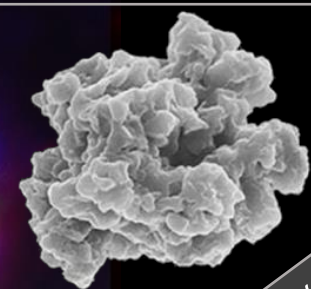
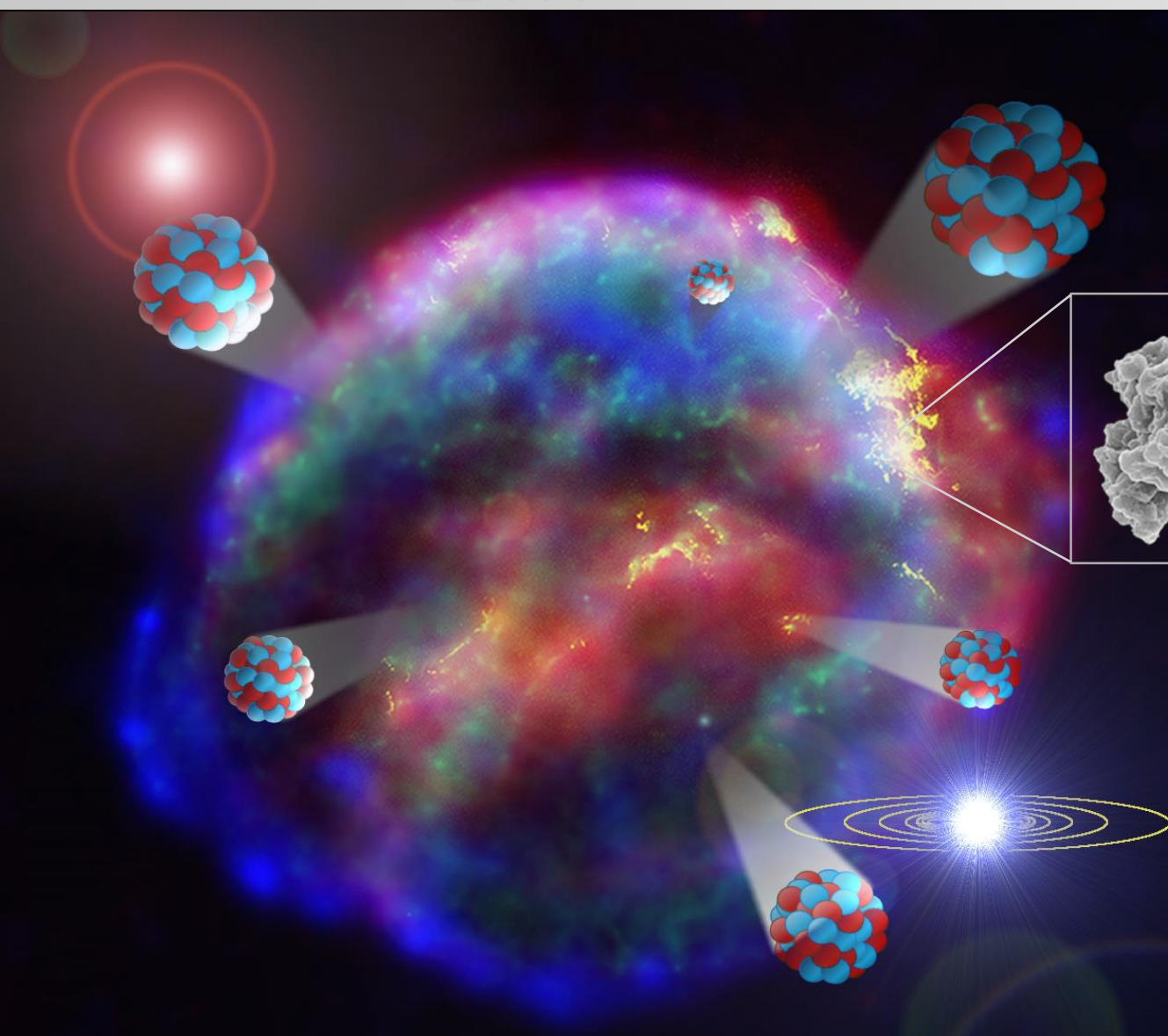


How to study stellar explosions with deep-sea sediments and ocean crusts?



Supernova-
Dust



The *Australopithecus afarensis* 2-3 million years ago



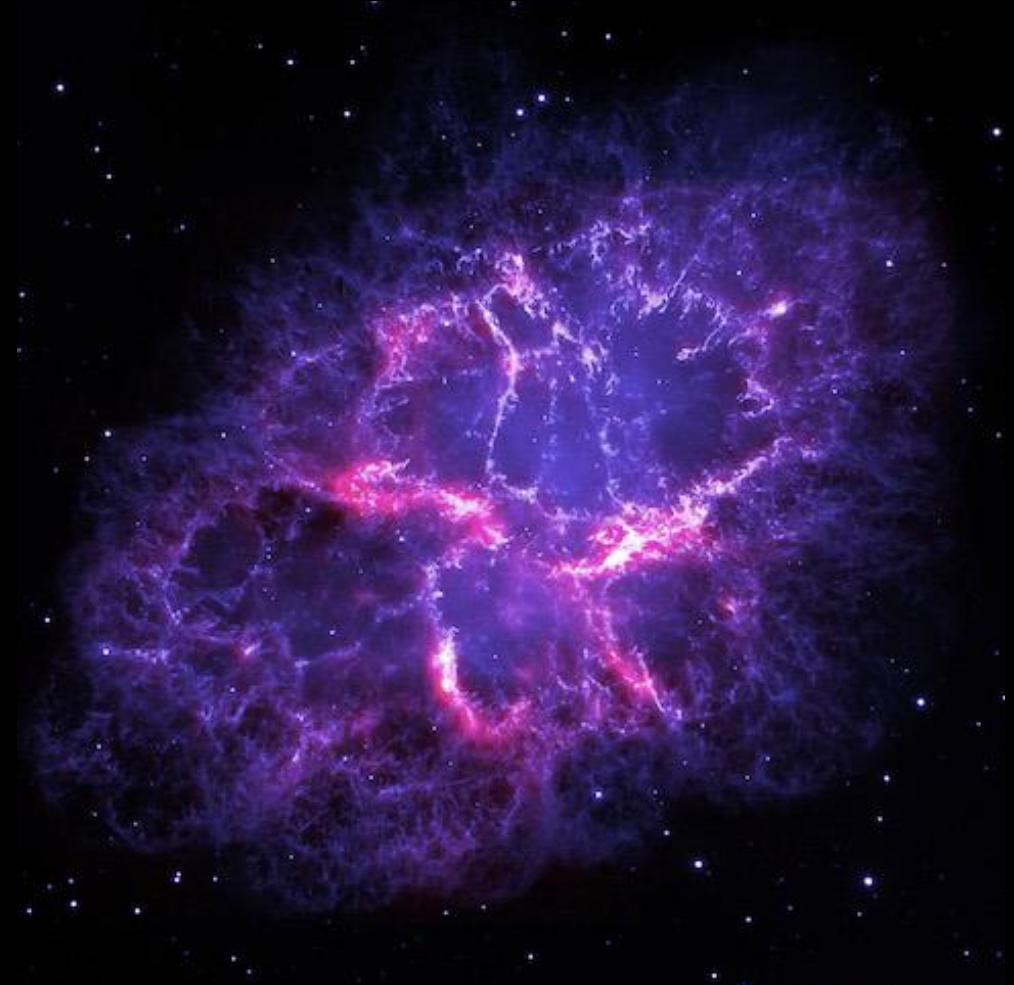
© Mark A. Garlick / No unauthorised usage

Supernova Remnants



Supernova Remnant SNR 0509-67.5
HST/ACS/WFC3 • CXO/ACIS

NASA, ESA, CXC, SAO, B. Schaefer and A. Pagnotta (Louisiana State University, Baton Rouge),
the Hubble Heritage Team (STScI/AURA), and J. Hughes (Rutgers University) STScI-PRC12-06a



Crab Nebula
Herschel (red) and Hubble (blue) composite

Credit: ESA/Herschel/PACS/MESS Key Programme Supernova Remnant Team; NASA, ESA
and Allison Loll/Jeff Hester (Arizona State University).

Observation of Supernova Remnants



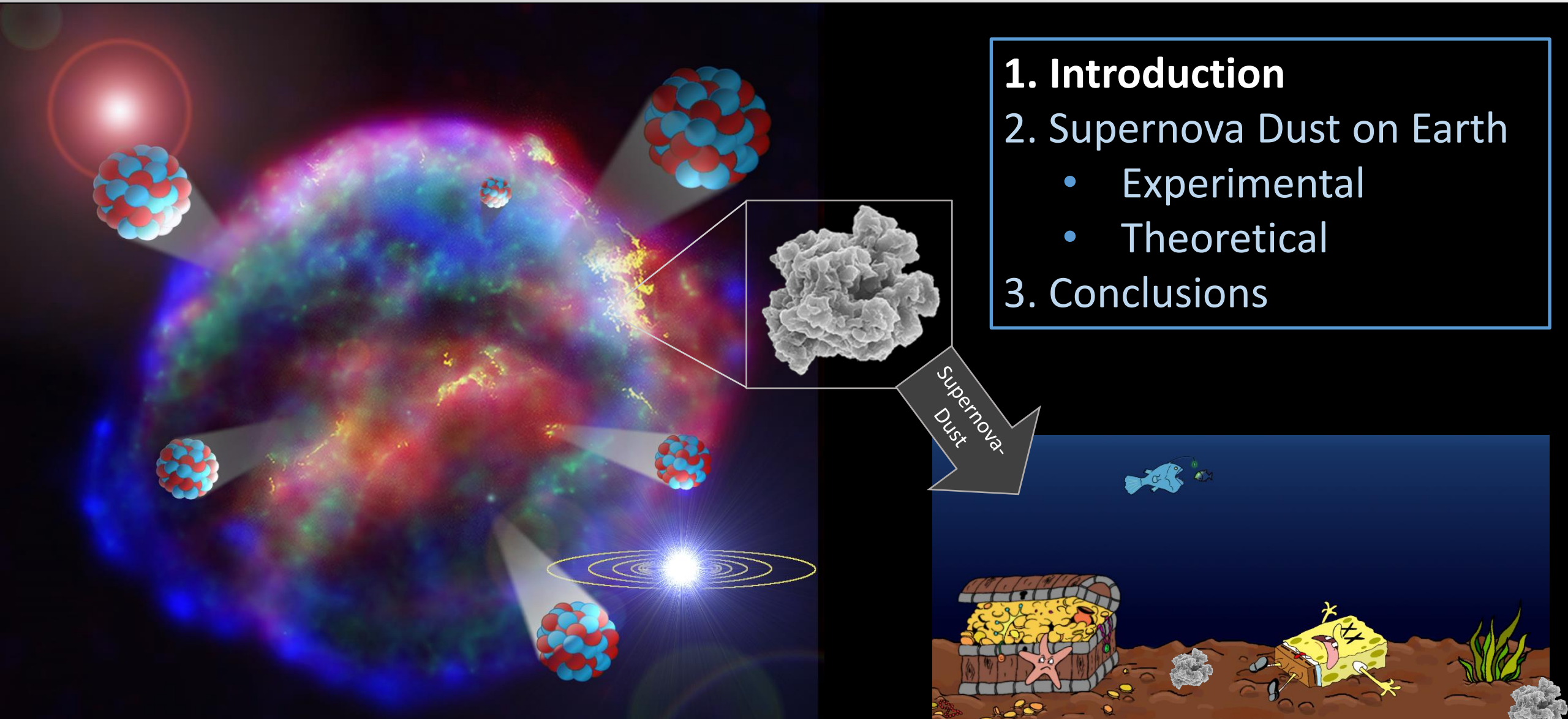
Outline

1. Introduction

2. Supernova Dust on Earth

- Experimental
- Theoretical

3. Conclusions



The Solar Environment

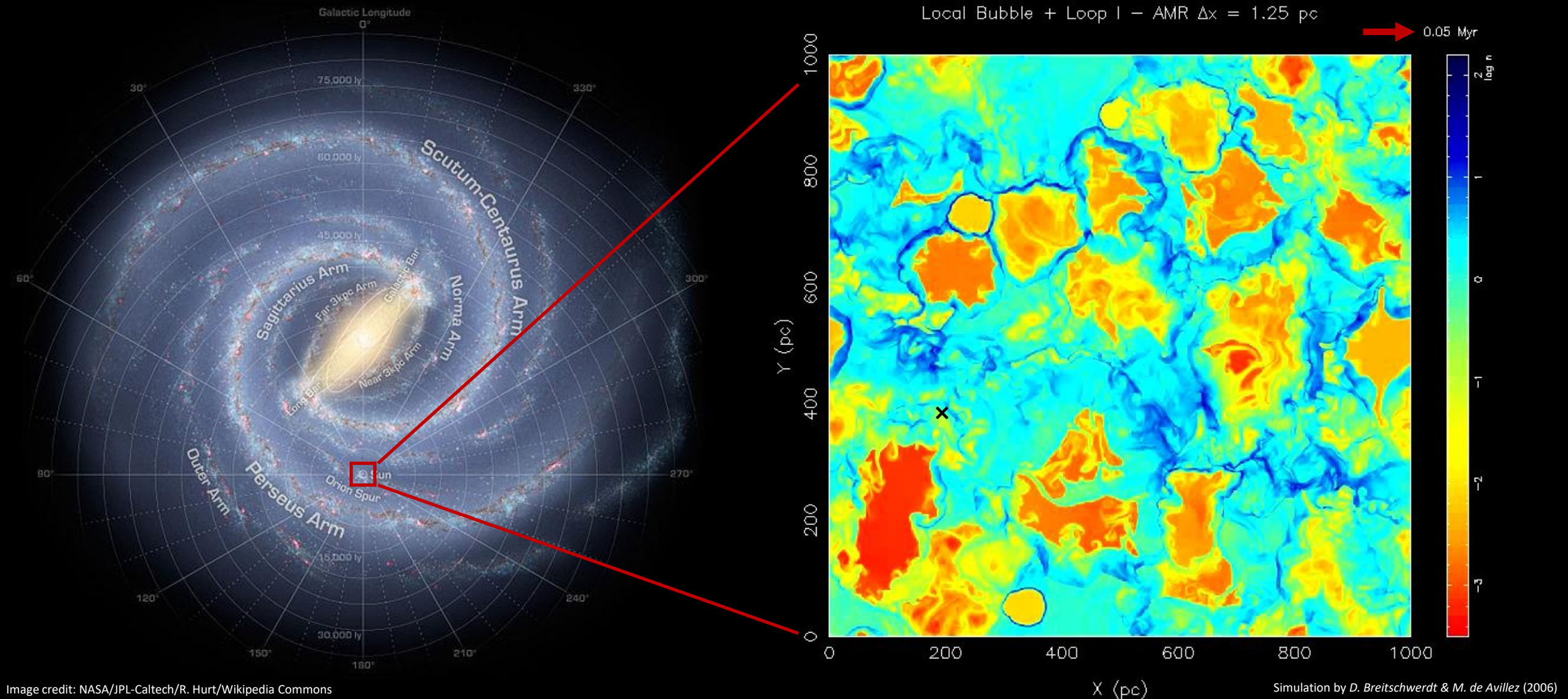
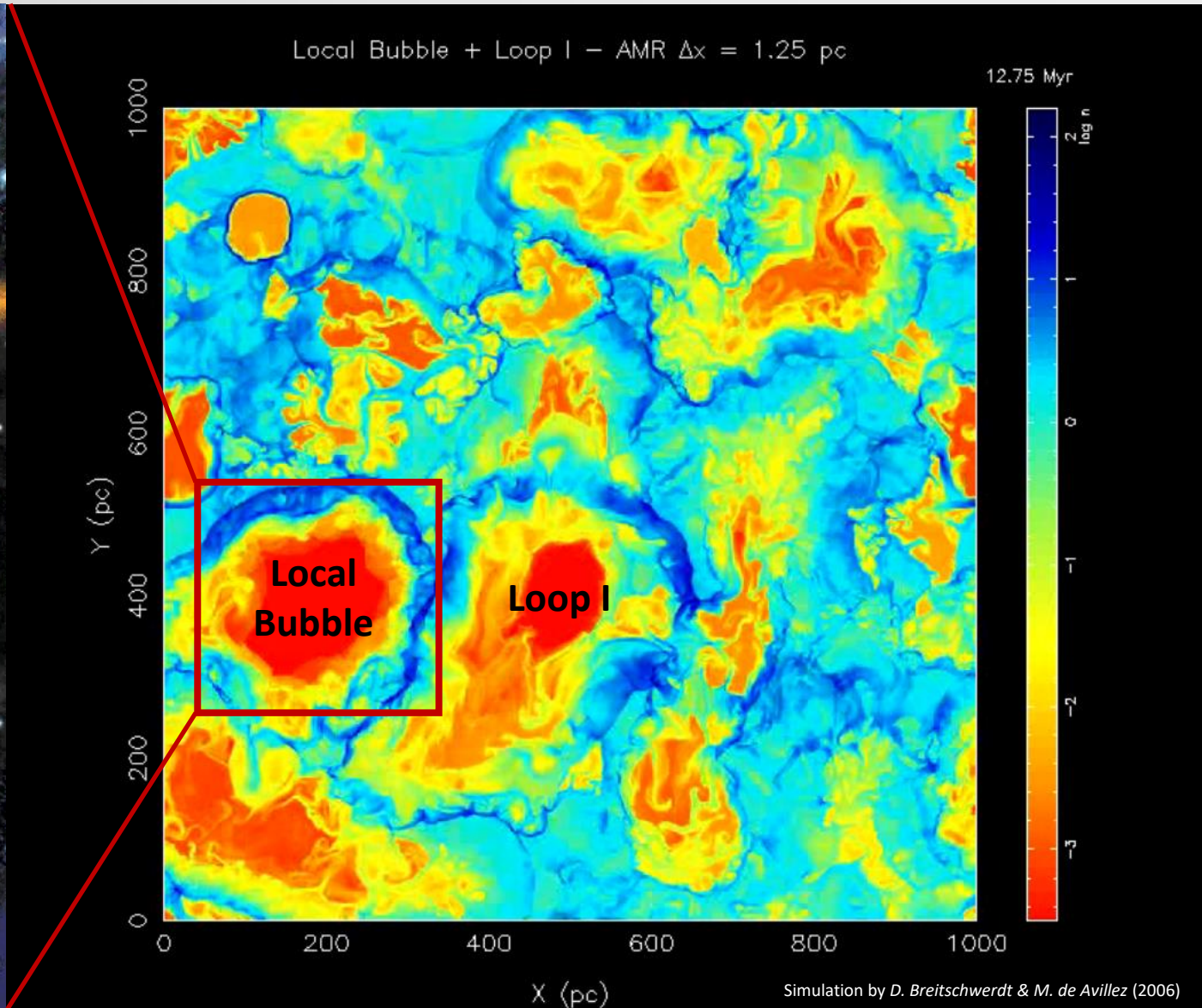


Image credit: NASA/JPL-Caltech/R. Hurt/Wikipedia Commons

The Solar Environment



The Solar Environment

Age: 14 Myr

Origin: 14-20 SNe

T = 10^6 K

n = 5×10^{-3} atoms cm^{-3}

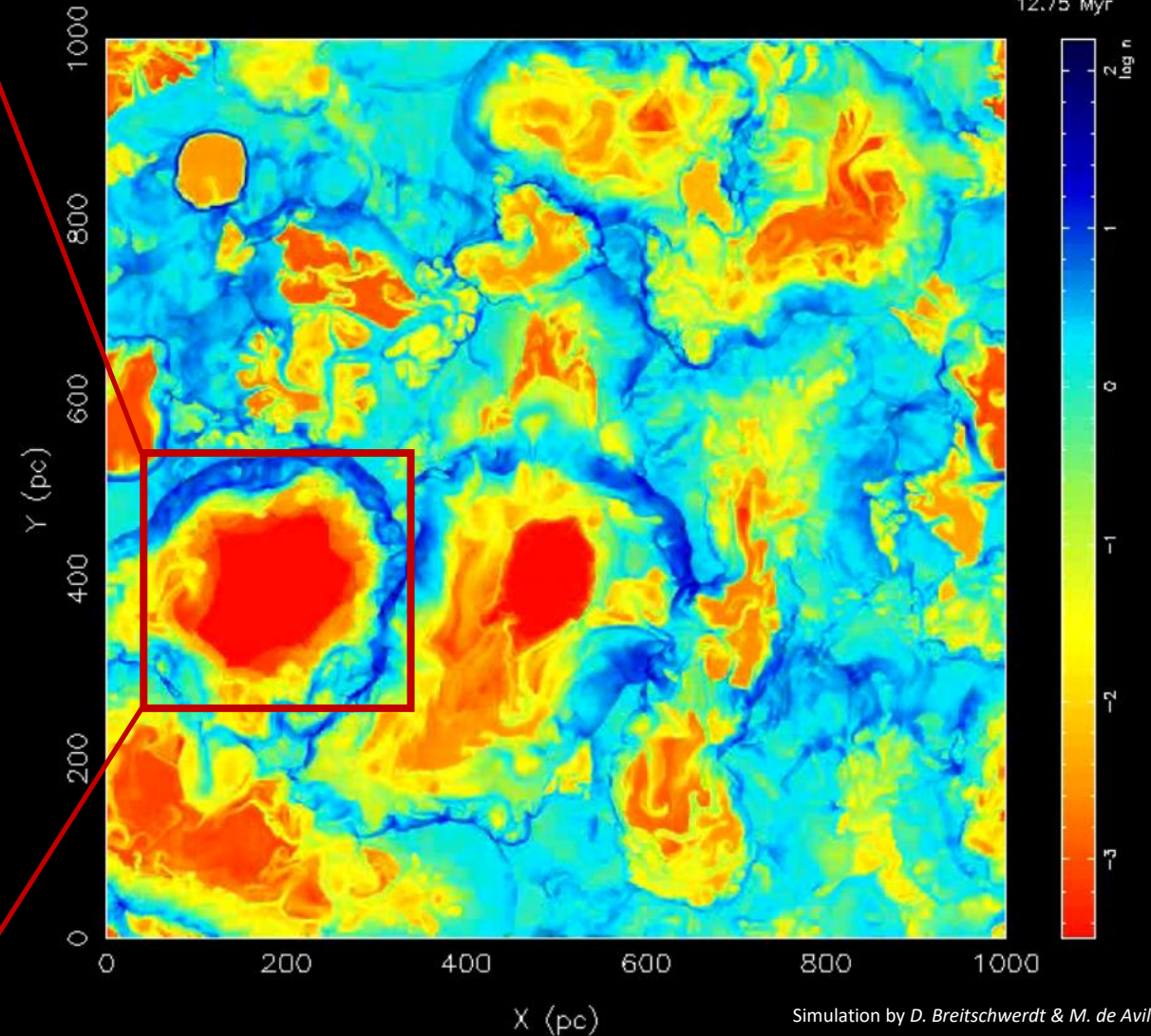
V = $200^2 \times 600$ pc³

Sun ☉

Local Bubble

Local Bubble + Loop I – AMR $\Delta x = 1.25$ pc

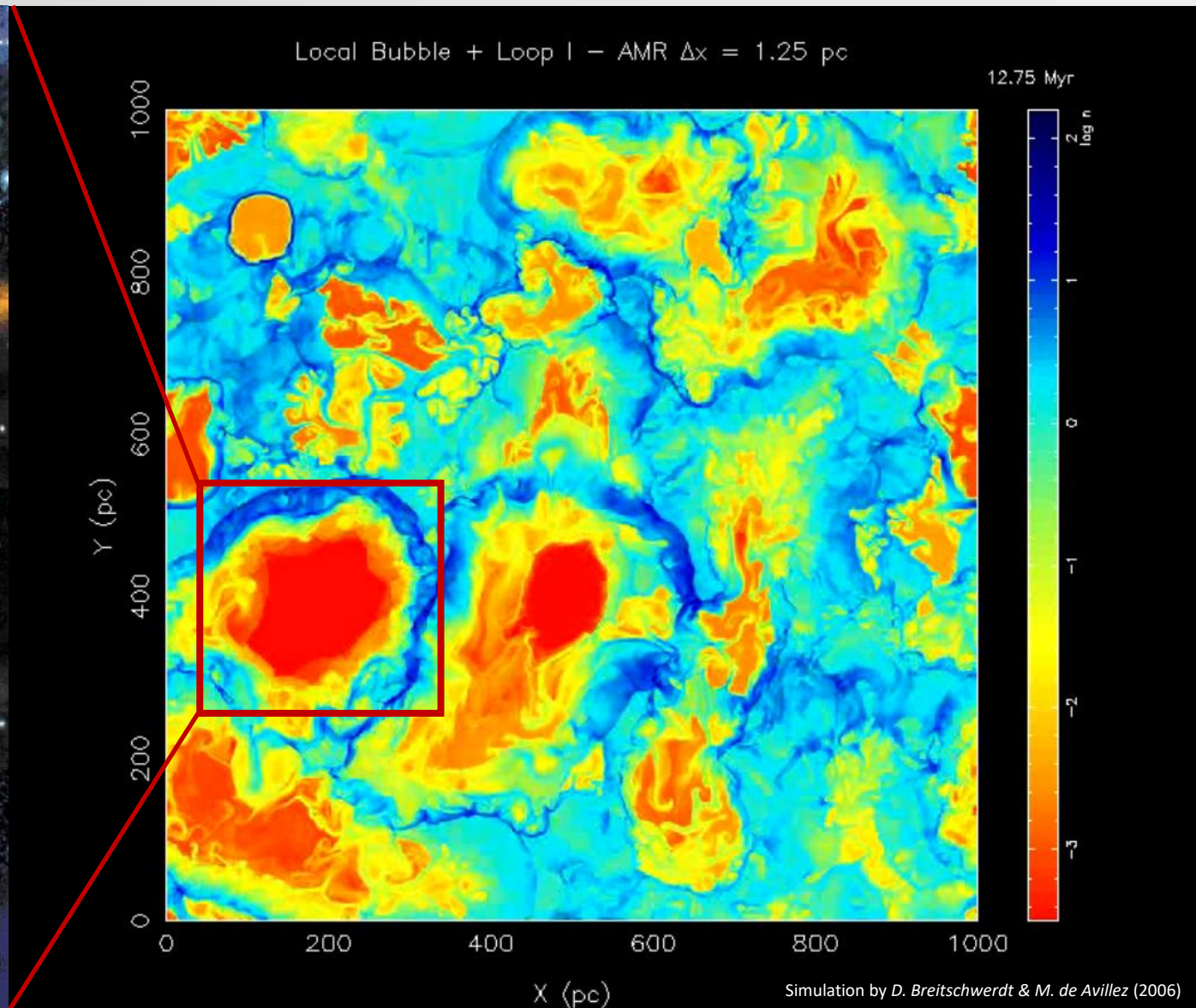
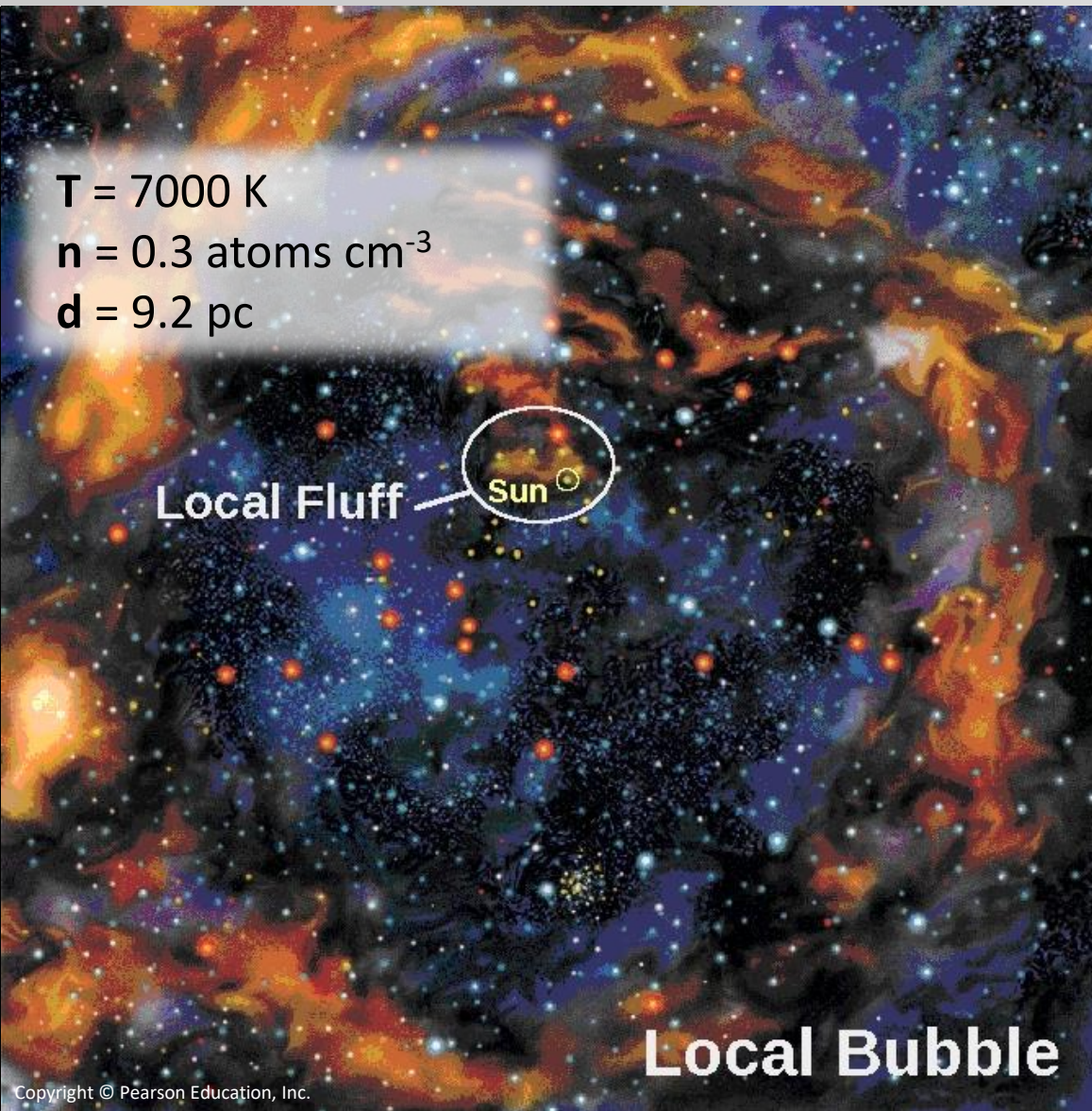
12.75 Myr



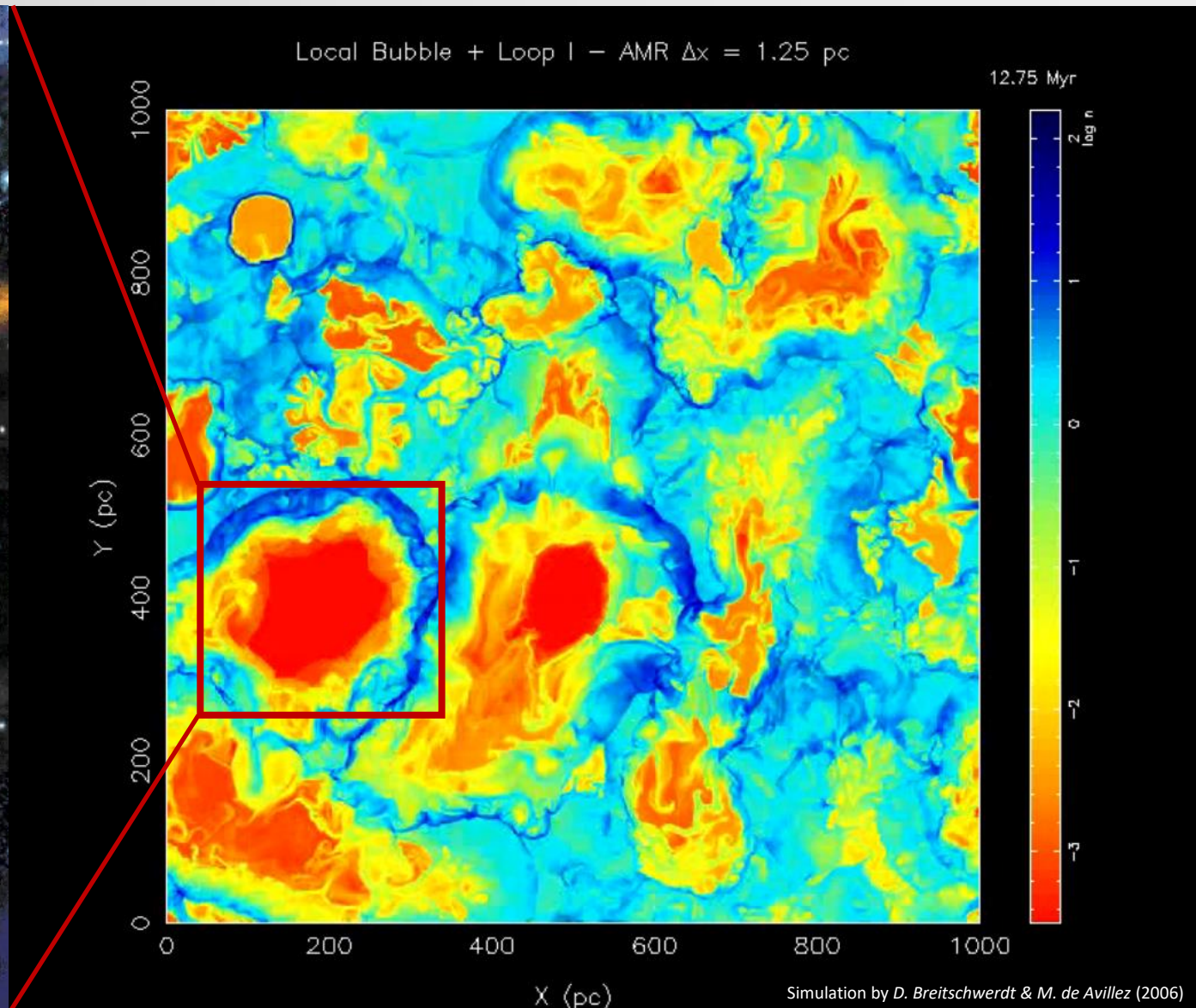
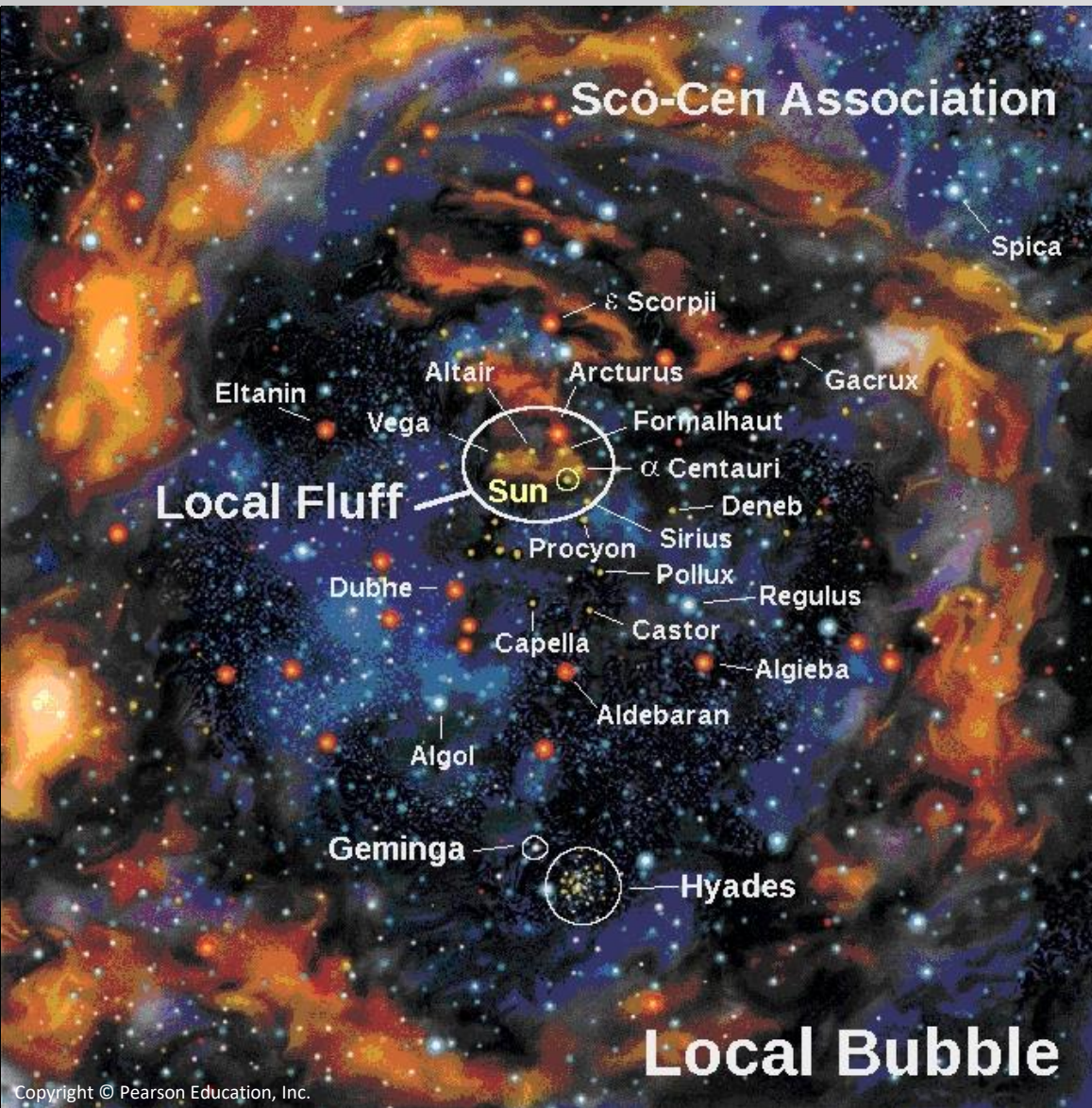
Simulation by D. Breitschwerdt & M. de Avillez (2006)

Copyright © Pearson Education, Inc.

The Solar Environment



The Solar Environment



Supernova Traces on Earth

Radiocarbon 38:68 (1996)

THE ASTROPHYSICAL JOURNAL, 470:1227–1236, 1996 October 20

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^{60}Fe , A PROMISING AMS ISOTOPE FOR MANY APPLICATIONS

G. KORSCHINEK, T. FAESTERMANN, K. KNIE and C. SCHMIDT

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**Half-lives in the order of
0.3-80 million years!**

GEOLOGICAL ISOTOPE ANOMALIES AS SIGNATURES OF NEARBY SUPERNOVAE

JOHN ELLIS

Theoretical Physics Division, CERN, Geneva, Switzerland

BRIAN D. FIELDS¹

Department of Physics, University of Notre Dame, Notre Dame, IN 46556

AND

DAVID N. SCHRAMM²

University of Chicago, 5640 South Ellis Avenue, Chicago, IL 60637

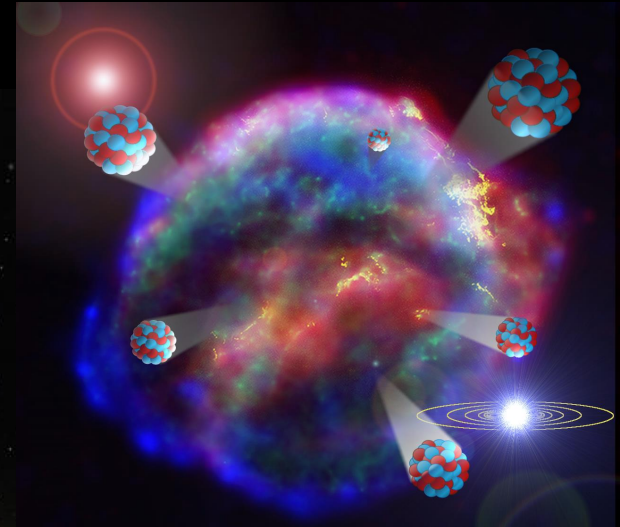
Received 1995 June 15; accepted 1996 May 21

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Solar System Formation 4.57 Gyr ago

Radioactive decay since Solar System formation
Detection on Earth: **Recent influx!**

Half-lives in the order of
0.3-80 million years!



Implications for Earth

Radiocarbon 38:68 (1996)

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Nucleosynthesis

^{60}Fe , $t_{1/2} = 2.6 \text{ Myr}$

(Rugel et al. 2009, Wallner et al. 2015)

- s- and r-process**

- (S)AGB stars, Core-collapse supernovae, Electron-capture supernovae

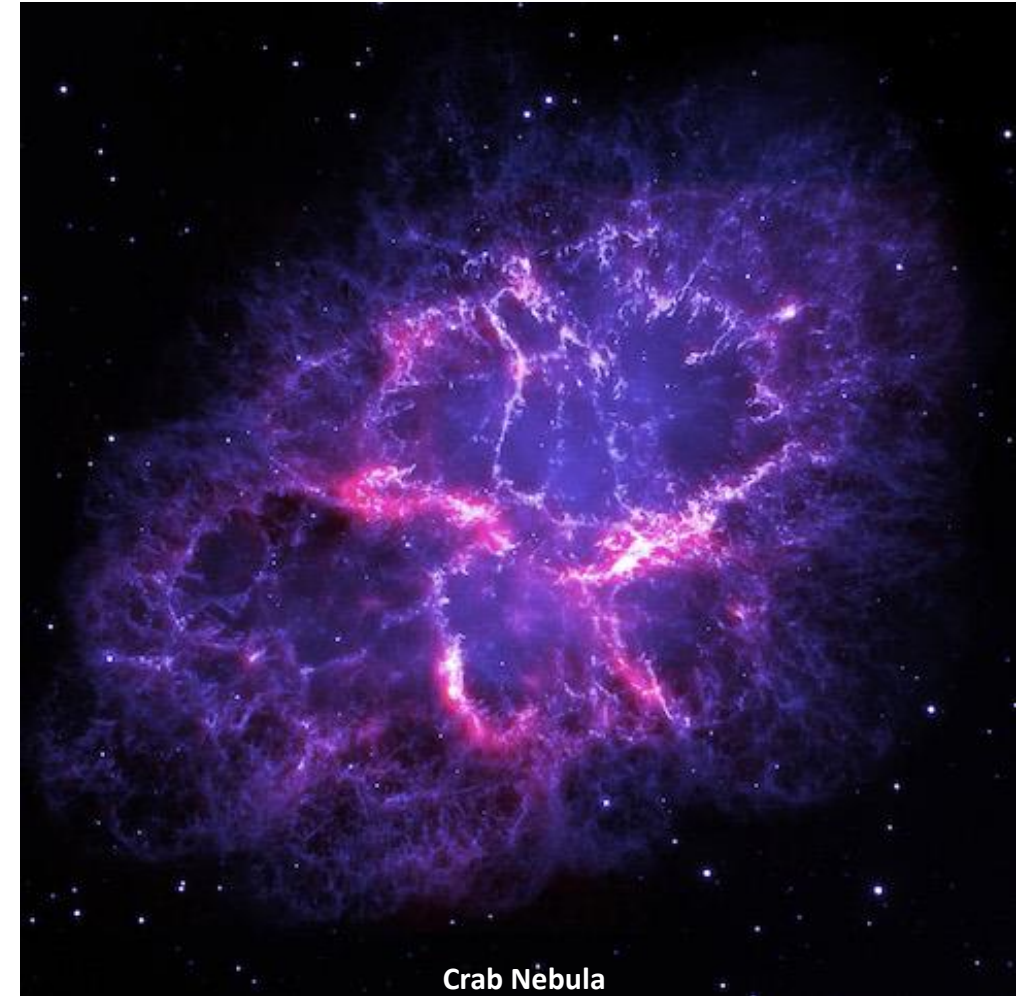
Ni 60 26.2231 σ 2.9	Ni 61 1.1399 σ 2.5 $\sigma_{n, \alpha}$ 0.00003	Ni 62 3.6345 σ 15	Ni 63 100 a β^- 0.07 $n\beta^- \gamma$ σ 20
Co 59 100 σ 20.7 + 16...	Co 60 10.5 m 5.272 a β^- 0.3... γ 1332... σ 58	Co 61 1.65 h β^- 1.2... γ 67; 909...	Co 62 14.0 m 1.5 m β^- 2.9... γ 1173; 1163; 2003...
Fe 58 0.282 σ 1.3	Fe 59 44.503 d β^- 1.6... γ 109; 129...	Fe 60 $1.5 \cdot 10^6 \text{ a}$ σ 0.1	Fe 61 6.0 m β^- 2.8... γ 1205; 1027; 298...

^{26}Al , $t_{1/2} = 0.7 \text{ Myr}$

- $^{25}\text{Mg}(p, \gamma)$ reaction**

- (S)AGB stars, Core collapse supernovae, Wolf-Rayet stars

Si 26 2.21 s β^+ 3.8... γ 829; 1622... m	Si 27 4.16 s β^+ 3.1... γ (22...)	Si 28 92.223 σ 0.17
Al 25 7.18 s β^+ 3.3... γ (1612...)	Al 26 6.35 s $716 \cdot 10^5 \text{ a}$ β^+ 1.2... γ 180... β^+ 3.2	Al 27 100 σ 0.230
Mg 24 78.99 σ 0.053	Mg 25 10.00 σ 0.20	Mg 26 11.01 σ 0.038

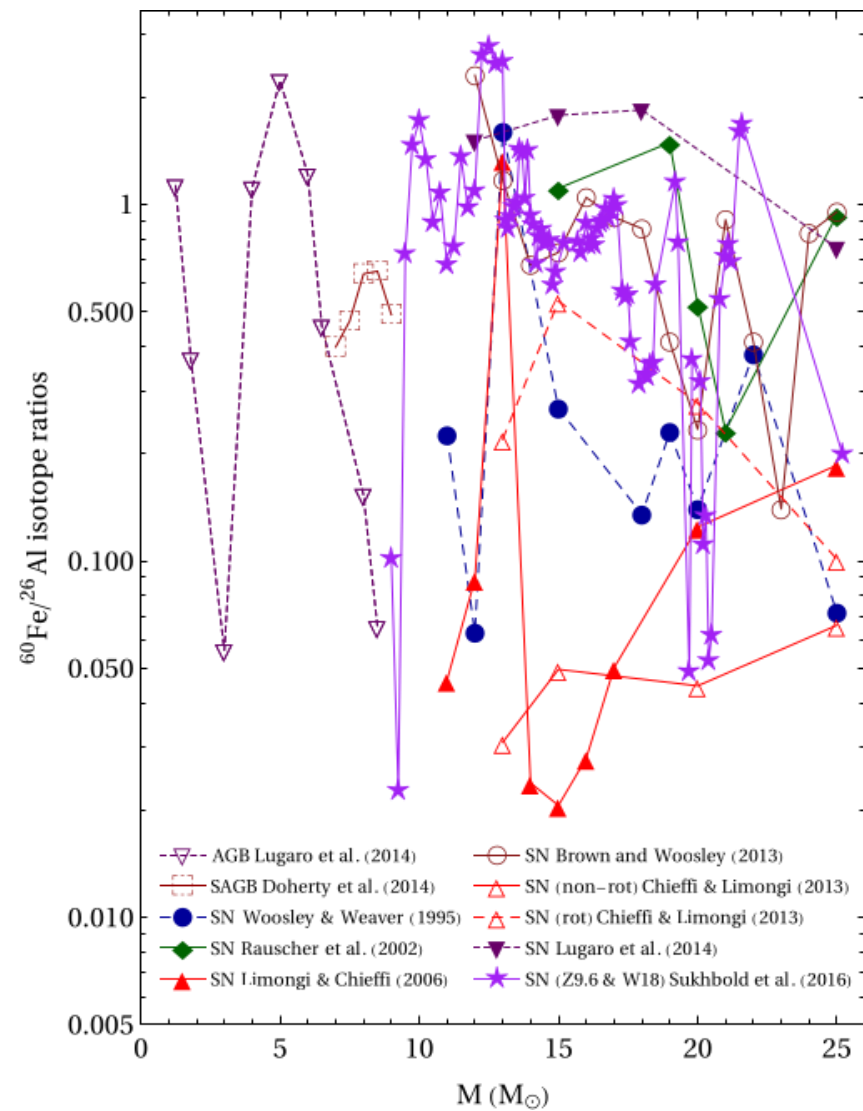


Crab Nebula

Herschel (red) and Hubble (blue) composite

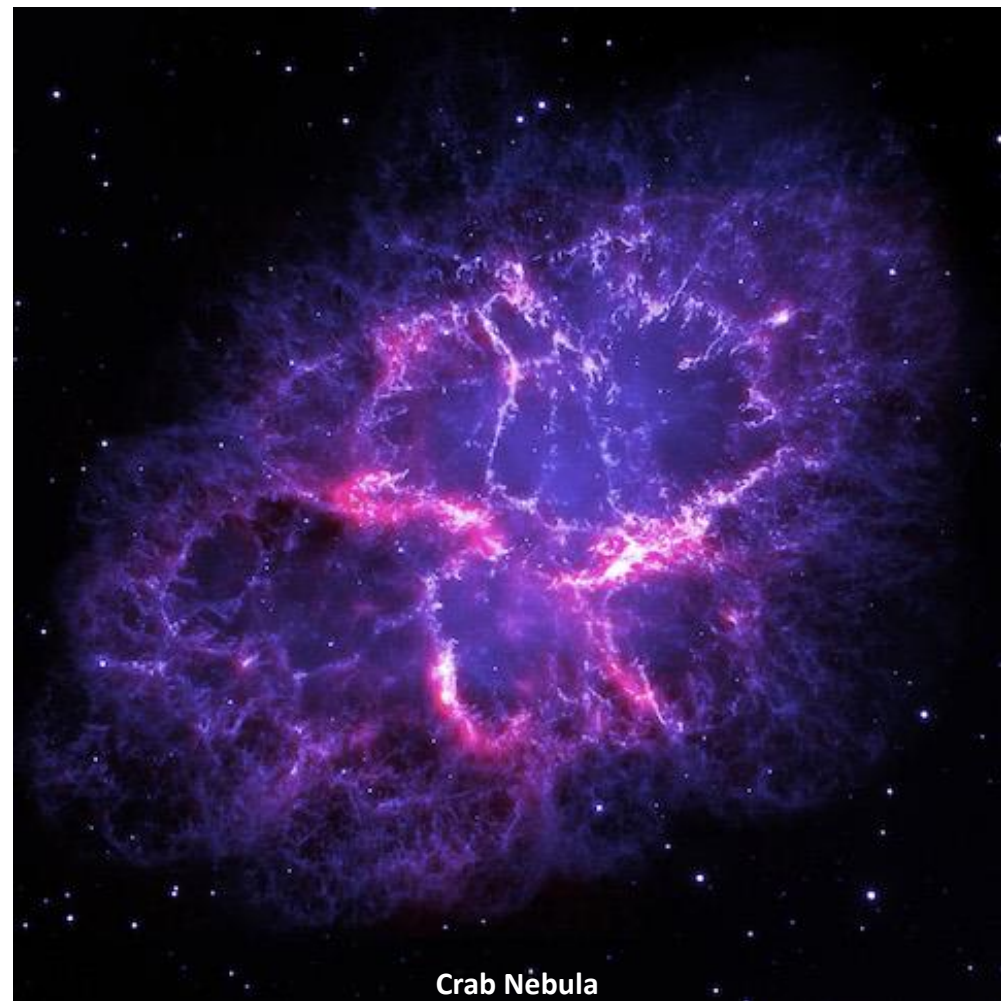
Credit: ESA/Herschel/PACS/MESSE Key Programme Supernova Remnant Team; NASA, ESA and Allison Loll/Jeff Hester (Arizona State University).

Nucleosynthesis



Ni 60 26.2231	Ni 61 1.1399	Ni 62 3.6345	Ni 63 100 a
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Co 59 100	Co 60 10.5 m 5.272 a	Co 61 1.65 h	Co 62 14.0 m 1.5 m
σ 20.7 + 16...	β^- 0.3... γ 1332... σ 58	β^- 1.2... γ 67; 909...	β^- 2.9... γ 1173; 1163; 2003... σ 1129...
Fe 58 0.282	Fe 59 44.503 d	Fe 60 1.5 · 10 ⁶ a	Fe 61 6.0 m
σ 1.3	β^- 1.6... γ 109; 1292... σ 13	σ 0.1	β^- 2.8... γ 1205; 1027; 298...

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β^+ 3.8... γ 829; 1622... m	β^+ 3.1... γ (22...)	σ 0.17
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β^+ 3.3... γ (1612...)	β^+ 3.2... γ 1.2... σ 1.8...	σ 0.230
Mg 24 78.99	Mg 25 10.00	Mg 26 11.01
σ 0.053	σ 0.20	σ 0.038

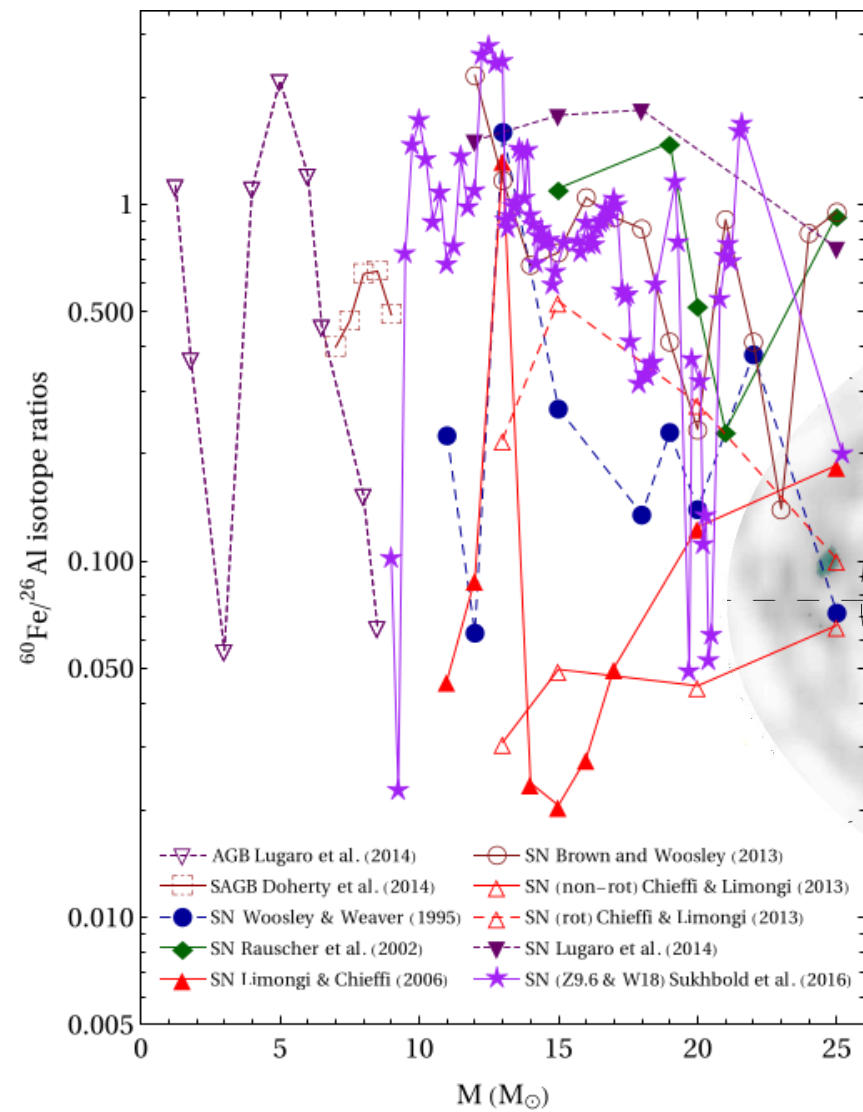


Herschel (red) and Hubble (blue) composite

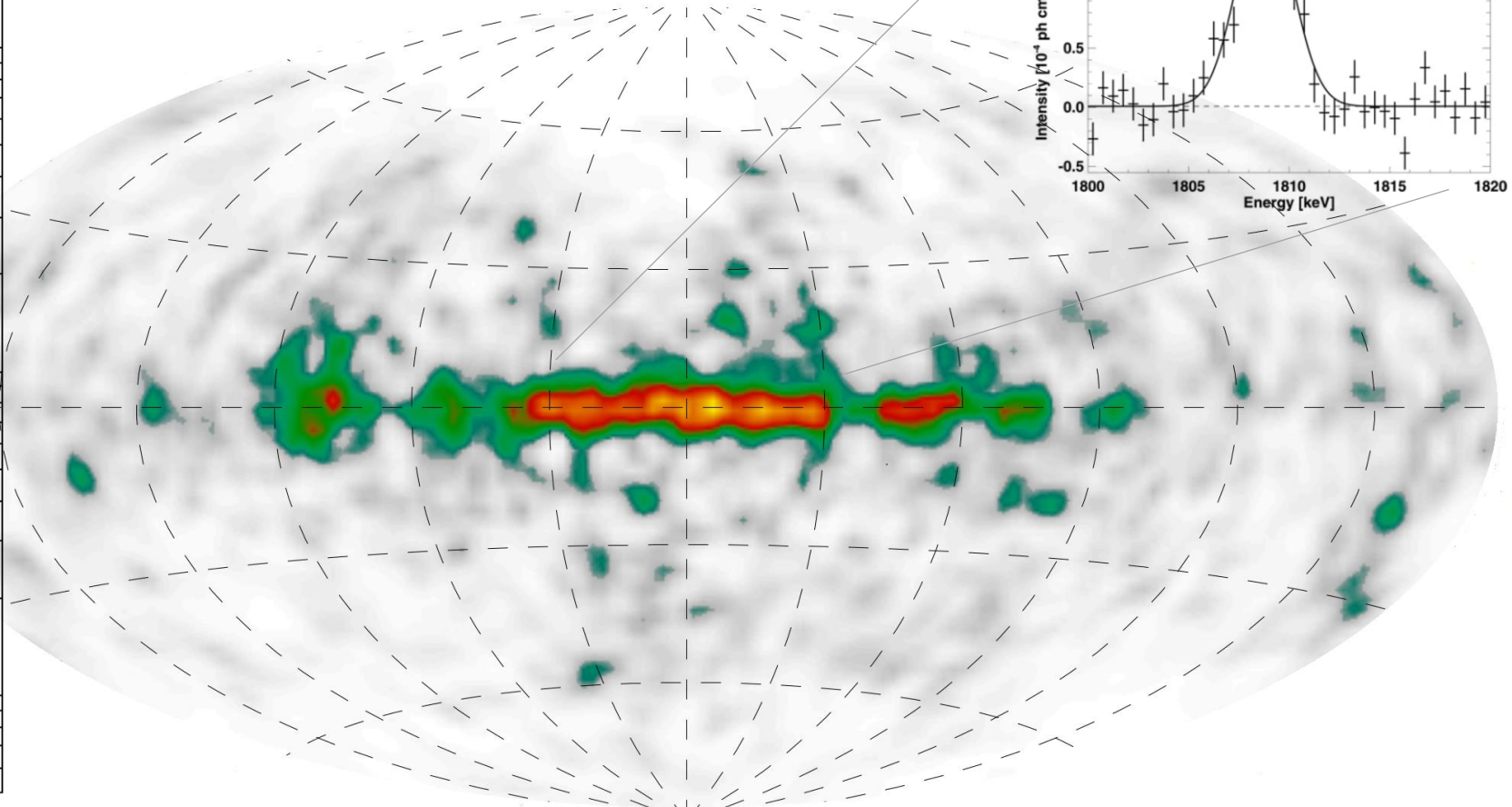
Credit: ESA/Herschel/PACS/MESSE Key Programme Supernova Remnant Team; NASA, ESA and Allison Loll/Jeff Hester (Arizona State University).

Modified from Feige et al., PRL 121, 221103 (2018)

Radioactivity in the Milky Way



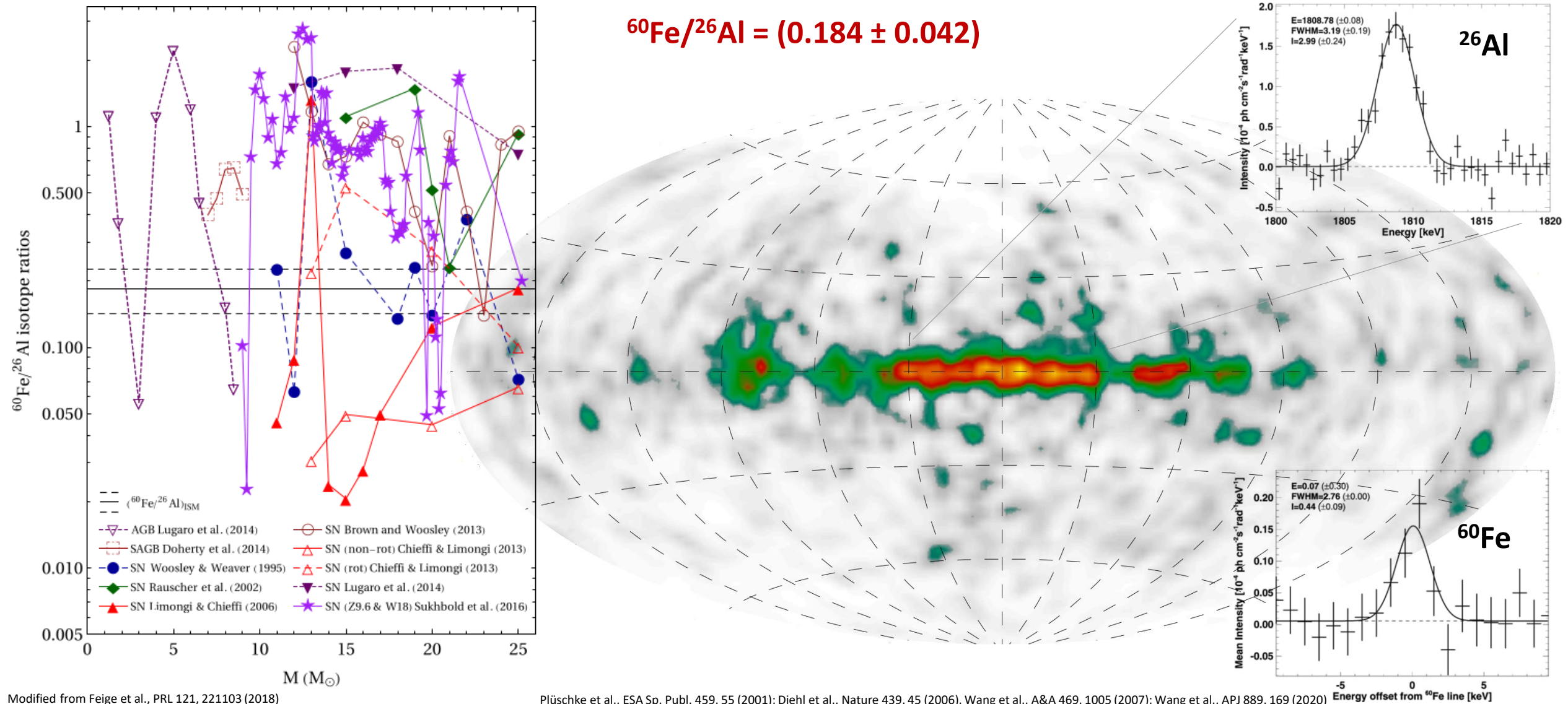
Supernova-rate in our Milky Way
2 supernovae per century



Modified from Feige et al., PRL 121, 221103 (2018)

Plüschke et al., ESA Sp. Publ. 459, 55 (2001); Diehl et al., Nature 439, 45 (2006)

Radioactivity in the Milky Way



Modified from Feige et al., PRL 121, 221103 (2018)

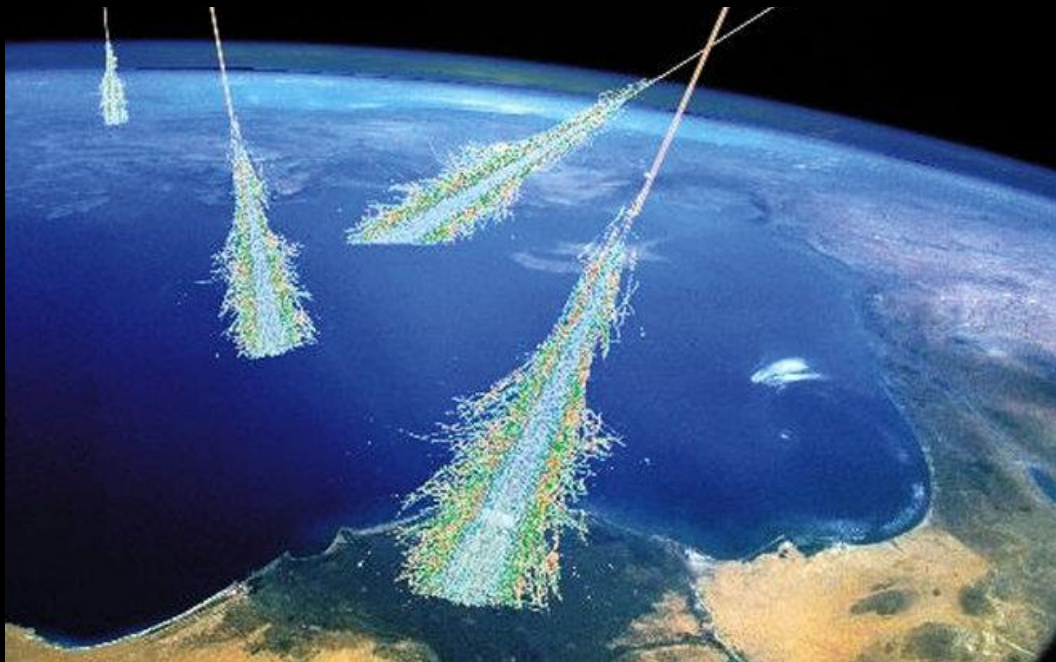
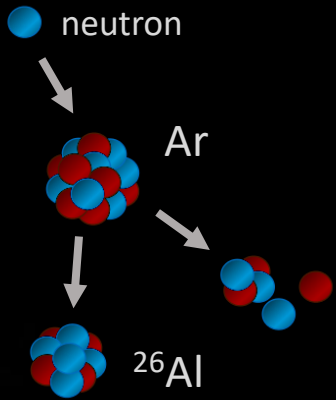
Plüschke et al., ESA Sp. Publ. 459, 55 (2001); Diehl et al., Nature 439, 45 (2006); Wang et al., A&A 469, 1005 (2007); Wang et al., APJ 889, 169 (2020)

Radioactive Supernova Traces

Other ^{60}Fe sources negligible

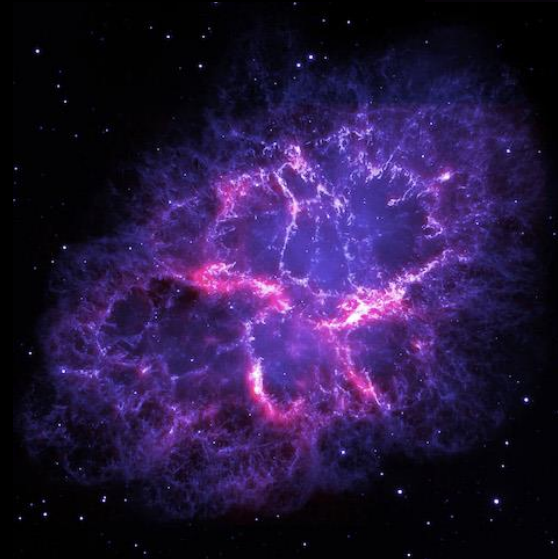
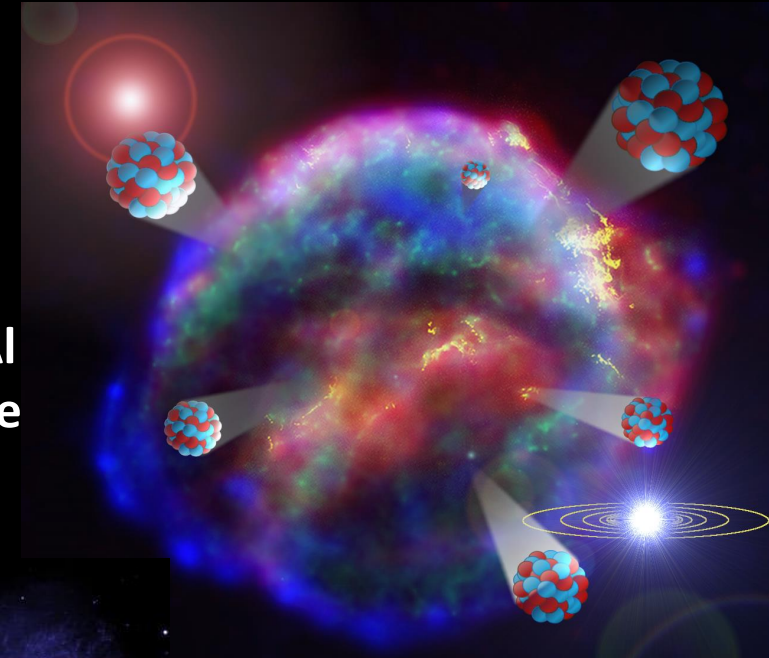
Continuous atmospheric
 ^{26}Al production:

$1250 \text{ ats cm}^{-2} \text{ yr}^{-1}$
(Auer et al., 2009)



Space.com

^{26}Al
 ^{60}Fe



10 – 150 pc



Supernova Traces on Earth

Radiocarbon 38:68 (1996)

THE ASTROPHYSICAL JOURNAL, 470:1227–1236, 1996 October 20

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Supernova Traces on Earth

Alessandro Airo demonstrates where not to search for Supernova traces on Earth.



Supernova Traces on Earth

Deep-sea sediments

- Accumulation rate: **mm/kyr**
- Ideal for resolving SN events!!!



Deep-sea crusts and nodules

- Growth rate: **mm/Myr**
 - Rough time estimation
- Cover a large time range



Undisturbed archives
with a long-term memory!

Take-Home Message I

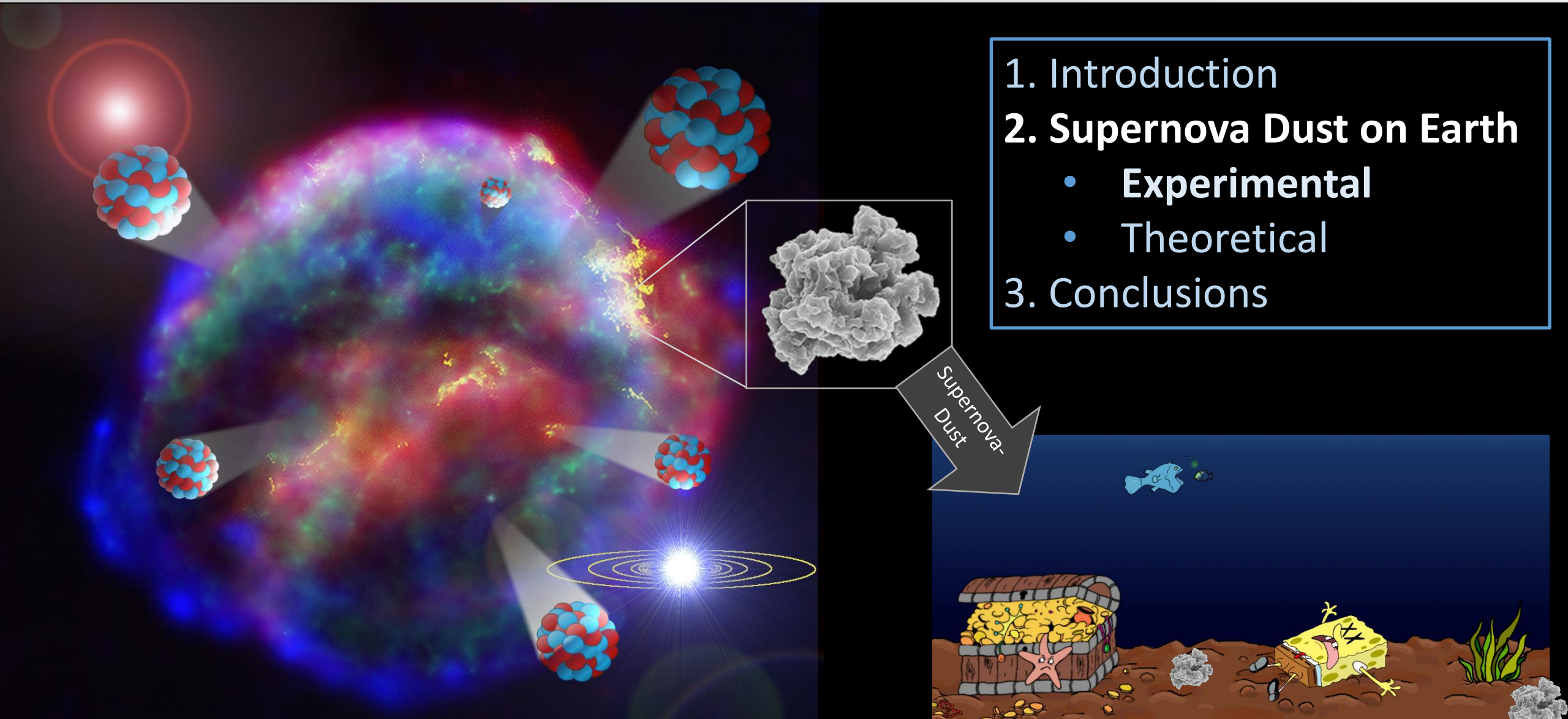
Million-year-old remnants of Supernova can leave traces on Earth if close enough

- Our Sun is located in a region with turbulent history: **The Local Bubble!**
- Suitable geological archives: **Deep-sea records**
- Suitable isotopes: **Long-lived radionuclides**
 - No leftover from the Solar System formation period
 - Negligible terrestrial production
 - **^{60}Fe and ^{26}Al** both have been observed within our Milky Way:
Continuous and fresh input from stellar explosions

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Outline

1. Introduction
2. **Supernova Dust on Earth**
 - Experimental
 - Theoretical
3. Conclusions



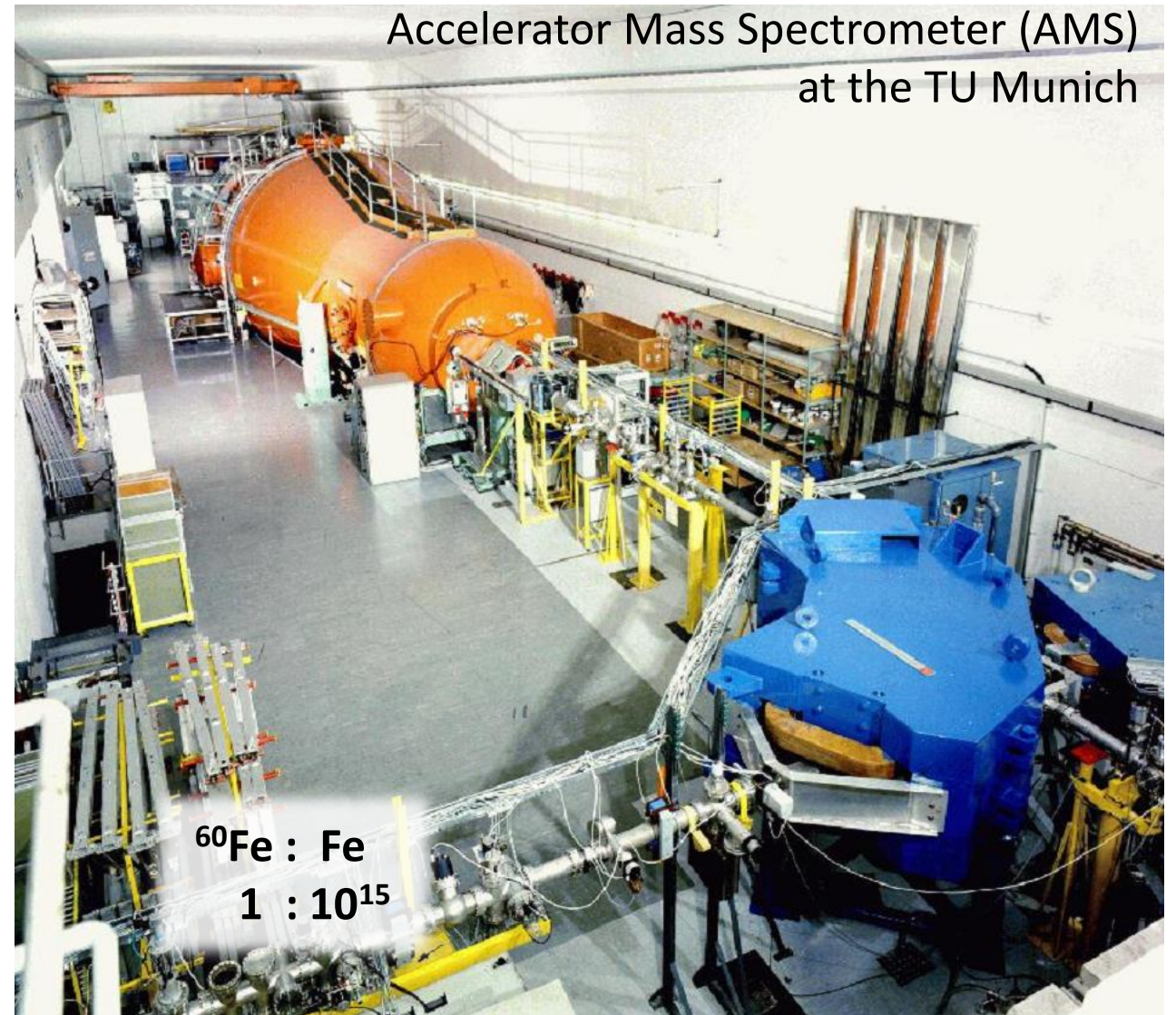
Supernova Traces in a Deep-Sea Crust



Supernova Traces in a Deep-Sea Crust



Supernova Traces in a Deep-Sea Crust



www.mll-muenchen.de/tandem/index.html

Supernova Traces in a Deep-Sea Crust



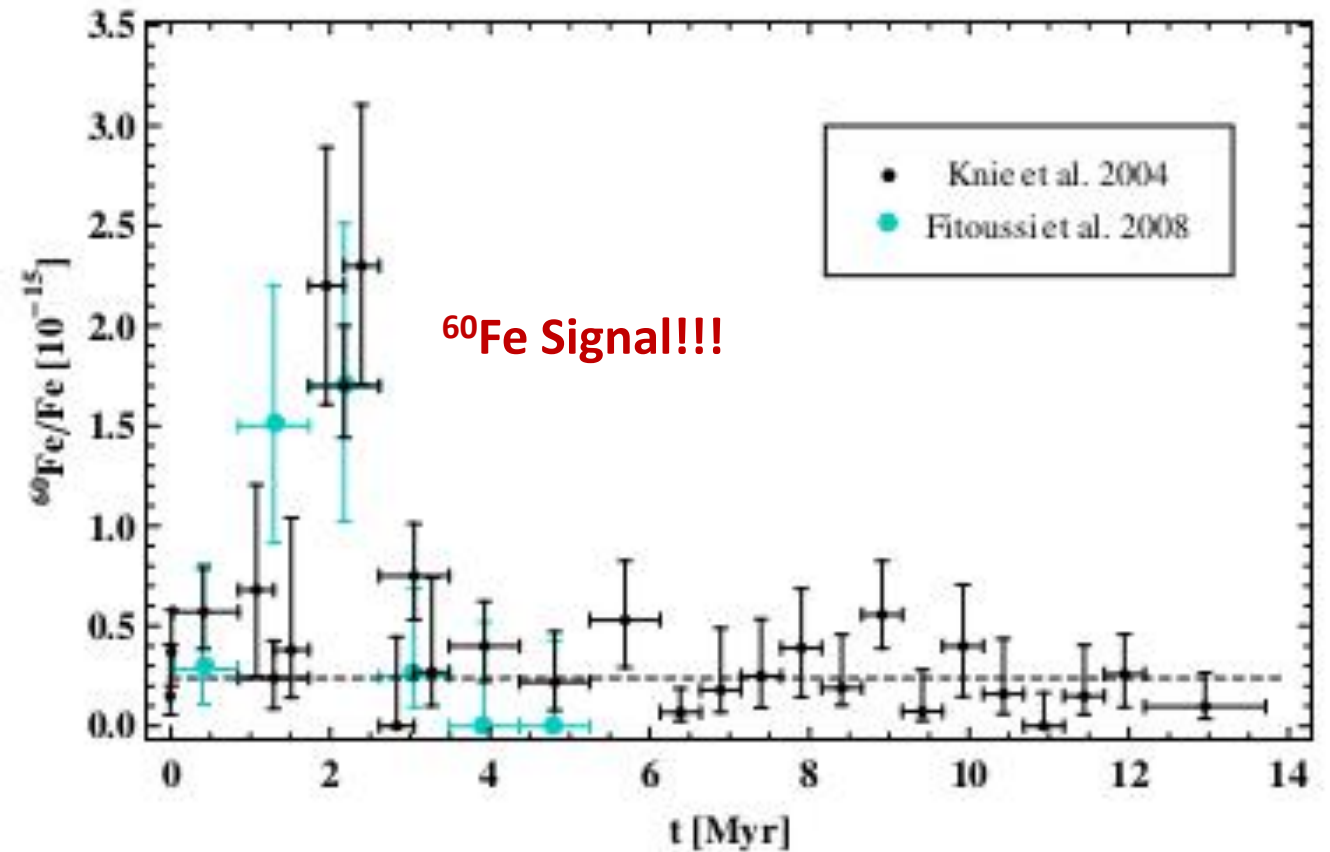
VOLUME 93, NUMBER 17

PHYSICAL REVIEW LETTERS

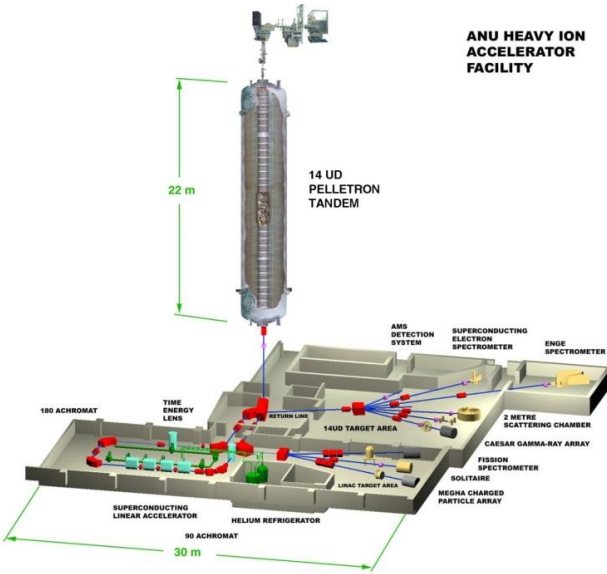
week ending
22 OCTOBER 2004

^{60}Fe Anomaly in a Deep-Sea Manganese Crust and Implications for a Nearby Supernova Source

K. Knie,¹ G. Korschinek,^{1,*} T. Faestermann,¹ E. A. Dorfi,² G. Rugel,^{1,3} and A. Wallner^{1,3}



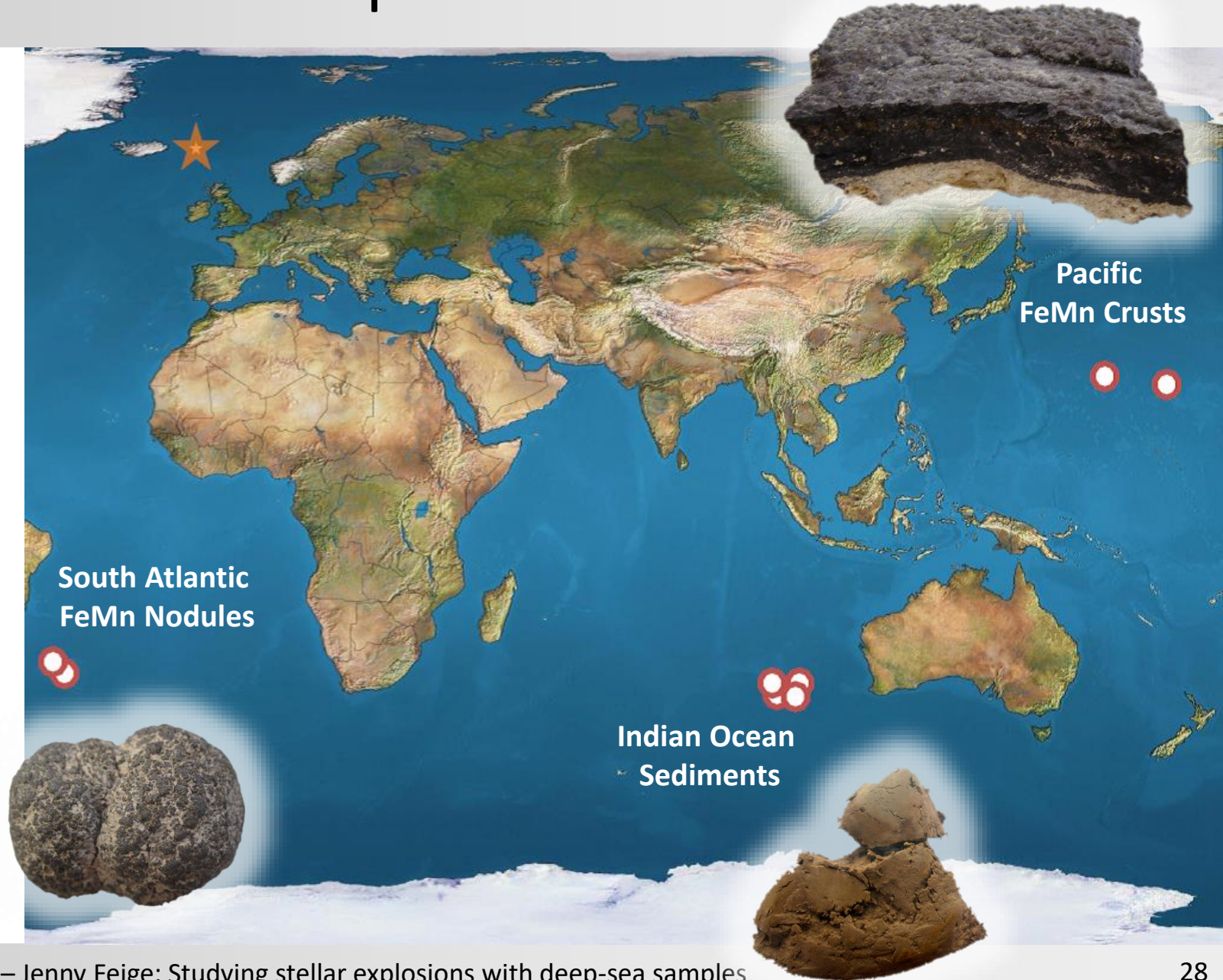
Searching for Global Supernova Traces



ANU HEAVY ION
ACCELERATOR
FACILITY

ANU
Canberra
Australia

Less than
 10^{60} Fe atoms
in a 3 g sediment
sample!



Pacific
FeMn Crusts

South Atlantic
FeMn Nodules

Indian Ocean
Sediments

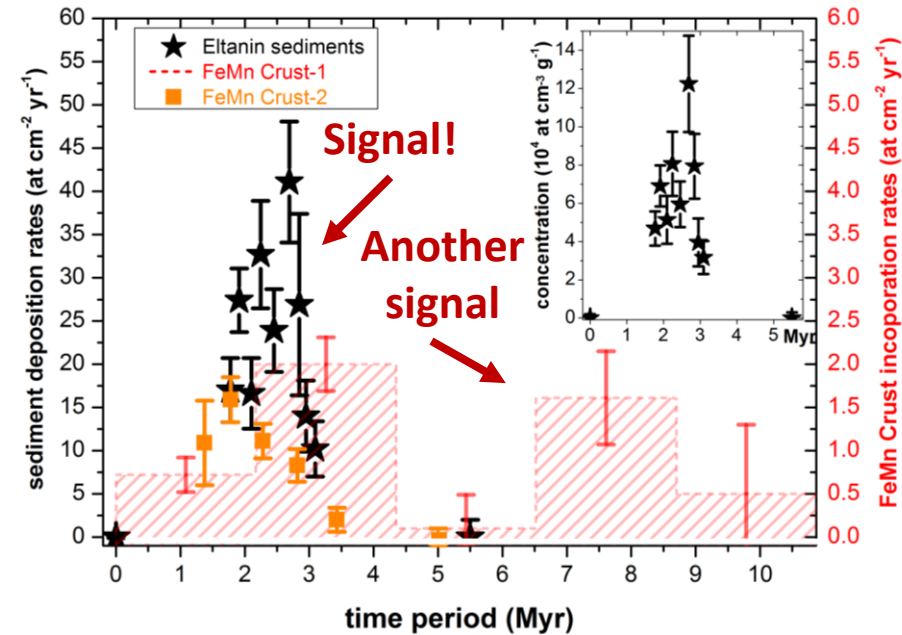
Global Supernova Traces

LETTER

doi:10.1038/nature17196

Recent near-Earth supernovae probed by global deposition of interstellar radioactive ^{60}Fe

A. Wallner¹, J. Feige^{2*}, N. Kinoshita³, M. Paul⁴, L. K. Fifield¹, R. Golser², M. Honda⁵, U. Linnemann⁶, H. Matsuzaki⁷, S. Merchel⁸, G. Rugel⁸, S. G. Tims¹, P. Steier², T. Yamagata⁹ & S. R. Winkler²



Peak-flux: About 30 ats/yr/cm^2

Absolute amount of ^{60}Fe
distributed on Earth: ~20 kg

Imagine **20 kg**
of flour spread
over Earth. Can
you detect it?

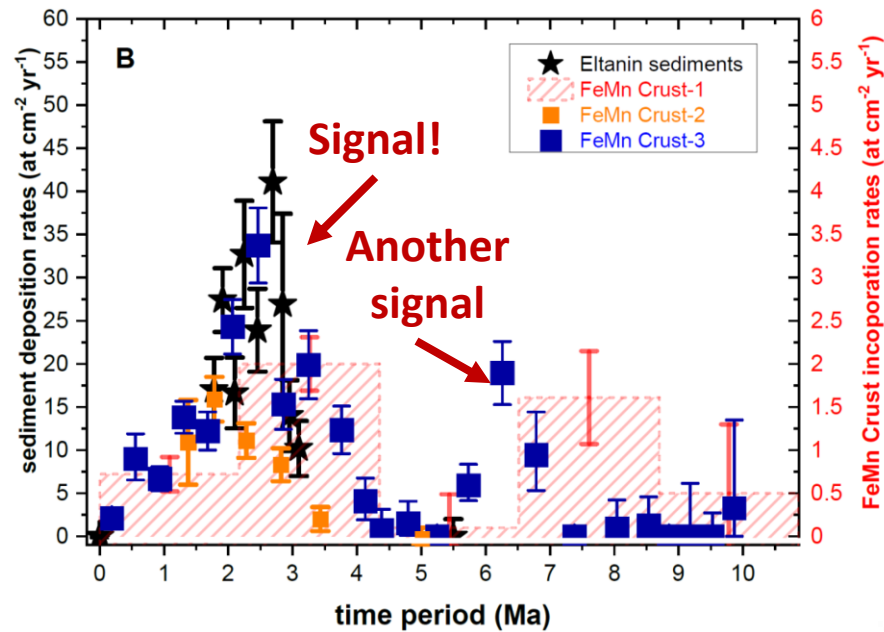


Global Supernova Traces

NUCLEAR ASTROPHYSICS Wallner *et al.*, *Science* **372**, 742–745 (2021)

^{60}Fe and ^{244}Pu deposited on Earth constrain the r-process yields of recent nearby supernovae

A. Wallner^{1,2*}, M. B. Froehlich¹, M. A. C. Hotchkis³, N. Kinoshita⁴, M. Paul⁵, M. Martschini^{1†}, S. Pavetich¹, S. G. Tims¹, N. Kivel⁶, D. Schumann⁶, M. Honda^{7‡}, H. Matsuzaki⁸, T. Yamagata⁸



The 2nd signal is real and has been confirmed in another crust!

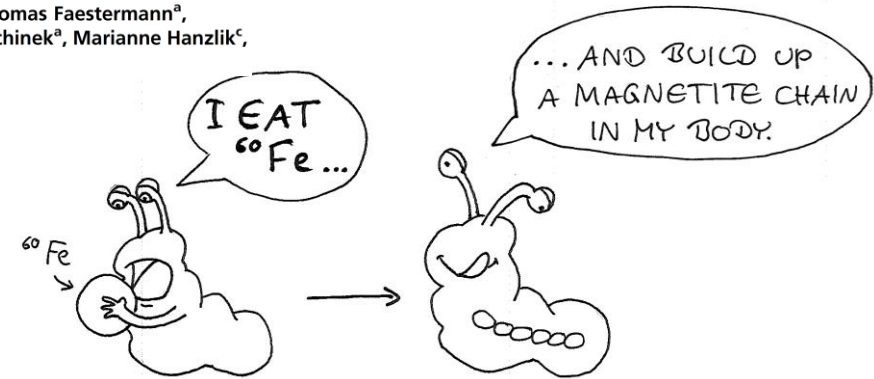
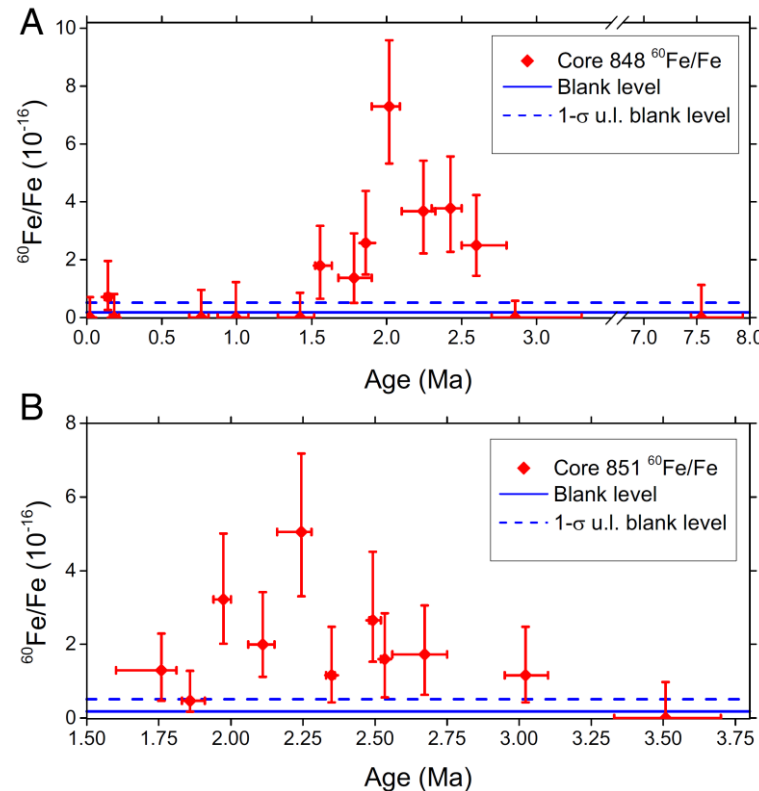


Supernova Traces in Fossil Bacteria



Time-resolved 2-million-year-old supernova activity discovered in Earth's microfossil record

Peter Ludwig^a, Shawn Bishop^{a,1}, Ramon Egli^b, Valentyna Chernenko^a, Boyana Deneva^a, Thomas Faestermann^a, Nicolai Famulok^a, Leticia Fimiani^a, José Manuel Gómez-Guzmán^a, Karin Hain^a, Gunther Korschinek^a, Marianne Hanzlik^c, Silke Merchel^d, and Georg Rugel^d



Two sediment cores from equatorial pacific

- Magnetotactic bacteria
- Chemical extraction of magnetofossils
- Time-resolved ^{60}Fe signal in biogenic reservoir

Supernova Traces on the Moon

PRL **116**, 151104 (2016)

PHYSICAL REVIEW LETTERS

week ending
15 APRIL 2016

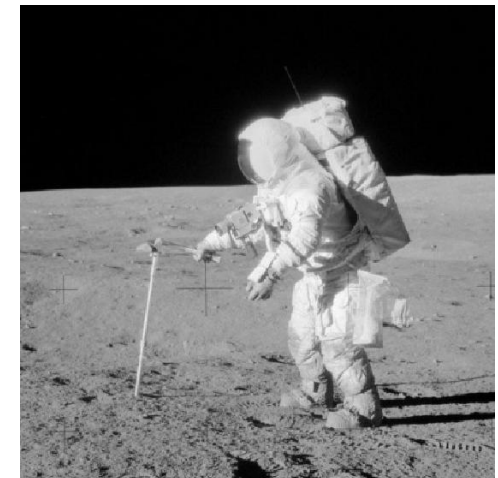
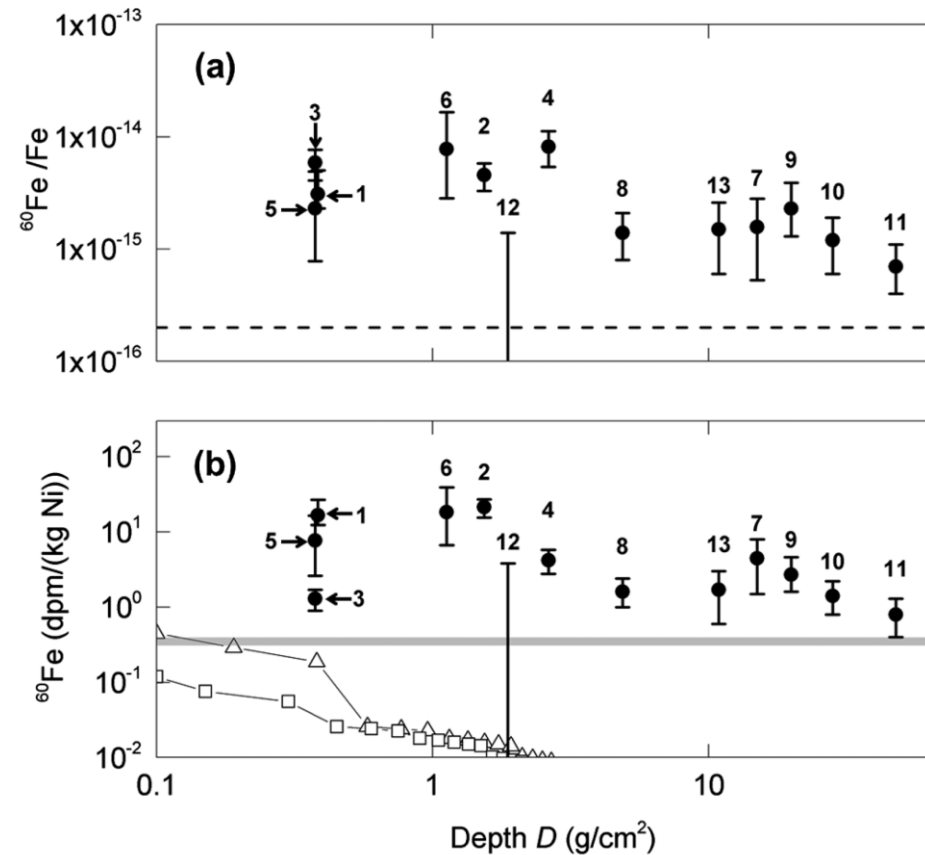
Interstellar ^{60}Fe on the Surface of the Moon

L. Fimiani,¹ D. L. Cook,^{2,*} T. Faestermann,¹ J. M. Gómez-Guzmán,¹ K. Hain,¹ G. Herzog,² K. Knie,^{1,†}
G. Korschinek,^{1,‡} P. Ludwig,¹ J. Park,² R. C. Reedy,³ and G. Rugel^{1,§}

- No sedimentation
- Gardening effect
- Cosmogenic ^{60}Fe

Everything points to one or more supernovae in the past Myr

- ^{60}Fe is everywhere it shouldn't be!
- We are living in the Local Bubble!

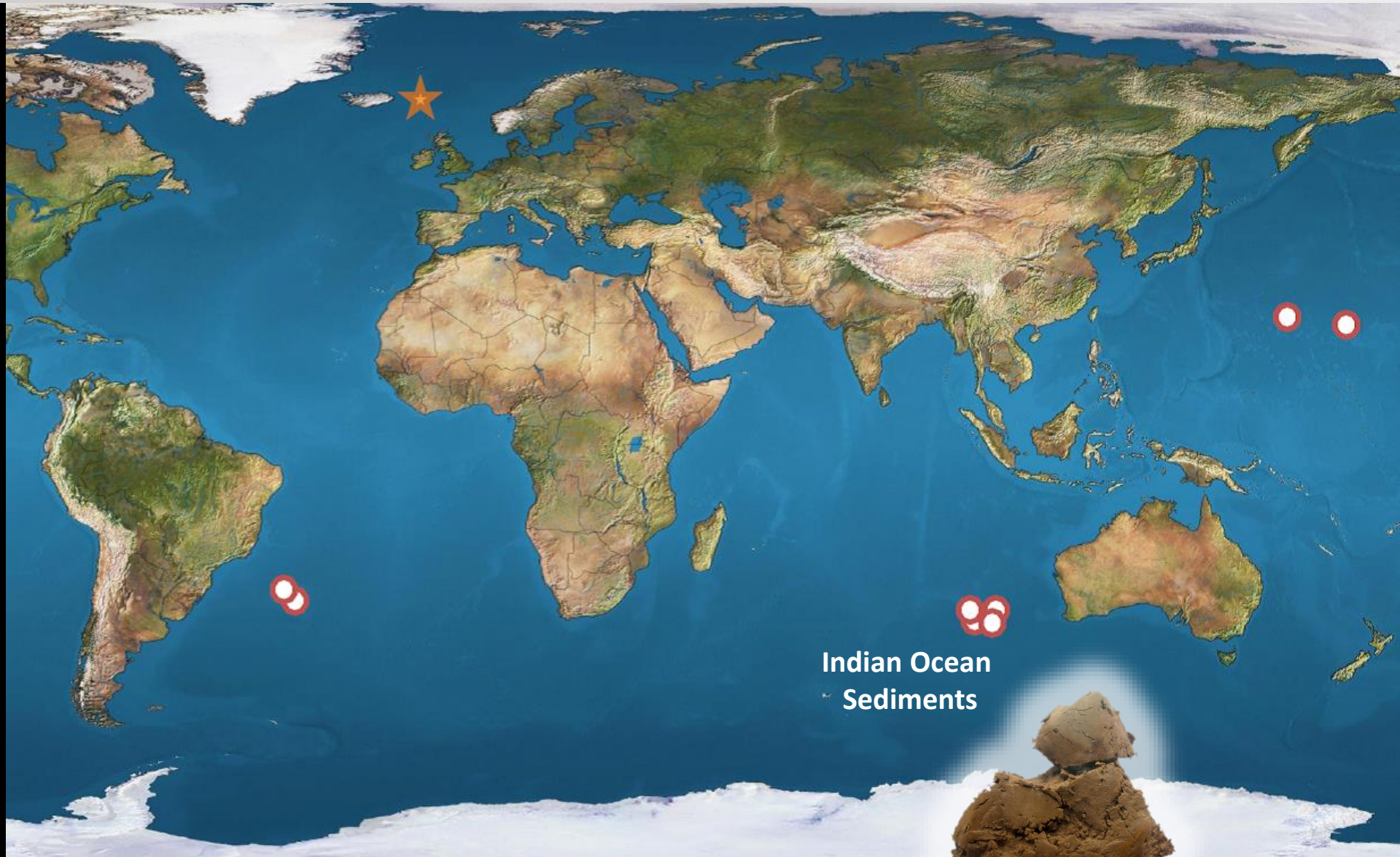
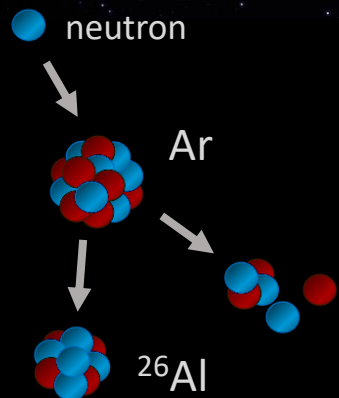
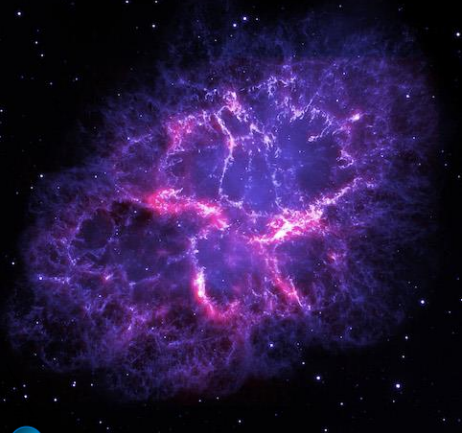


Alan Bean takes **Lunar samples**
(Apollo 12, 1969)

^{26}Al in deep-sea sediments

Sources of ^{26}Al

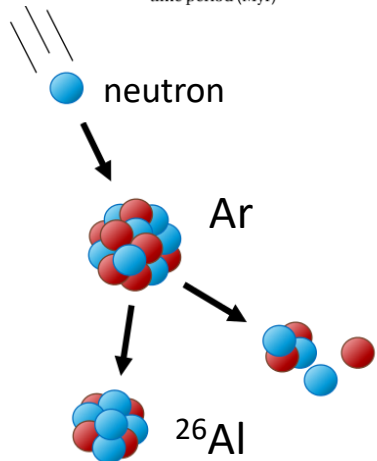
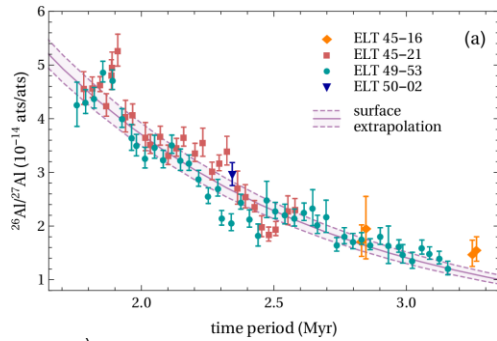
- Supernova
- Atmospheric



^{26}Al in deep-sea sediments

Sources of ^{26}Al

- Supernova
- Atmospheric

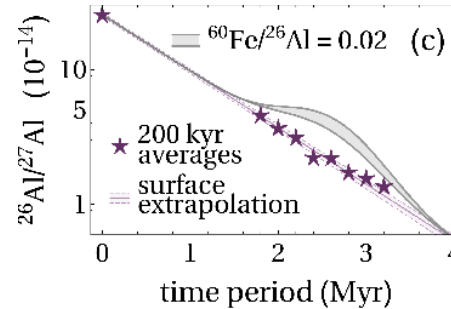
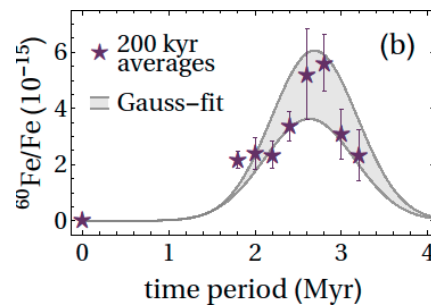


PHYSICAL REVIEW LETTERS **121**, 221103 (2018)

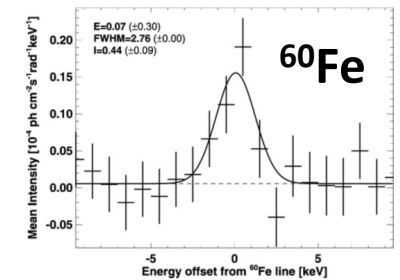
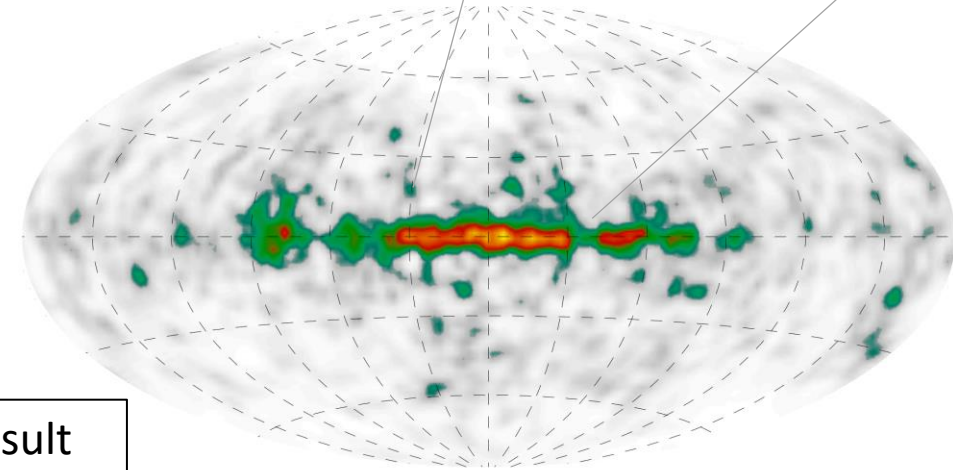
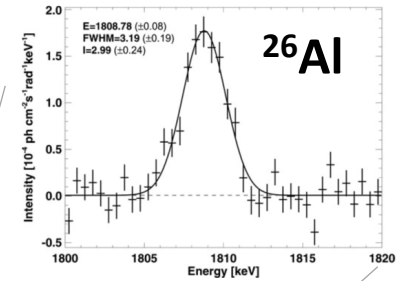
Editors' Suggestion

Limits on Supernova-Associated $^{60}\text{Fe}/^{26}\text{Al}$ Nucleosynthesis Ratios from Accelerator Mass Spectrometry Measurements of Deep-Sea Sediments

Jenny Feige,^{1,2,*} Anton Wallner,³ Randolph Altmeyer,⁴ L. Keith Fifield,³ Robin Golser,² Silke Merchel,⁵ Georg Rugel,⁵ Peter Steier,² Stephen G. Tims,³ and Stephan R. Winkler^{2,6}



- A low $^{60}\text{Fe}/^{26}\text{Al}$ nucleosynthesis ratio should result in a visible ^{26}Al signature.
- The measured data suggests a limit of $0.18^{+0.15}_{-0.08}$.
- Comparable to observations (0.184 ± 0.042).

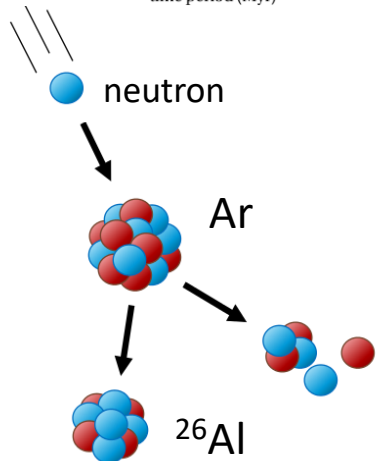
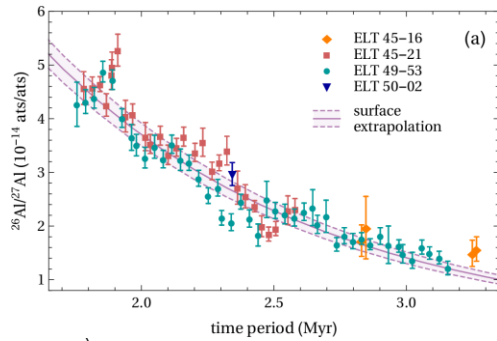


Plüschke et al., ESA Sp. Publ. 459, 55 (2001); Diehl et al., Nature 439, 45 (2006); Wang et al., A&A 469, 1005 (2007); Wang et al., APJ 889, 169 (2020)

^{26}Al in deep-sea sediments

Sources of ^{26}Al

- Supernova
- Atmospheric

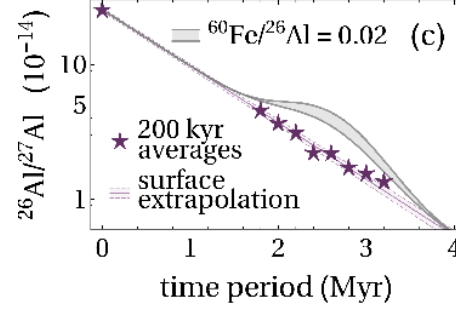
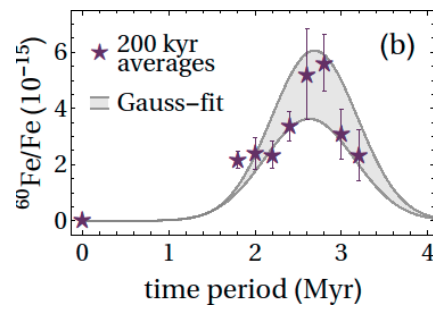


PHYSICAL REVIEW LETTERS **121**, 221103 (2018)

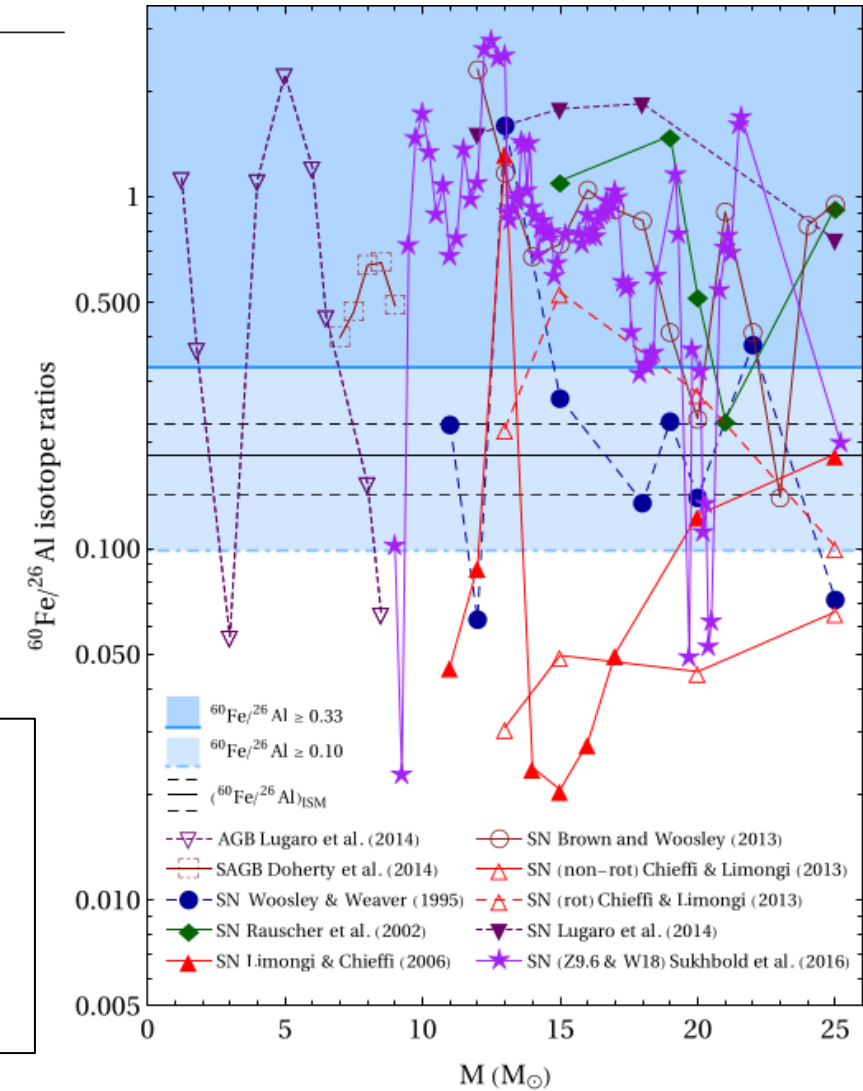
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- Comparable to observations (0.184 ± 0.042).
- Comparable to stellar nucleosynthesis models.



Recent Supernova Influx

PHYSICAL REVIEW LETTERS 123, 072701 (2019)

Editors' Suggestion

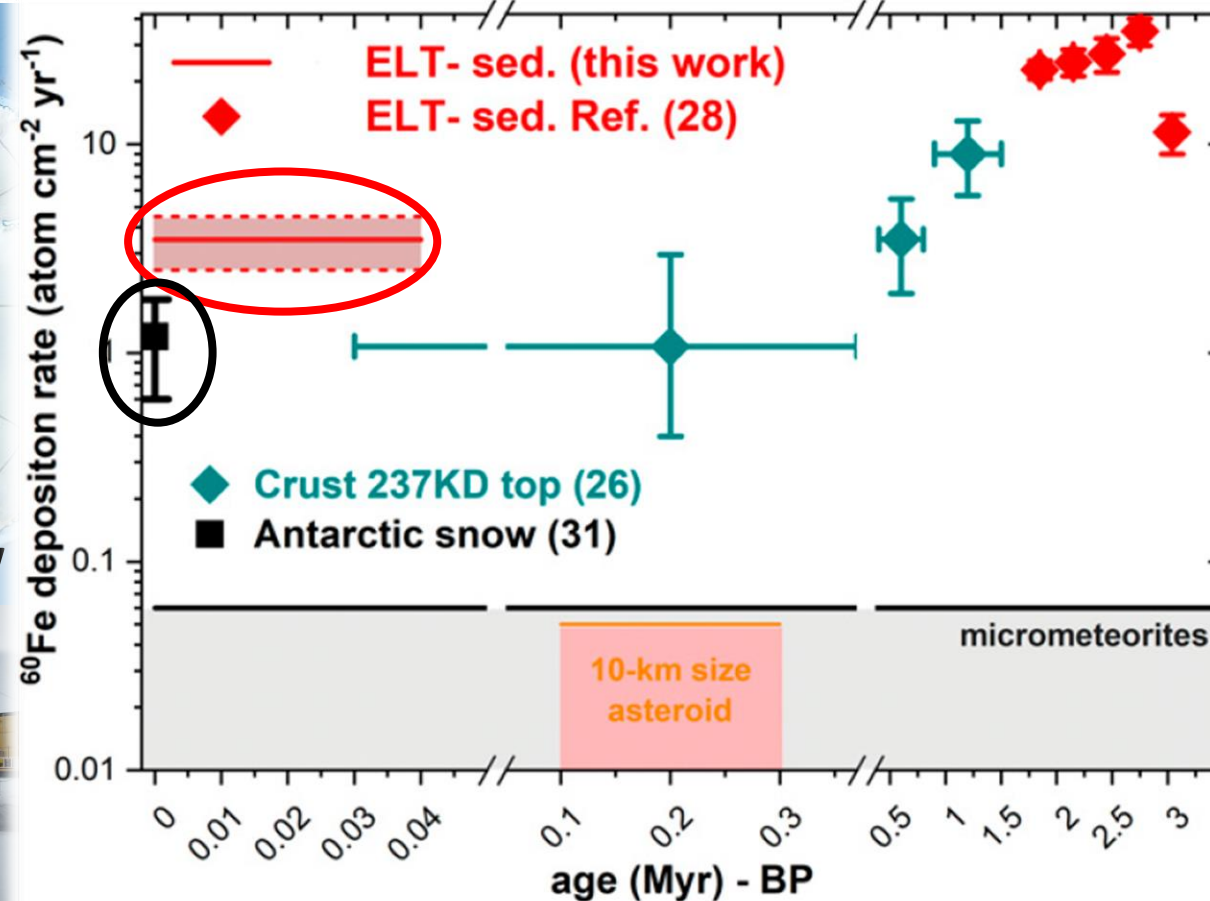
Featured in Physics

Interstellar ^{60}Fe in Antarctica

Dominik Koll,^{1,*} Gunther Korschinek,^{1,2} Thomas Faestermann,^{1,2} J. M. Gómez-Guzmán,¹ Sepp Kipfstuhl,³ Silke Merchel,⁴ and Jan M. Welch⁵

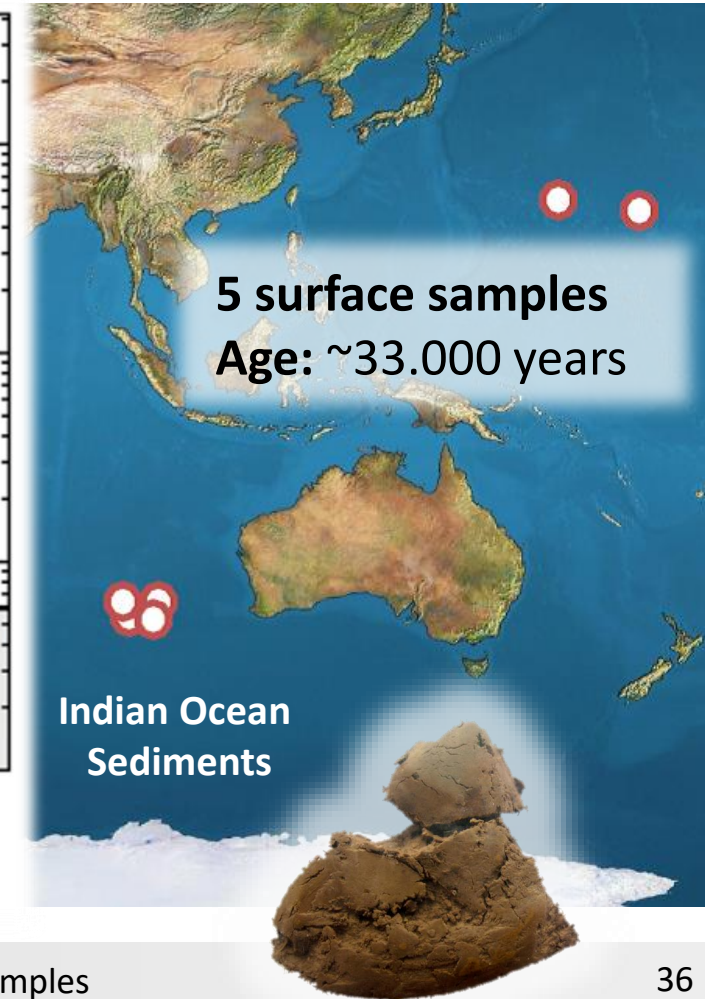


^{60}Fe Flux today
1/10 of peak flux



^{60}Fe deposition during the late Pleistocene and the Holocene echoes past supernova activity

A. Wallner^{a,b,1}, J. Feige^{c,d}, L. K. Fifield^a, M. B. Froehlich^a, R. Golser^c, M. A. C. Hotchkis^a, D. Koll^a, G. Leckenby^a, M. Martschini^{a,c}, S. Merchel^b, S. Panjkov^a, S. Pavetich^a, G. Rugel^b, and S. G. Tims^a



Take-Home Message II

Million-year-old Supernova traces have been detected in

- Deep-sea crusts, nodules and sediments all over the world
- Fossil magnetotactic bacteria
- Lunar samples

Two distinct ^{60}Fe signals around 2-3 and 6-7 Myr ago

- SN-derived ^{26}Al is overwhelmed by atmospheric background
 - Lower $^{60}\text{Fe}/^{26}\text{Al}$ limit of ~ 0.18

Still today ^{60}Fe is reaching our Earth

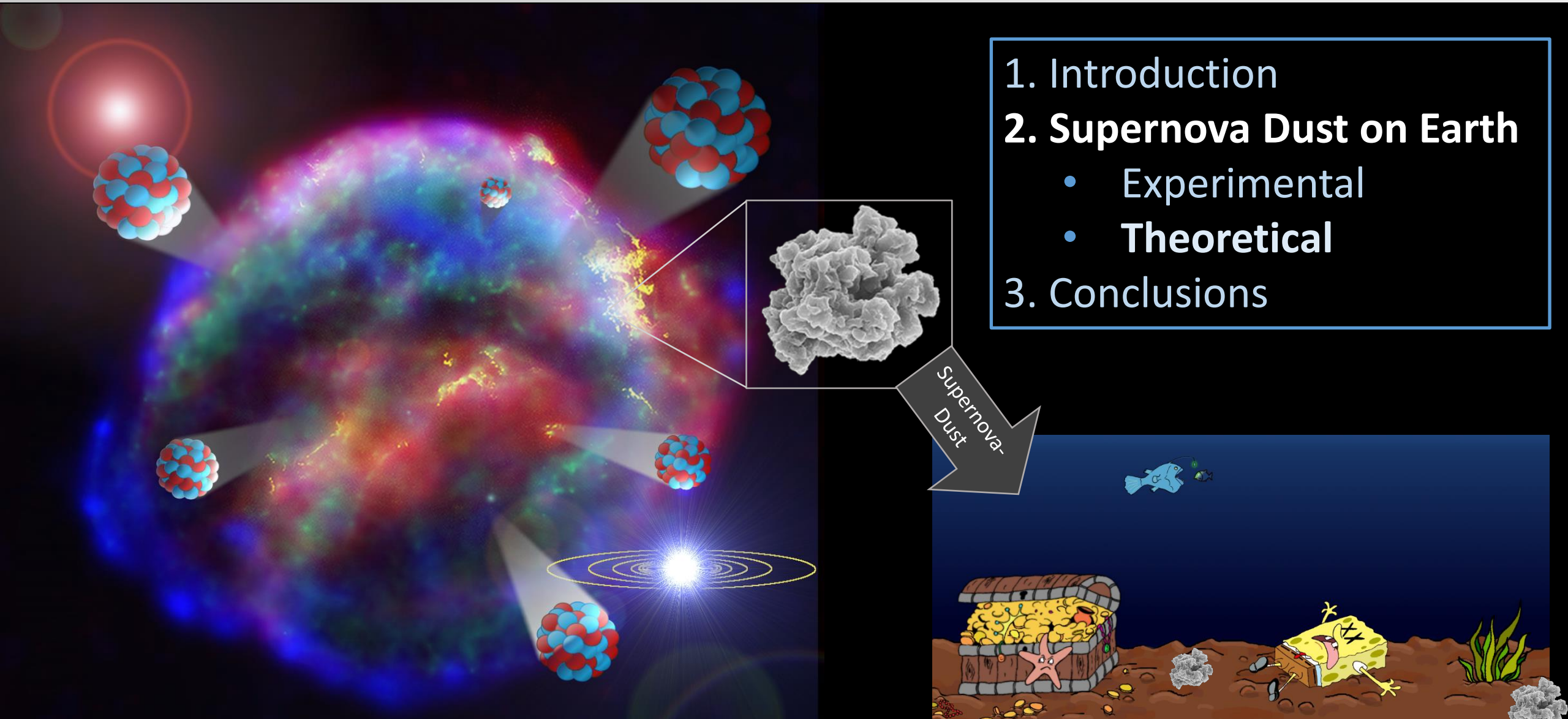
Further findings

- ^{53}Mn in deep-sea crusts (Korschinek et al., PRL 125, 031101, 2020)
- ^{244}Pu in deep-sea crusts (Wallner et al., Science 372, 742, 2021; D. Koll's talk)

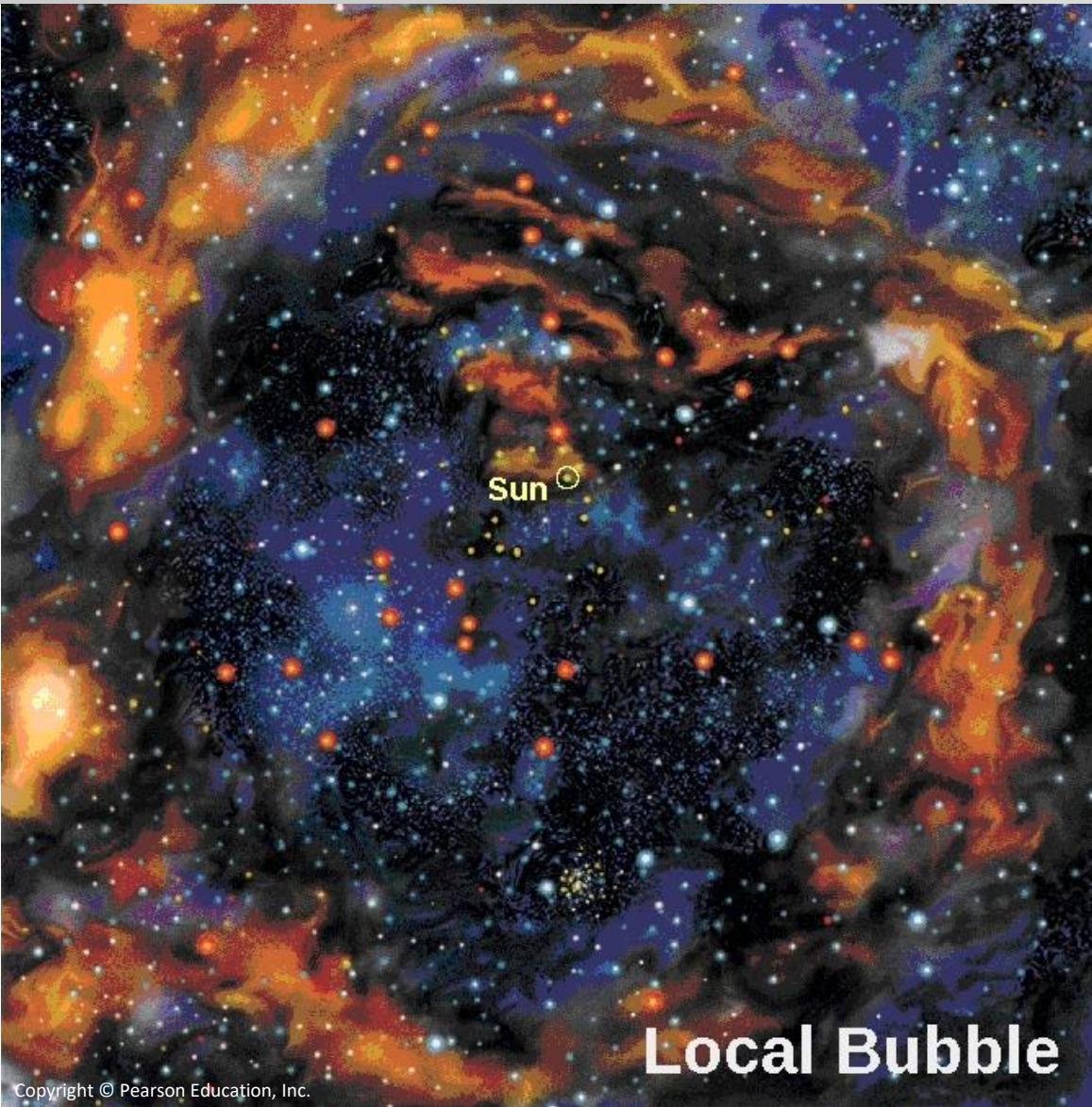
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Outline

1. Introduction
- 2. Supernova Dust on Earth**
 - Experimental
 - **Theoretical**
3. Conclusions

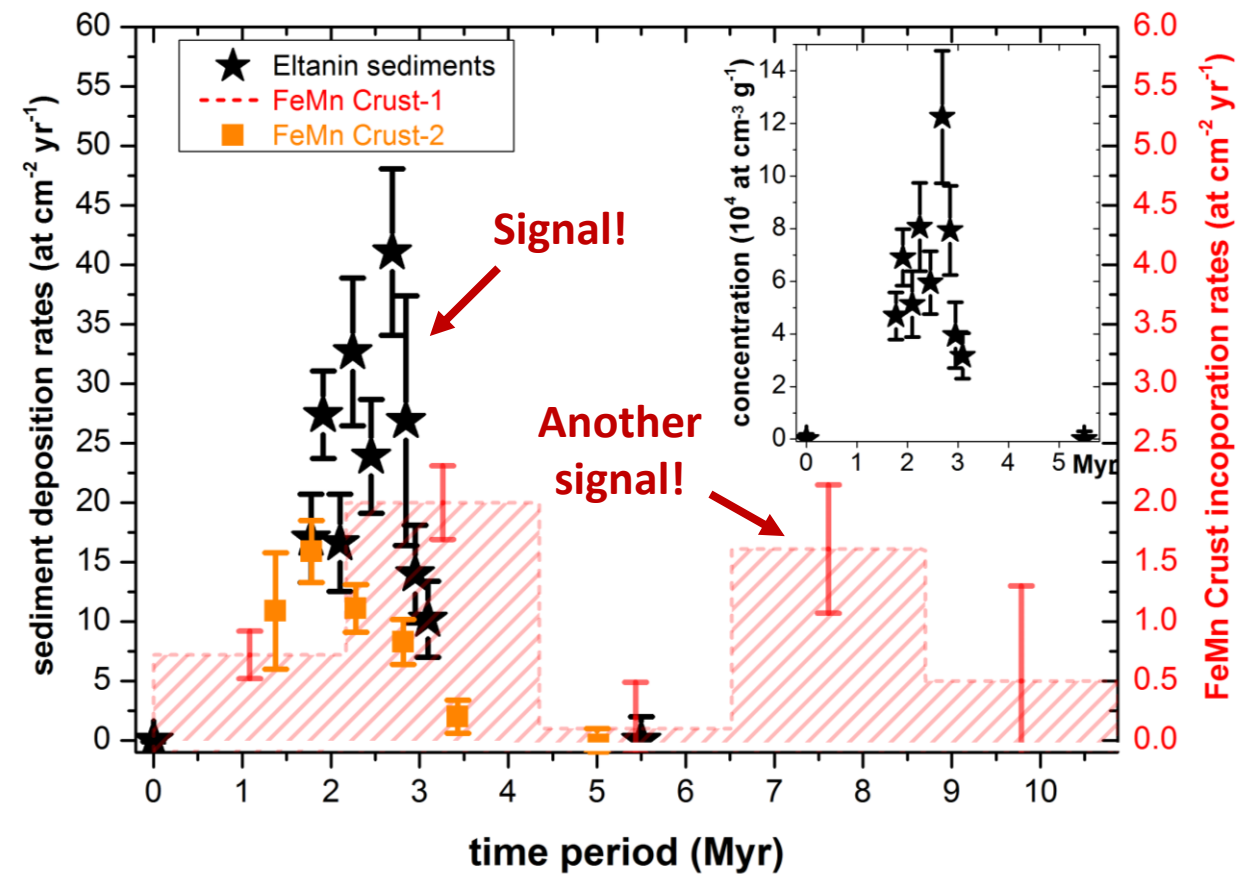


Supernova Signatures



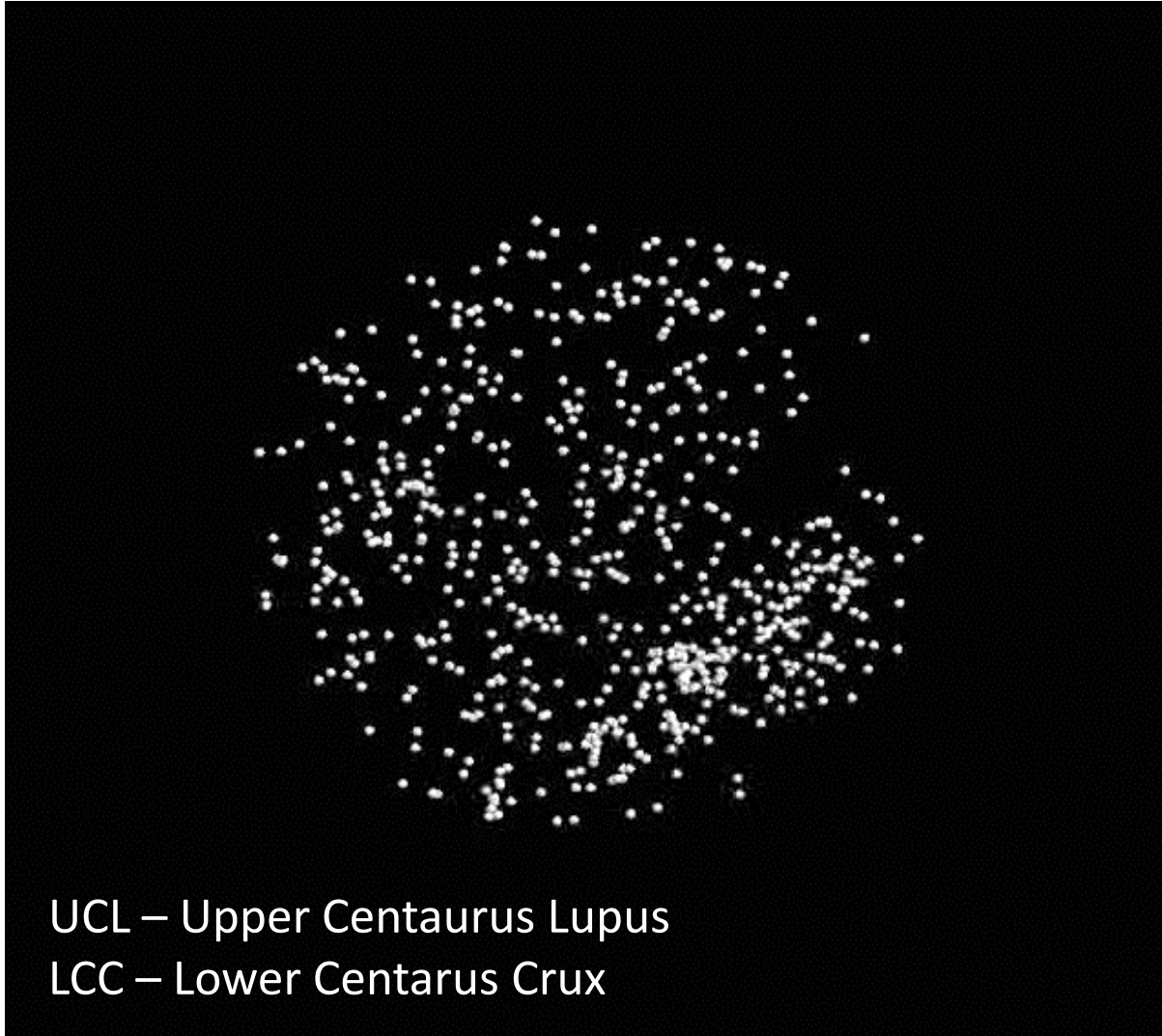
Which stars created the Local Bubble and the ^{60}Fe signals?

Why is ^{60}Fe still arriving on Earth today?



Wallner, Feige et al., Nature 532, 69 (2016)

Nearby Stellar Moving Groups



Mon. Not. R. Astron. Soc. **373**, 993–1003 (2006)

doi:10.1111/j.1365-2966.2006.11044.x

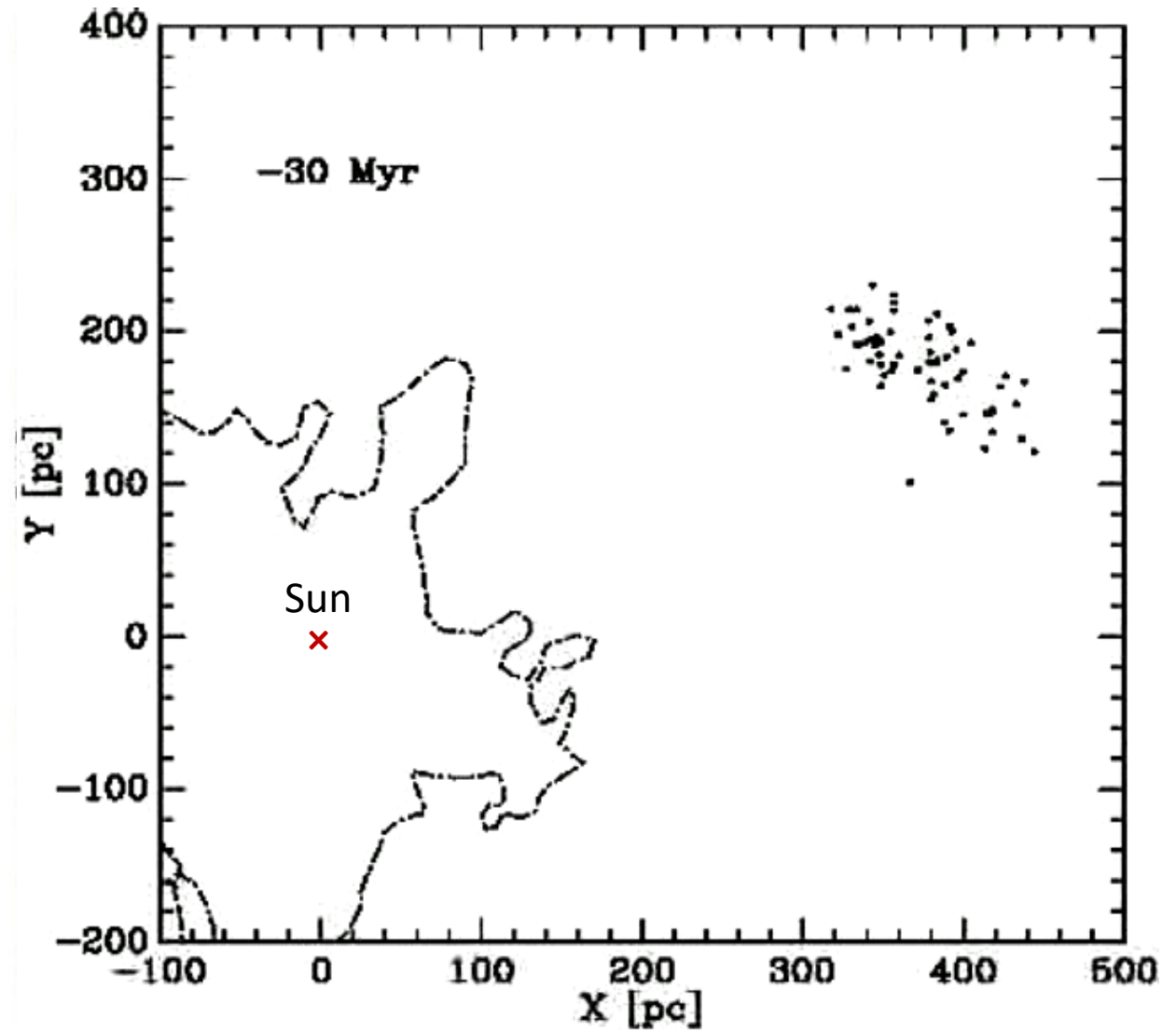
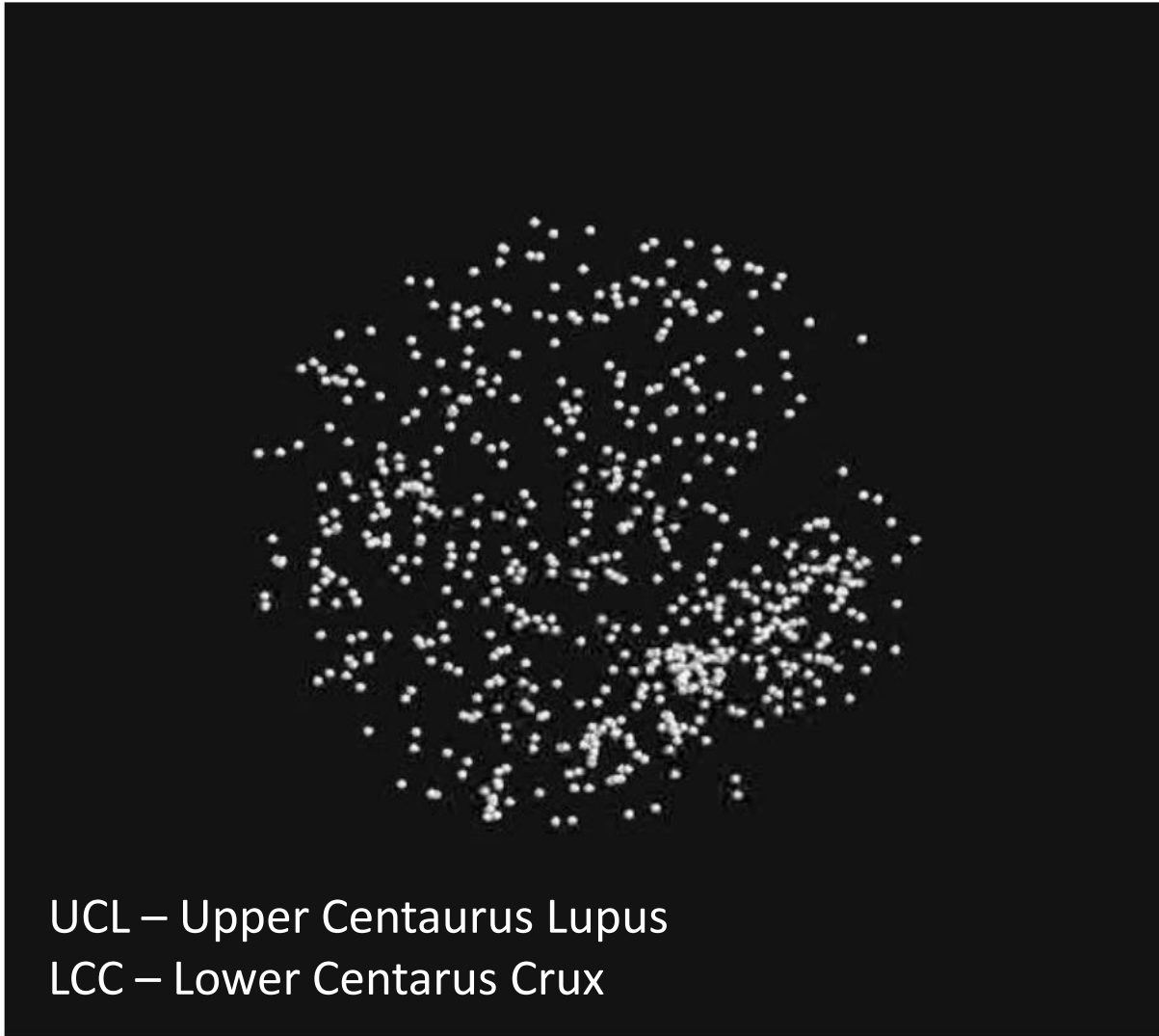
The search for the origin of the Local Bubble redivivus

B. Fuchs,^{1*} D. Breitschwerdt,² M. A. de Avillez,^{2,3} C. Dettbarn¹ and C. Flynn⁴

Solar environment ($r = 200$ pc) was searched for **suspicious stars**.

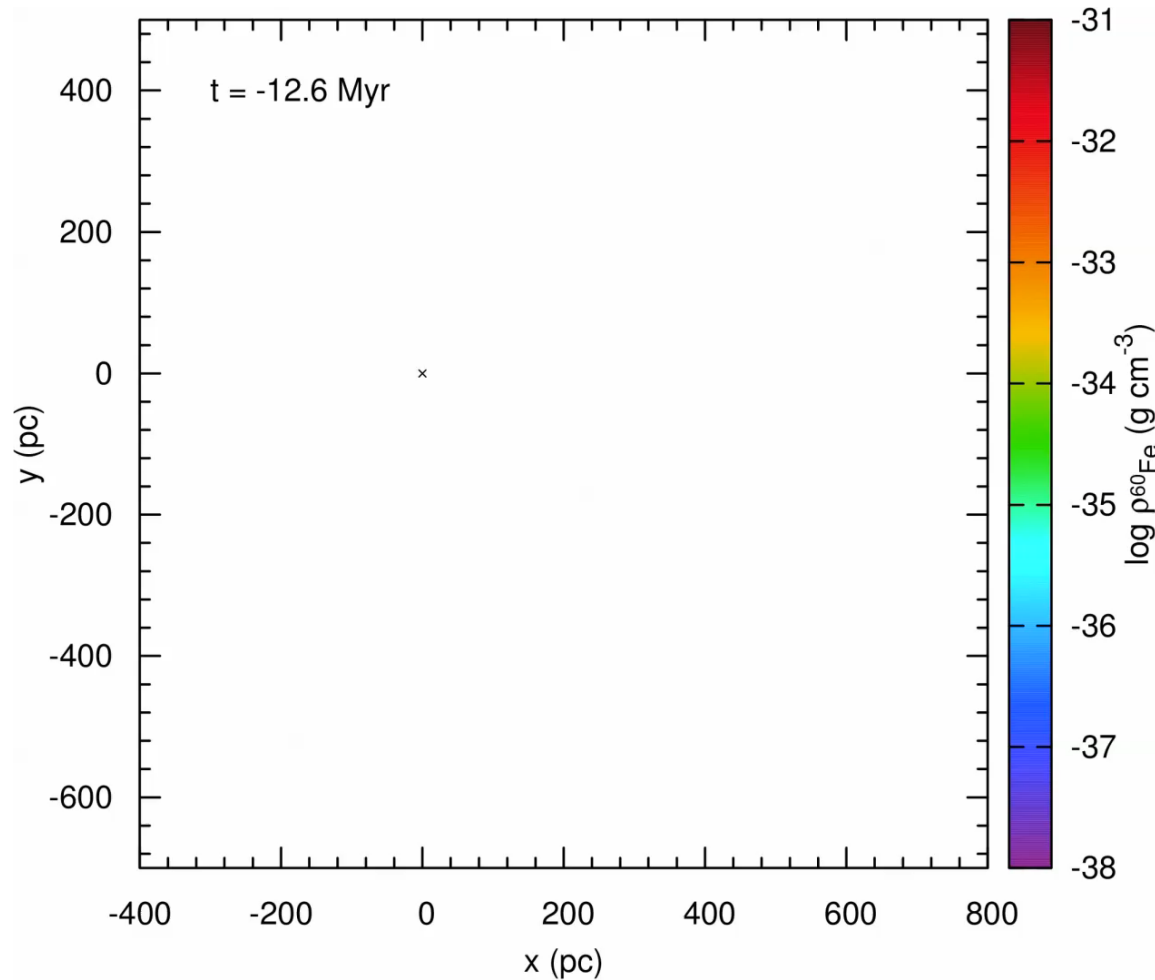
Subgroups of the **Scorpius-Centaurus association!**
Ages: ~ 20 Myr

Nearby Stellar Moving Groups



Fuchs et al., MNRAS 373, 993 (2006)

Modelling the ^{60}Fe Signal and the Local Bubble



Homogeneous ambient medium ($0.3 \text{ atoms cm}^{-3}$)

Resolution: 0.7 pc

Number of Supernovae: 16

Local Bubble shell crosses Solar System 2.2 Myr ago!

Fluence:

- Number of atoms that reach the Earth per cm^2

$$F = \frac{U}{4} \frac{M_{\text{ej}}}{4\pi r^2 A m_p} e^{-t/\tau}$$

t time since explosion

M_{ej} ... ejected ^{60}Fe mass

r distance of the explosion site

A mass number

m_p proton mass

τ ^{60}Fe mean lifetime

U uptake factor

Modelling the ^{60}Fe Signal and the Local Bubble

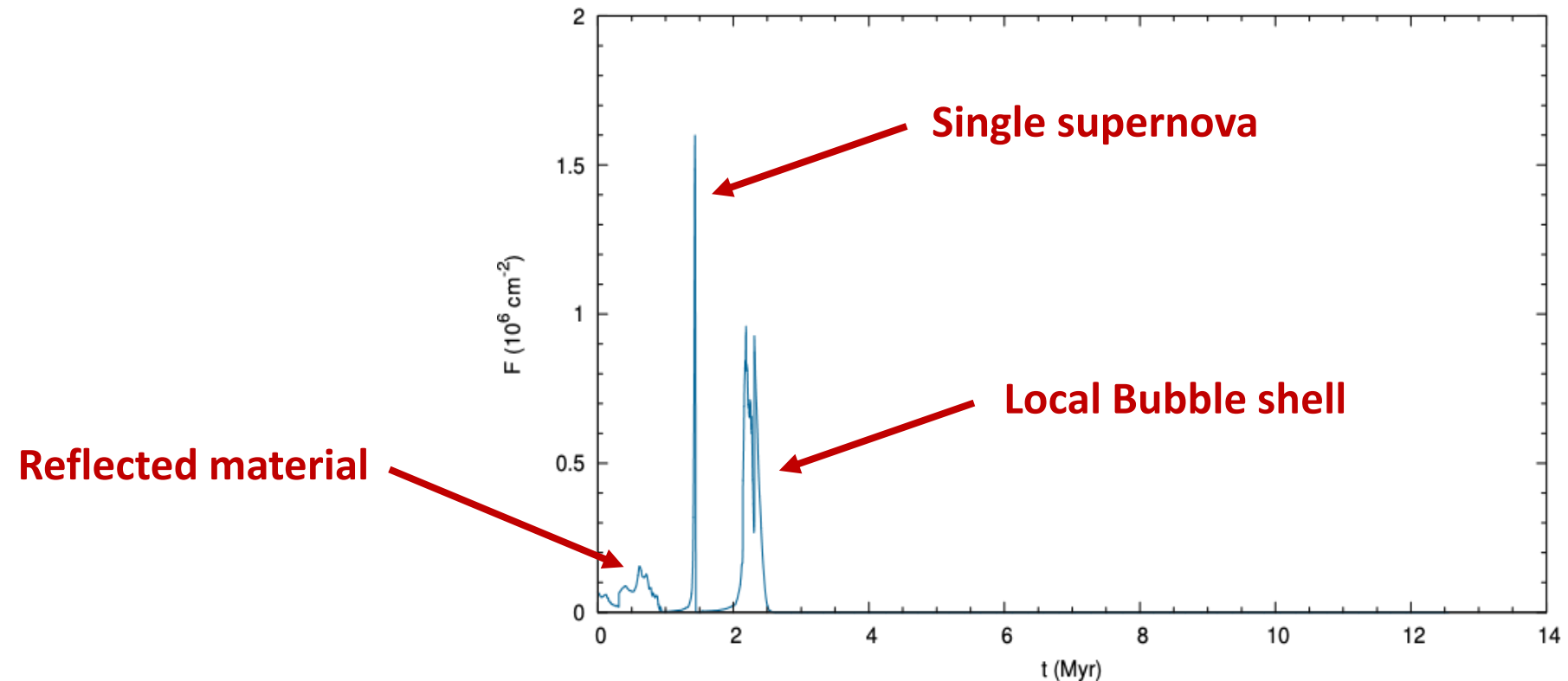
Astronomy
&
Astrophysics

A&A 604, A81 (2017)
DOI: [10.1051/0004-6361/201629837](https://doi.org/10.1051/0004-6361/201629837)
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Numerical studies on the link between radioisotopic signatures
on Earth and the formation of the Local Bubble

I. ^{60}Fe transport to the solar system by turbulent mixing of ejecta from nearby
supernovae into a locally homogeneous interstellar medium

M. M. Schulreich¹, D. Breitschwerdt¹, J. Feige¹, and C. Dettbarn²



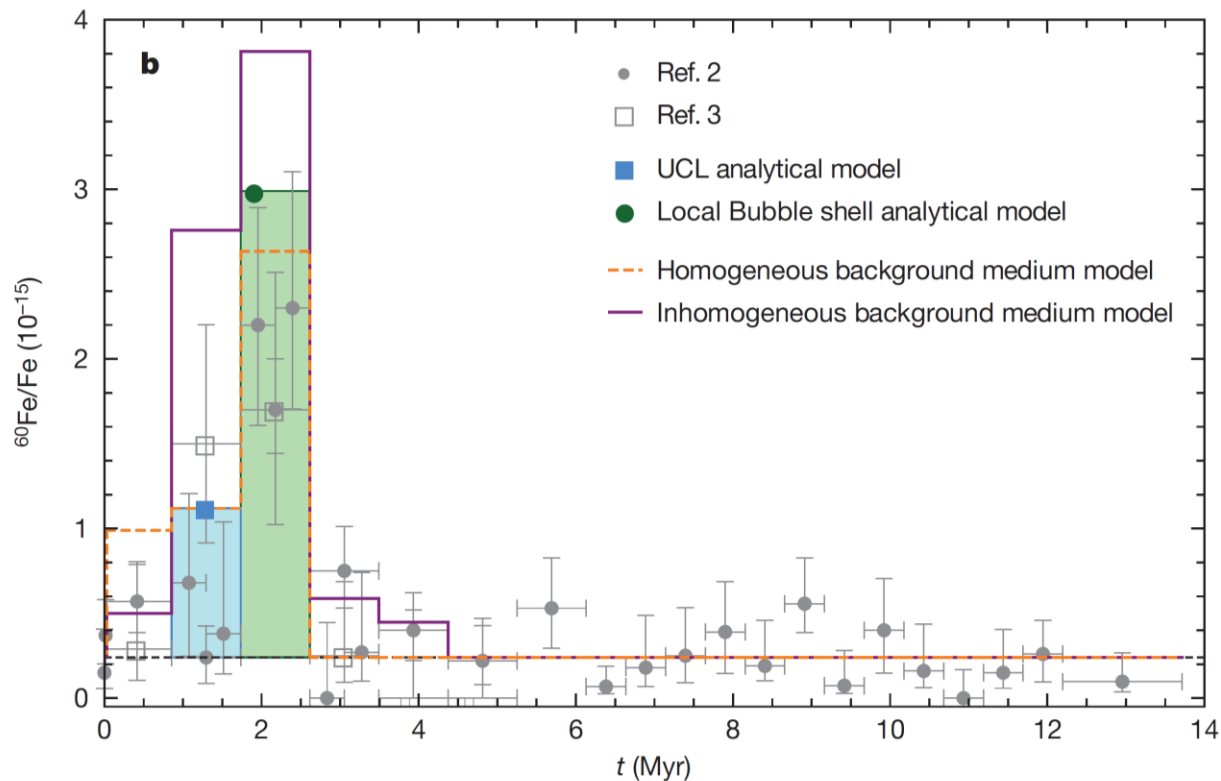
Modelling the ^{60}Fe Signal and the Local Bubble

LETTER

doi:10.1038/nature17424

The locations of recent supernovae near the Sun from modelling ^{60}Fe transport

D. Breitschwerdt¹, J. Feige¹, M. M. Schulreich¹, M. A. de Avillez^{1,2}, C. Dettbarn³ & B. Fuchs³



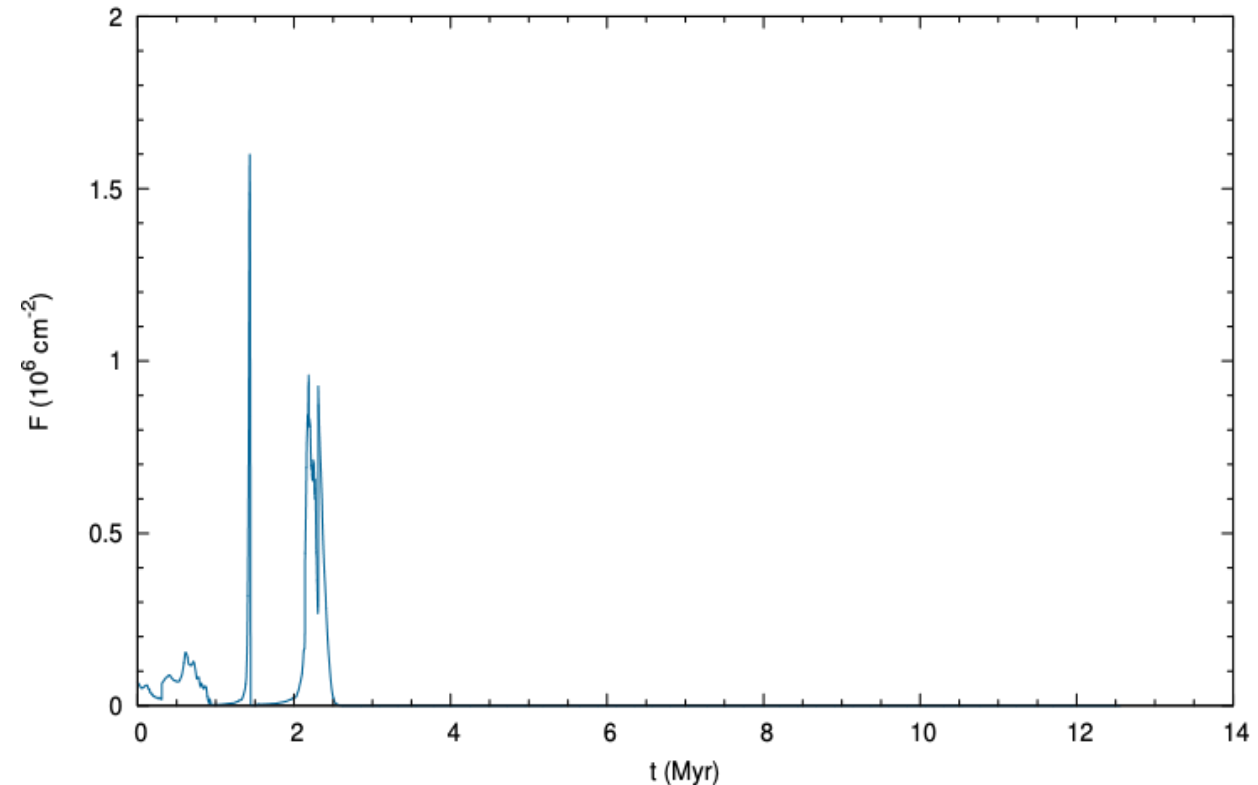
Astronomy
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Modelling the ^{60}Fe Signal and the Local Bubble

Most recent supernovae

- UCL: 2.6 Myr at 106 pc, $9.4 M_{\odot}$
- LCC: 2.3 Myr at 91 pc, $9.2 M_{\odot}$
- UCL: 1.5 Myr at 96 pc, $8.8 M_{\odot}$

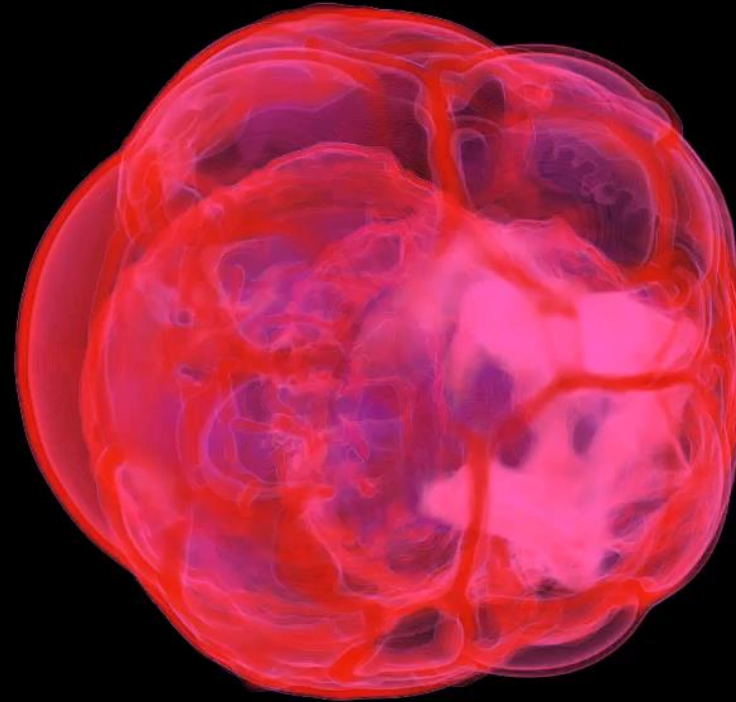
Observable today in

- Libra (UCL)
- Lupus (LCC)

Antares

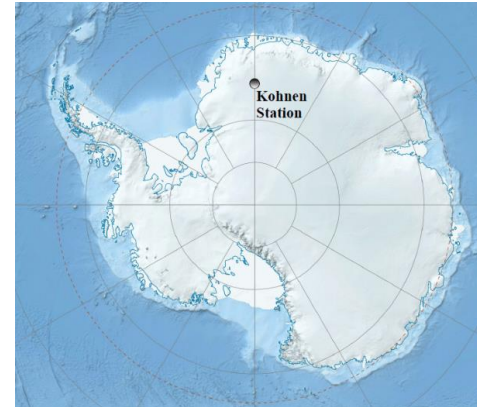
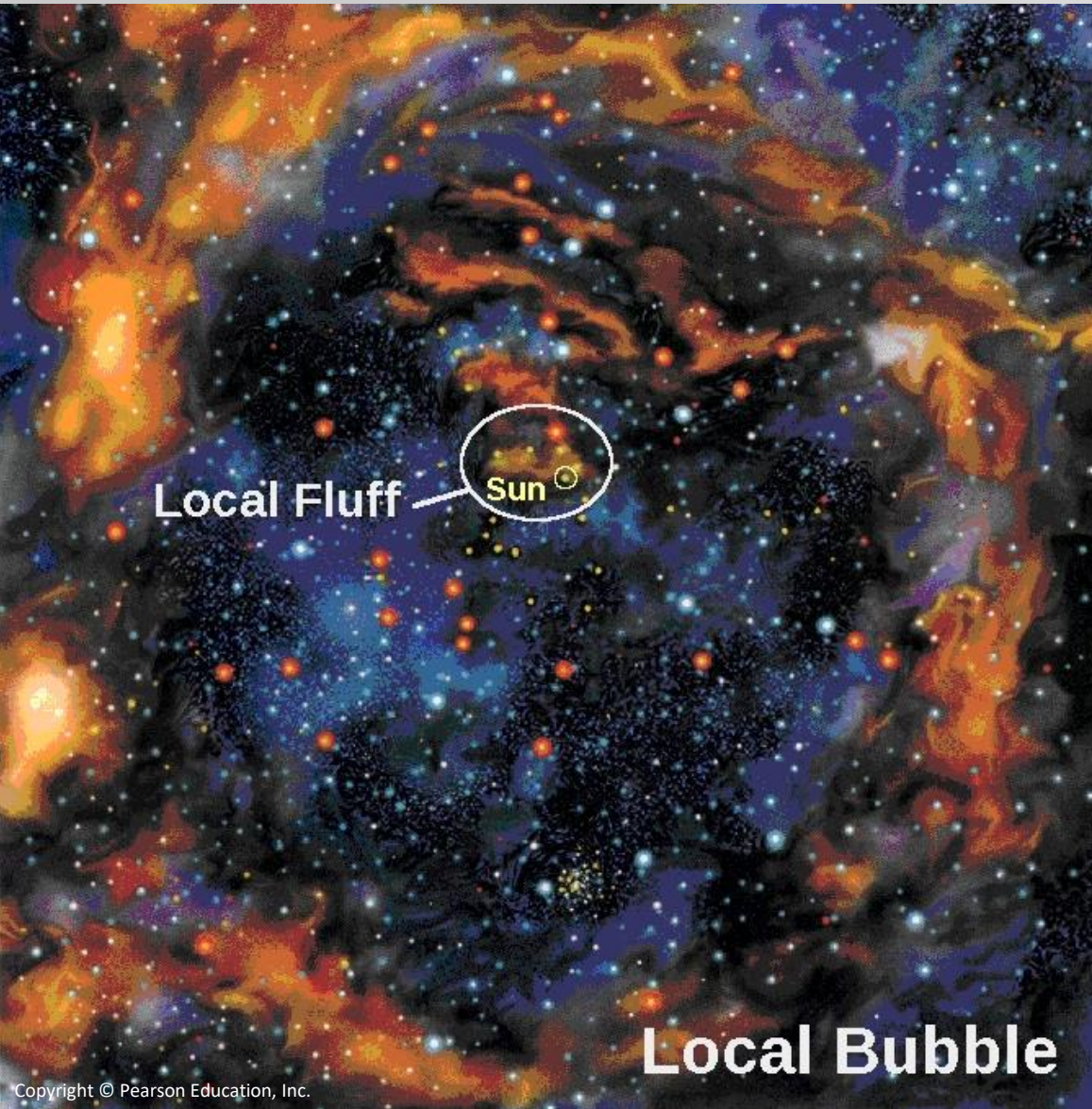


3D simulation of the Local Bubble and the Loop I

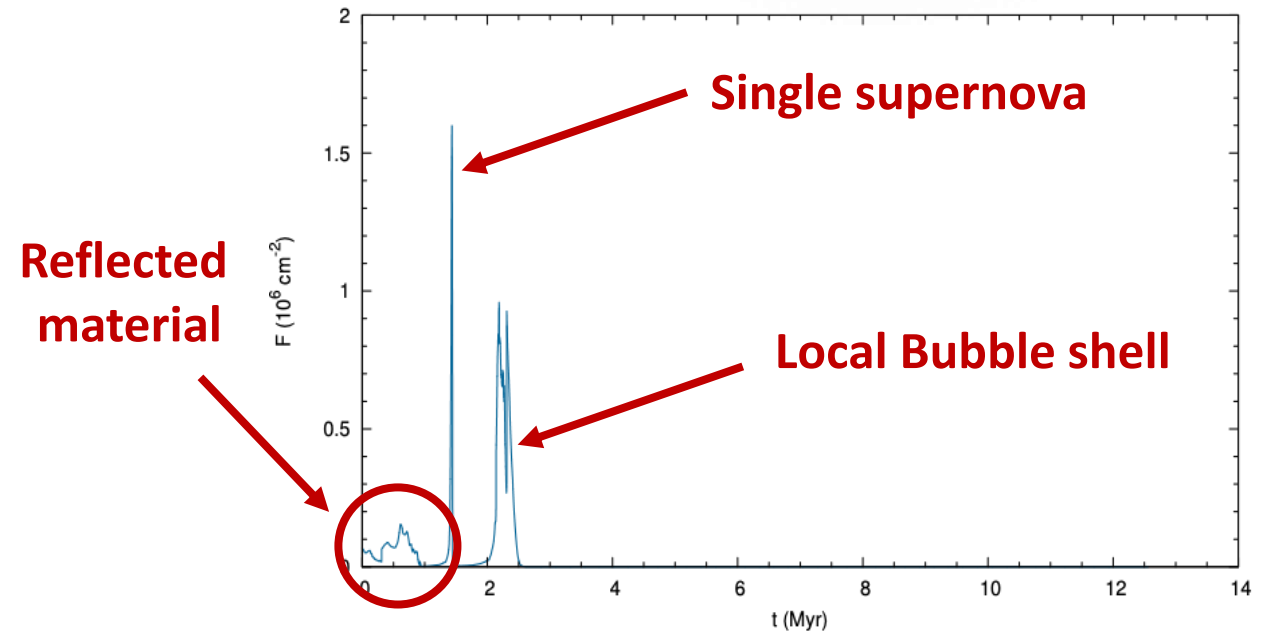


Courtesy: Michael M. Schulreich

Recent Supernova Influx



Koll et al., PRL 123, 072701 (2019), Wallner et al., PNAS 117 (36) 21873 (2020)



Schulreich et al., A&A 604, A81 (2017)

Recent Supernova Influx



Our Sun is moving through the Local Fluff

Local Interstellar Cloud

Sun

G Cloud

Alpha Centauri

Sirius

Koll et al., PRL 123, 072701 (2019)

Take-Home Message III

Subgroups of Sco-Cen association most prominent for producing the 2-3 Myr old ^{60}Fe signal as well as our Local Bubble!

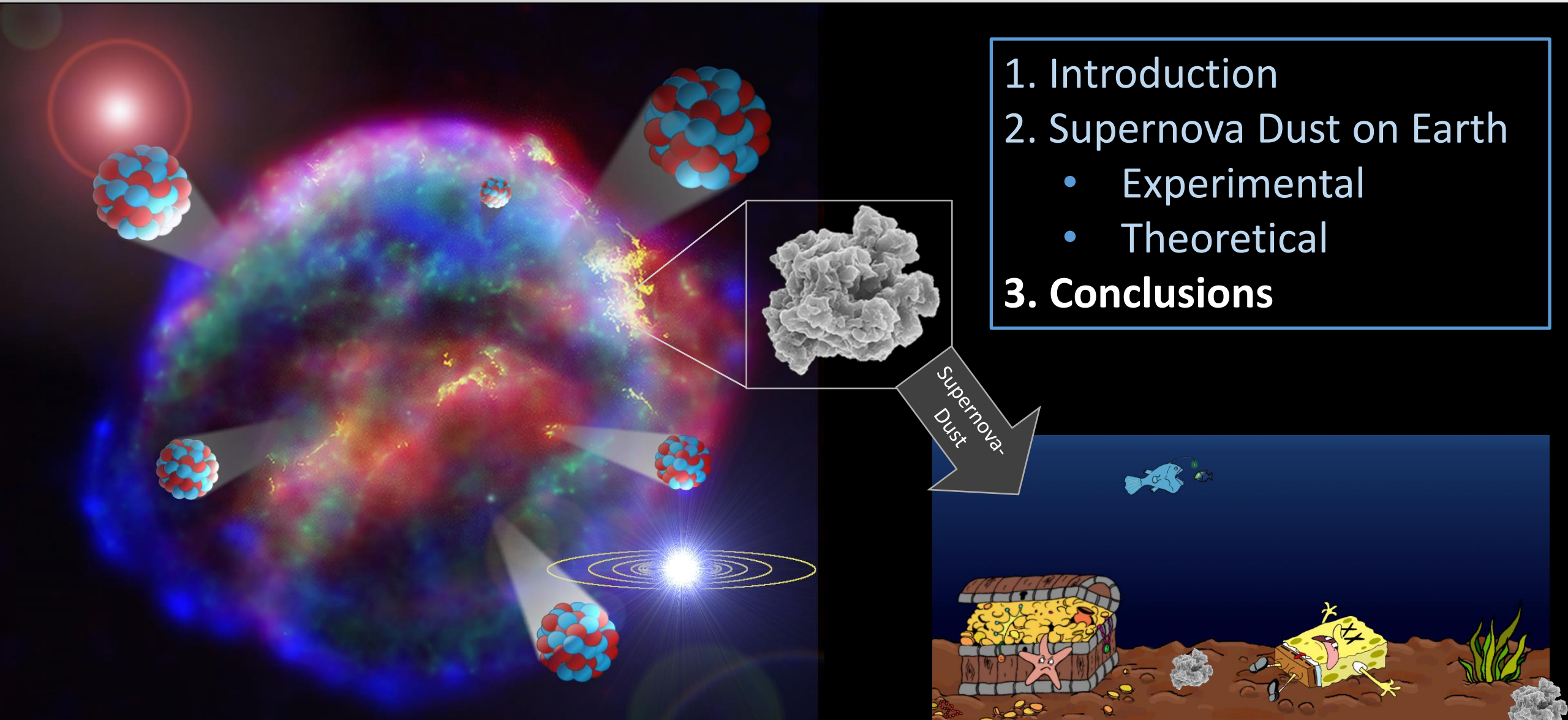
Numerical models point to recent ^{60}Fe Supernova influx on Earth!

Is this influx coinciding with our Sun's travel through the Local Fuff?

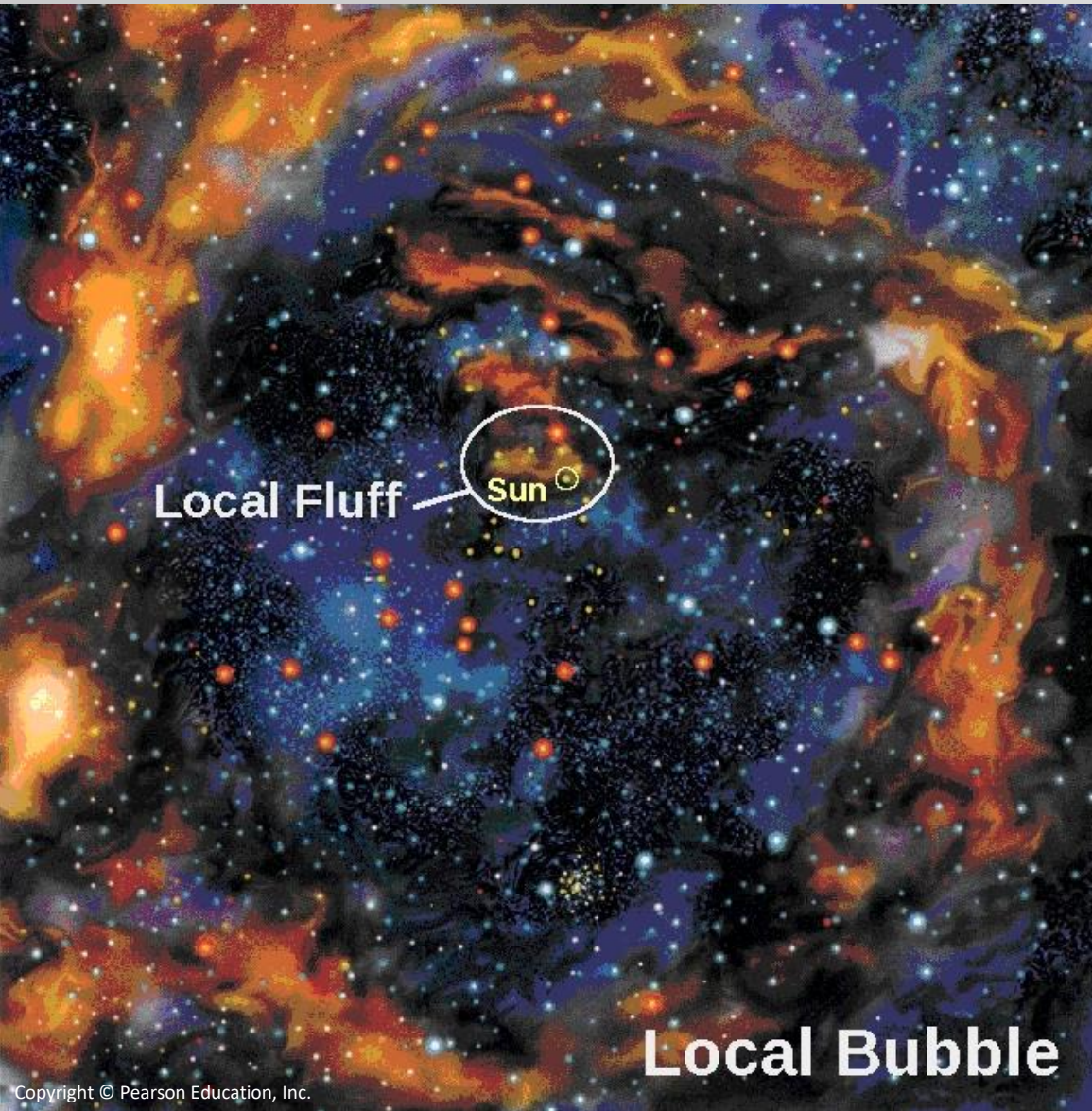
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Supernova Signatures

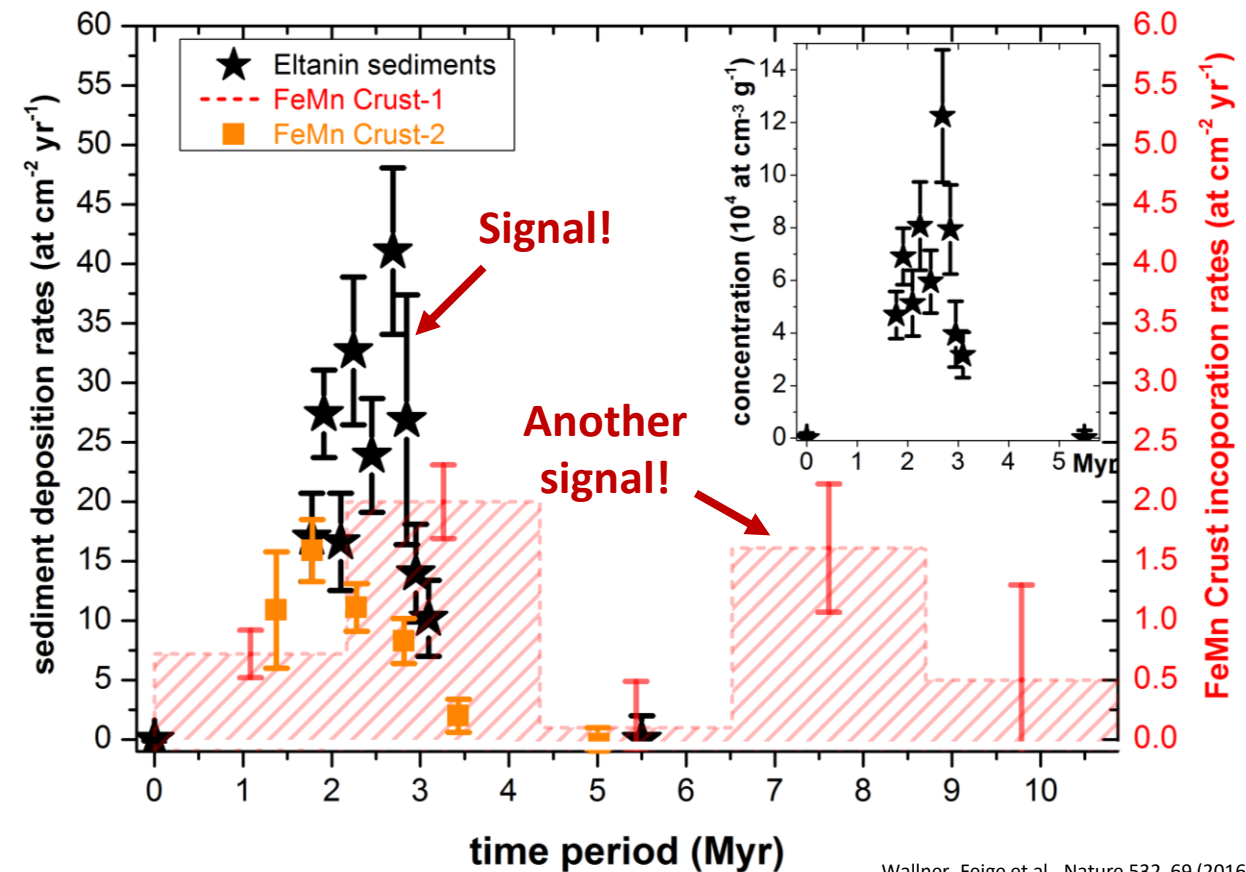


Open questions:


Which stars created the second signal?

Where does the recent influx come from?

What is the $^{60}\text{Fe}/^{26}\text{Al}$ ratio?



Related Studies

PRL 115 , 181103 (2015)	PHYSICAL REVIEW LETTERS	week ending 30 OCTOBER 2015
<hr/>		
		
Signatures of a Two Million Year Old Supernova in the Spectra of Cosmic Ray Protons, Antiprotons, and Positrons		
M. Kachelrieß, ¹ A. Neronov, ² and D. V. Semikoz ³		

Related Studies

PRL **115**, 181103 (2015)

PHYSICAL REVIEW LETTERS

week ending
30 OCTOBER 2015



Signatures of a Two Million Year Old Supernova in the Spectra of Cosmic Ray Protons,
Antiprotons, and Positrons

M. Kachelrieß,¹ A. Neronov,² and D. V. Semikoz³

Science

RESEARCH ARTICLES

Cite as: W. R. Binns *et al.*, *Science*
10.1126/science.aad6004 (2016).

Observation of the ^{60}Fe nucleosynthesis-clock isotope in galactic cosmic rays

W. R. Binns,^{1*} M. H. Israel,^{1*} E. R. Christian,² A. C. Cummings,³ G. A. de Nolfo,² K. A. Lave,¹ R. A. Leske,³
R. A. Mewaldt,³ E. C. Stone,³ T. T. von Rosenvinge,² M. E. Wiedenbeck⁴

Related Studies

PRL **115**, 181103 (2015)

PHYSICAL REVIEW LETTERS

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M. Kachelrieß,¹ A. Neronov,² and D. V. Semikoz³

Science

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MNRAS **498**, 899–917 (2020)

doi:10.1093/mnras/stz2629

Advance Access publication 2019 September 23

Observations of
galactic supernovae

W. R. Binns,^{1*} M.
R. A. Mewaldt,³

A nearby recent supernova that ejected the runaway star ζ Oph, the pulsar PSR B1706–16, and ^{60}Fe found on Earth

R. Neuhäuser,¹  F. Gießler¹ and V. V. Hambaryan ^{1,2}

Related Studies

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M. Kachelrieß,¹ A. Neronov,² and D. V. Semikoz³

PHYSICAL REVIEW LETTERS **125**, 031101 (2020)

Supernova-Produced ^{53}Mn on Earth

G. Korschinek^{1,*} T. Faestermann¹ M. Poutivtsev,^{1,†} A. Arazi^{2,3} K. Knie,^{1,‡} G. Rugel^{1,4} and A. Wallner^{1,4}

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M. Kachelrieß,¹ A. Neronov,² and D. V. Semikoz³

PHYSICAL REVIEW LETTERS **125**, 031101 (2020)

Supernova-Produced ⁵³Mn on Earth

G. Korschinek^{1,*}, T. Faestermann¹, M. Poutivtsev,^{1,†} A. Arazi^{2,3}, K. Knie,^{1,‡} G. Rugel^{1,4} and A. Wallner^{1,4}

Science

Observations
galactic

W. R. Binns,^{1,*}
R. A. Mewaldt,¹

ARTICLE

Received 30 Mar 2014 | Accepted 26 Nov 2014 | Published 20 Jan 2015

DOI: 10.1038/ncomms6956

OPEN

Abundance of live ²⁴⁴Pu in deep-sea reservoirs on Earth points to rarity of actinide nucleosynthesis

A. Wallner^{1,2}, T. Faestermann³, J. Feige², C. Feldstein⁴, K. Knie^{3,5}, G. Korschinek³, W. Kutschera², A. Ofan⁴,
M. Paul⁴, F. Quinto^{2,†}, G. Rugel^{3,†} & P. Steier²

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PHYSICAL REVIEW LETTERS **125**, 031101 (2020)

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THE ASTROPHYSICAL JOURNAL, 827:48 (17pp), 2016 August 10

doi:10.3847/0004-637X/827/1/48

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RADIOACTIVE IRON RAIN: TRANSPORTING ⁶⁰Fe IN SUPERNOVA DUST TO THE OCEAN FLOOR

BRIAN J. FRY¹, BRIAN D. FIELDS¹, AND JOHN R. ELLIS²

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Science

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Supernova-Produced ⁵³Mn on Earth

G. Korschinek^{1,*}, T. Faestermann¹, M. Poutivtsev,^{1,†} A. Arazi^{2,3}, K. Knie,^{1,‡} G. Rugel^{1,4} and A. Wallner^{1,4}

THE ASTROPHYSICAL JOURNAL, 840:105 (9pp), 2017 May 10

<https://doi.org/10.3847/1538-4357/aa6c57>

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A Supernova at 50 pc: Effects on the Earth's Atmosphere and Biota

A. L. Melott¹, B. C. Thomas², M. Kachelrieß³, D. V. Semikoz^{4,5}, and A. C. Overholt⁶

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BRIAN J. FRY¹, BRIAN D. FIELDS¹, AND JOHN R. ELLIS²

Conclusions

The *Australopithecus afarensis*
2-3 Myr ago...



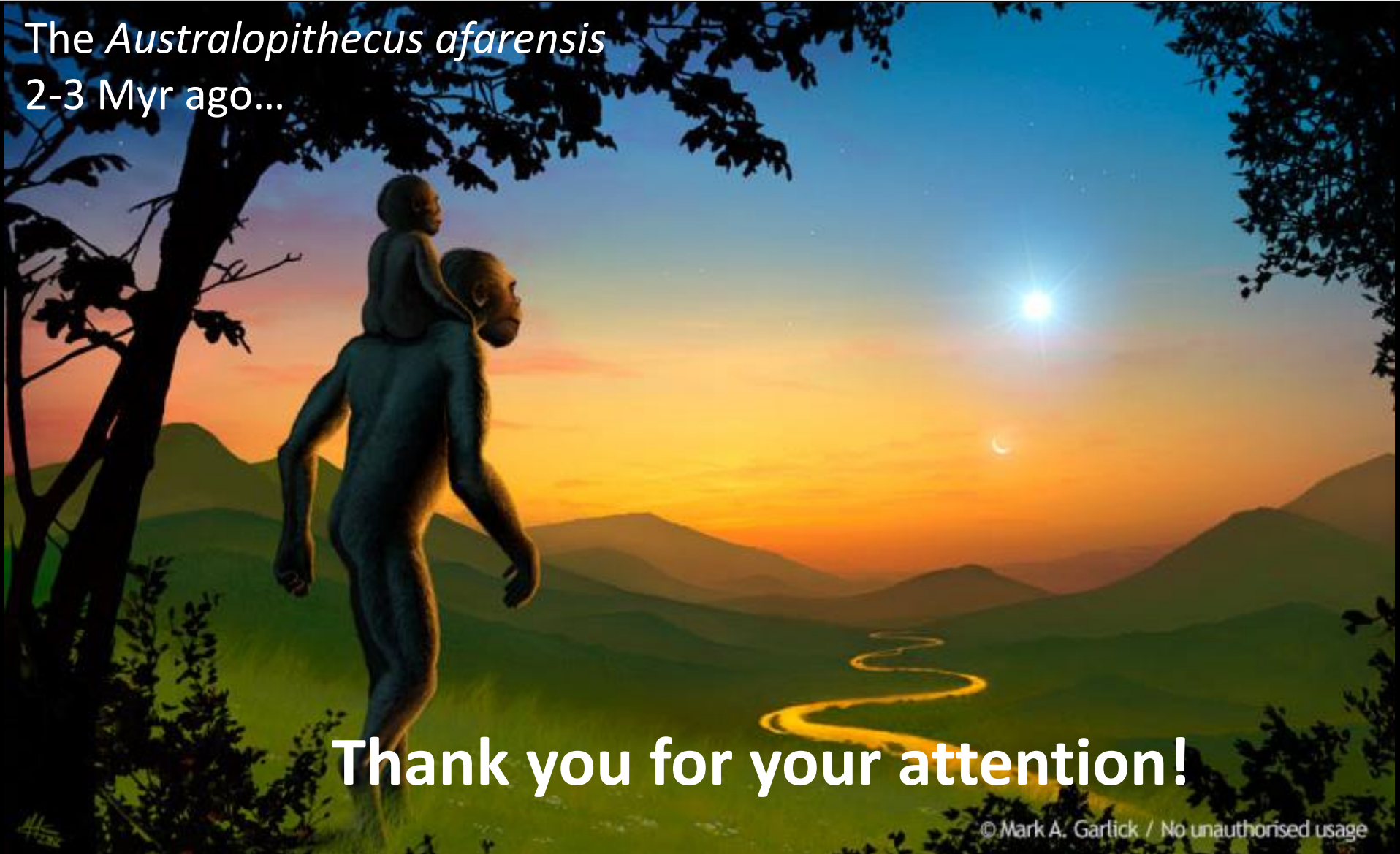
**Nearby Supernova Explosions May
Have Affected Human Evolution**

**When Stars Attack! Near-Earth Supernova
Explosions Threat, Effects and Kill Zone**

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Conclusions

The *Australopithecus afarensis*
2-3 Myr ago...



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