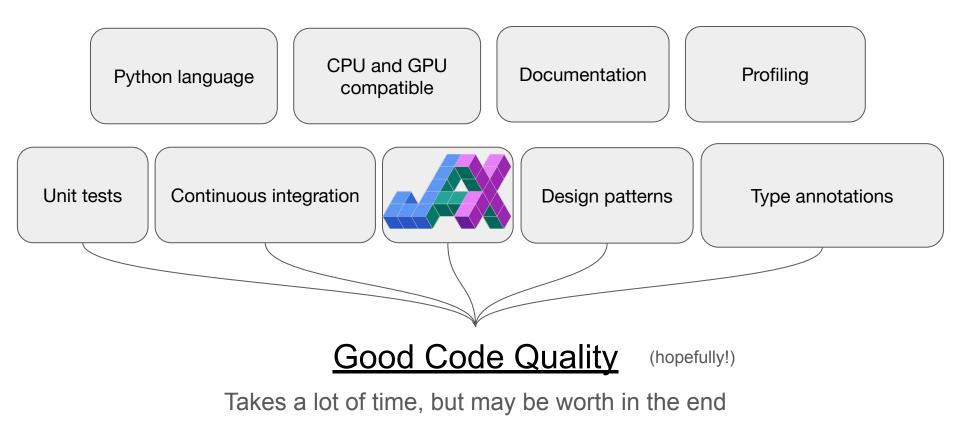


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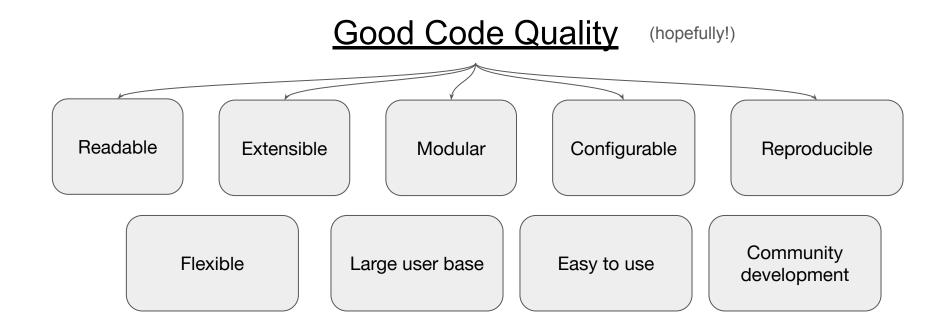


Guidelines we try to follow for RUBIX



	from typing import List, Optional, Union
	from jaxtyping import Int, Float, Array, jaxtyped
	from beartype import beartype as typechecker
	import numpy as np
	import equinox as eqx
	@jaxtyped(typechecker=typechecker)
	class BaseTelescope(eqx.Module):
11	
12	Base class for the telescope module.
13	This class contains the base parameters for the telescope module.
14	
15	Args:
	fov (float): The field of view of the telescope.
17	spatial_res (float): The spatial resolution of the telescope.
18	wave_range (list): The wavelength range of the telescope.
19	wave_res (float): The wavelength resolution of the telescope.
	lsf_fwhm (float): The full width at half maximum of the line spread function.
21	signal_to_noise (float): The signal to noise ratio of the telescope.
22	sbin (int): The size of the spatial bin in each direction for the aperture mask
23	aperture_region (jnp.ndarray): The aperture region of the telescope.
	pixel_type (str): The type of pixel used in the telescope.
25	wave_seq (jnp.ndarray): The wavelength sequence of the telescope.
	wave_edges (jnp.ndarray): The wavelength edges of the telescope.
27	
28	
29	fov: Union[float, int]
	<pre>spatial_res: Union[float, int]</pre>
31	<pre>wave_range: List[float] # upper and lower limits</pre>
32	wave_res: Union[float, int]
33	<pre>lsf_fwhm: Union[float, int]</pre>
34	<pre>signal_to_noise: Optional[float]</pre>
35	sbin: np.int64
	<pre>aperture_region: Union[Float[Array, ""], Int[Array, ""]]</pre>
37	pixel_type: str
	<pre>wave_seq: Float[Array, ""]</pre>
	wave_edges: Float[Array, ""]

Benefits of good code quality



What we can conclude



RUBIX

- Forward and inverse modeling of simulated galaxies and IFU data
- With help of the SSC Heidelberg: project developed to a good codebase
- Lots of effort into testing, integration, documentation
- Hopefully become a broadly used community tool



https://astro-rubix .web.app/

Thank you!