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Is Dark Matter made of Primordial Black Holes?

Günther Hasinger

Polish-German-WE-Heraeus Seminar

MAX BORN SYMPOSIUM

Synagoge Görlitz, 5. December 2023

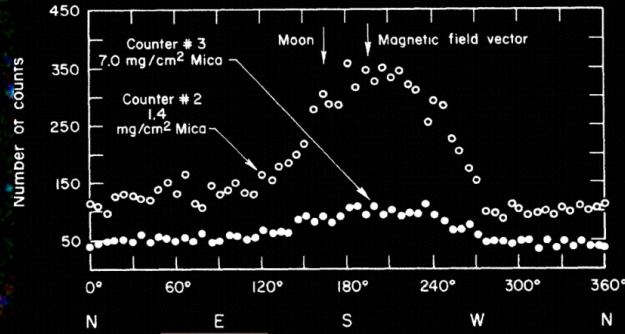
Wissenschaftsjahr 2023

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The X-ray background

ROSAT 1998

Giacconi 1962



ROSAT 1991

XMM-Newton 2012

The background is the echo of black hole formation and growth over cosmic time.

ROSAT 1998

Chandra 2011



Keck >1994

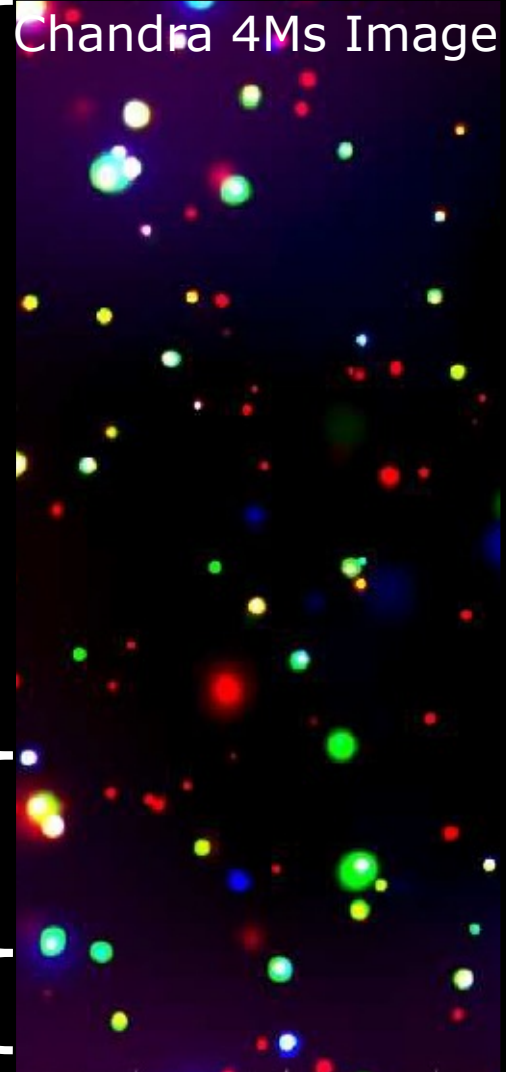
VLT >2001

Resolving the 0.5-2 keV X-ray background

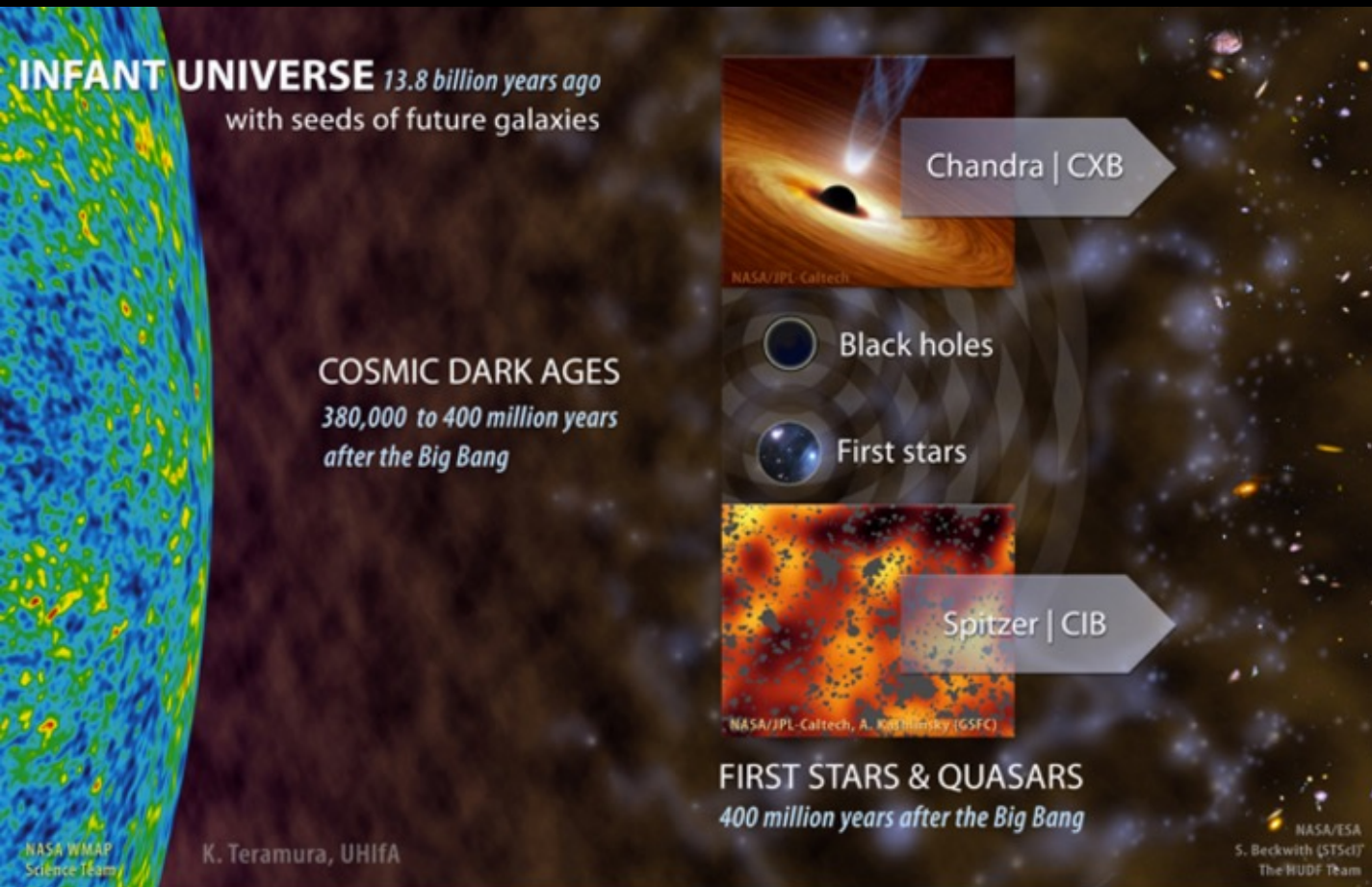
75-80% ROSAT sources

90-95% Chandra/XMM sources

Few % Stacking Galaxies
1% Correlation with CIB →



CIB x CXB fluctuations indicate high-z BH population



Significant cosmic background fluctuations have been found both in the NIR and in X-rays.

The strong CIB/CXB cross-correlation signal indicates a substantial contribution of Black Holes to the signal.

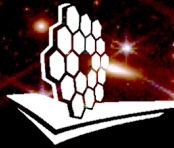
There is no correlation with fluctuations in the deepest HST images, therefore the signal likely comes from redshifts $z > 13$.

Large angular scale also points to high- z origin.

Could these be primordial?

Cappelluti et al., 2013

The first JWST Deep Image revealed by President Biden!



→ THE EUROPEAN SPACE AGENCY

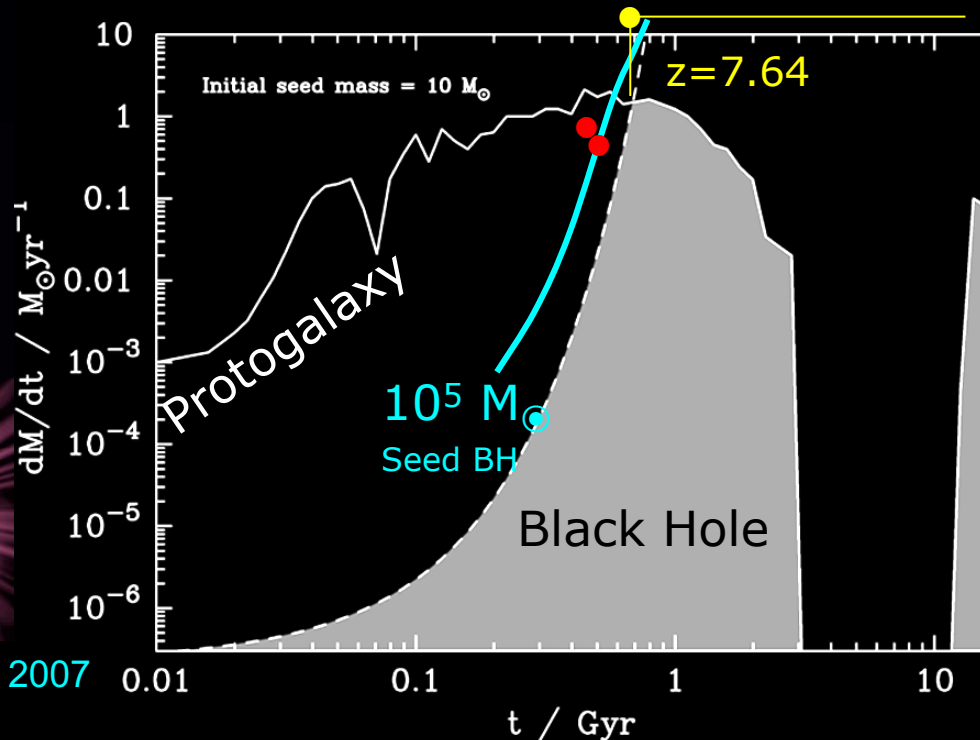
How to produce the first proto-quasars



$z=12.75$



Li, Hernquist, et al. 2007



$10^9 M_{\odot}$
known QSO

$z = 7.54$ 7.642
 $8 \times 10^8 M_{\odot}$ $1.6 \times 10^9 M_{\odot}$

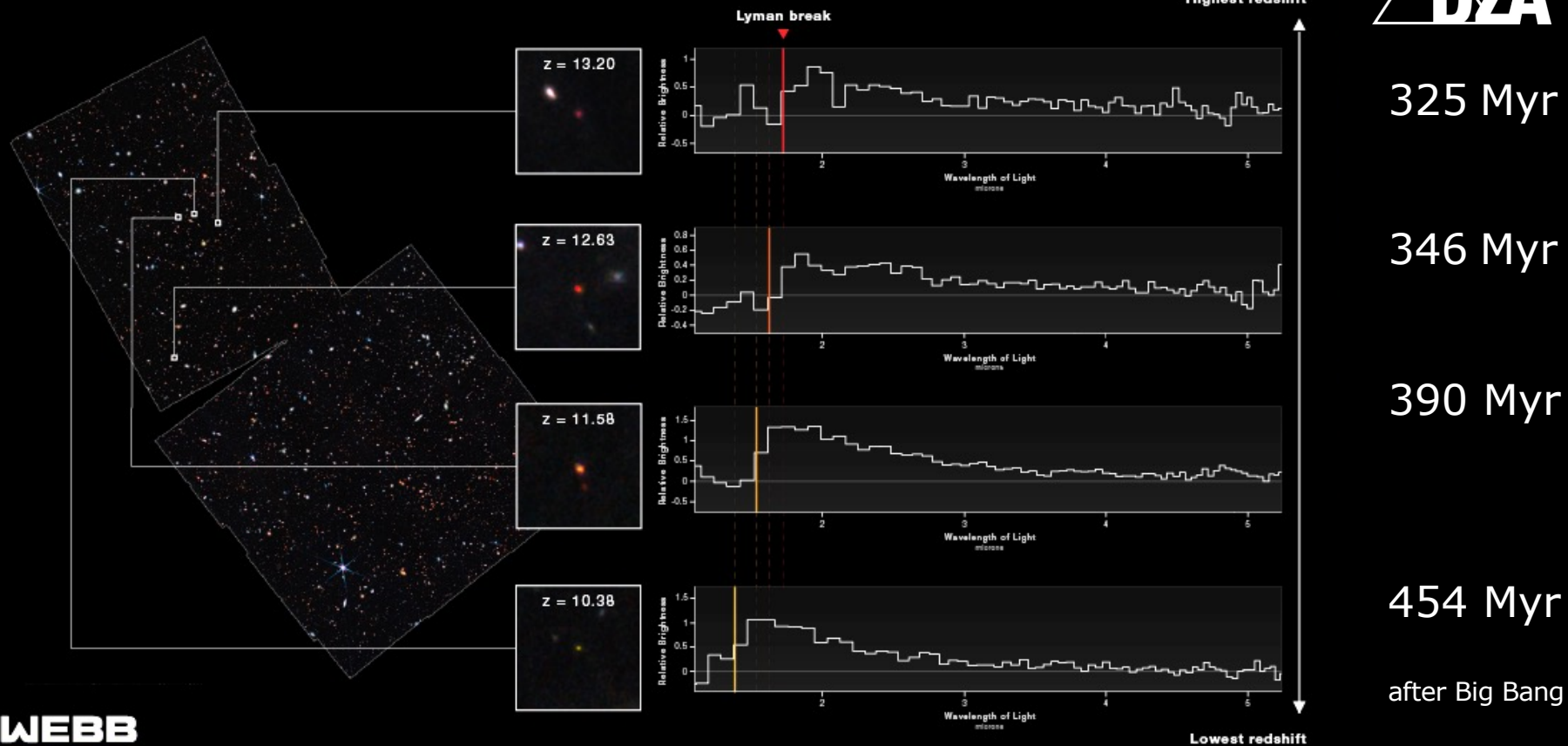
ULAS J1342+0928
(Bañados et al. 2018)

J0313-1806
(Wang et al. 2021)

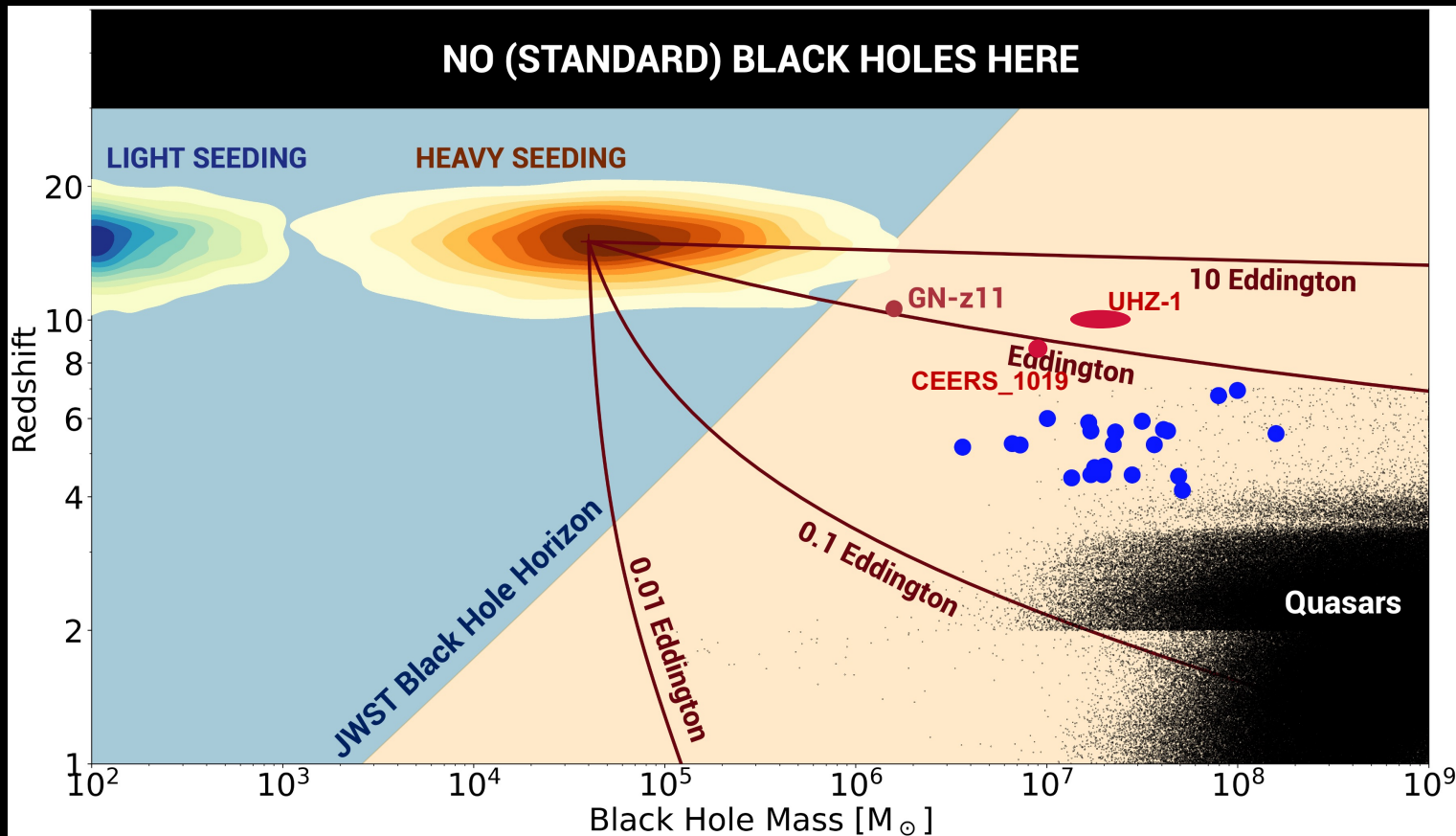
Larsen et al. 2023 &
Goulding et al. 2023:
JWST discoveries of
most distant BHs at
 $z=9$ & 10.1 at $10^{7-8} M_{\odot}$

Archibald et al., 2001

Need massive ($10^{5-7} M_{\odot}$) seed Black Holes early in the Universe !



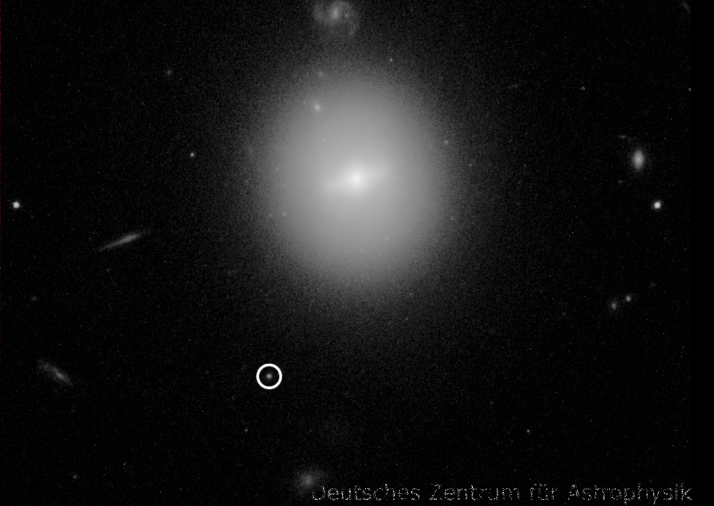
Evidence for Heavy Seeds gets stronger!



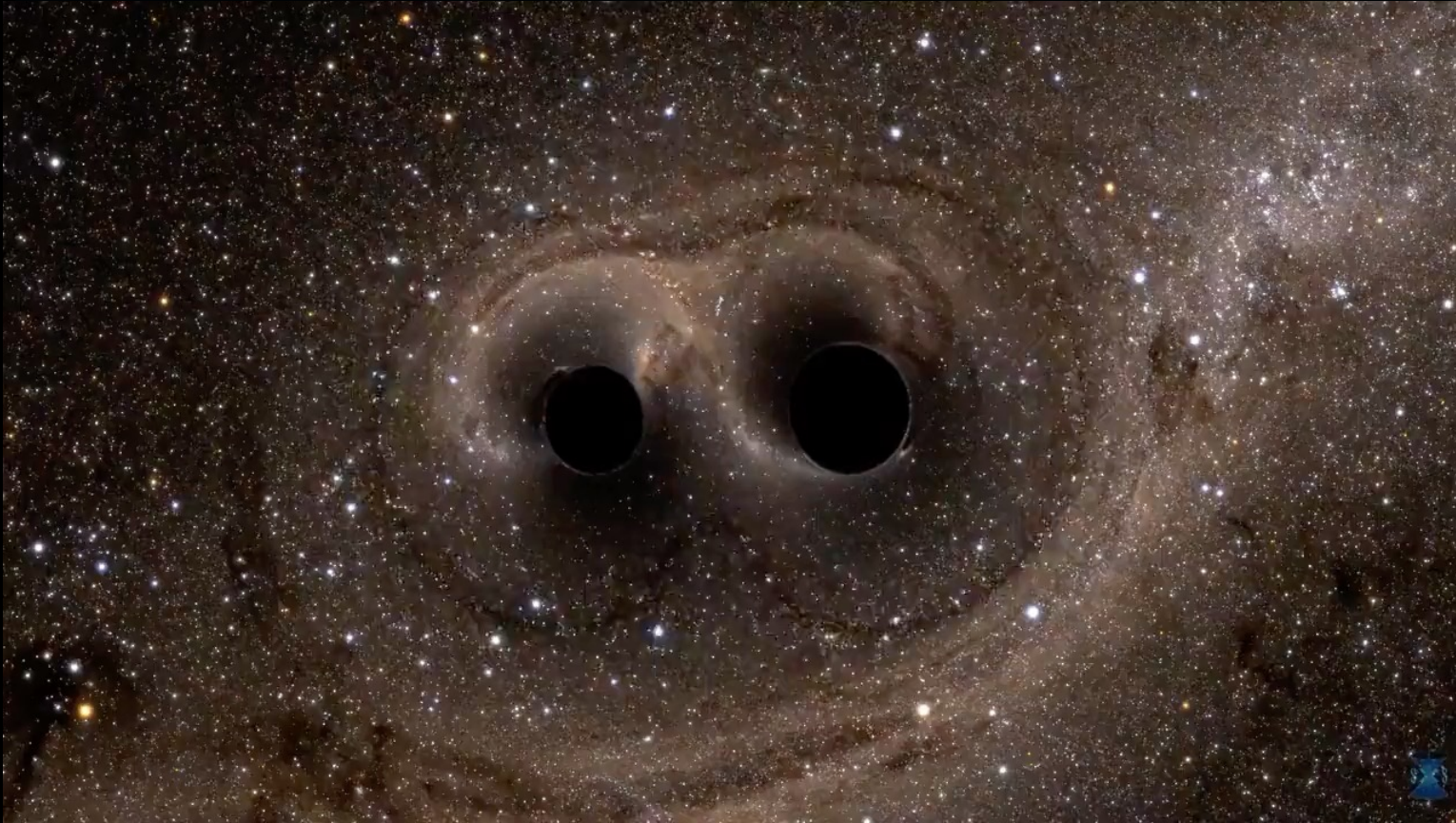
Hubble reveals an “Intermediary” Black Hole



The X-ray telescopes XMM-Newton and Chandra have detected a Black Hole, which is just in the process of disrupting and swallowing a star. The Hubble Space Telescope showed that this is an “Intermediary” Black Hole of about 50,000 solar masses in a globular cluster of another galaxy.

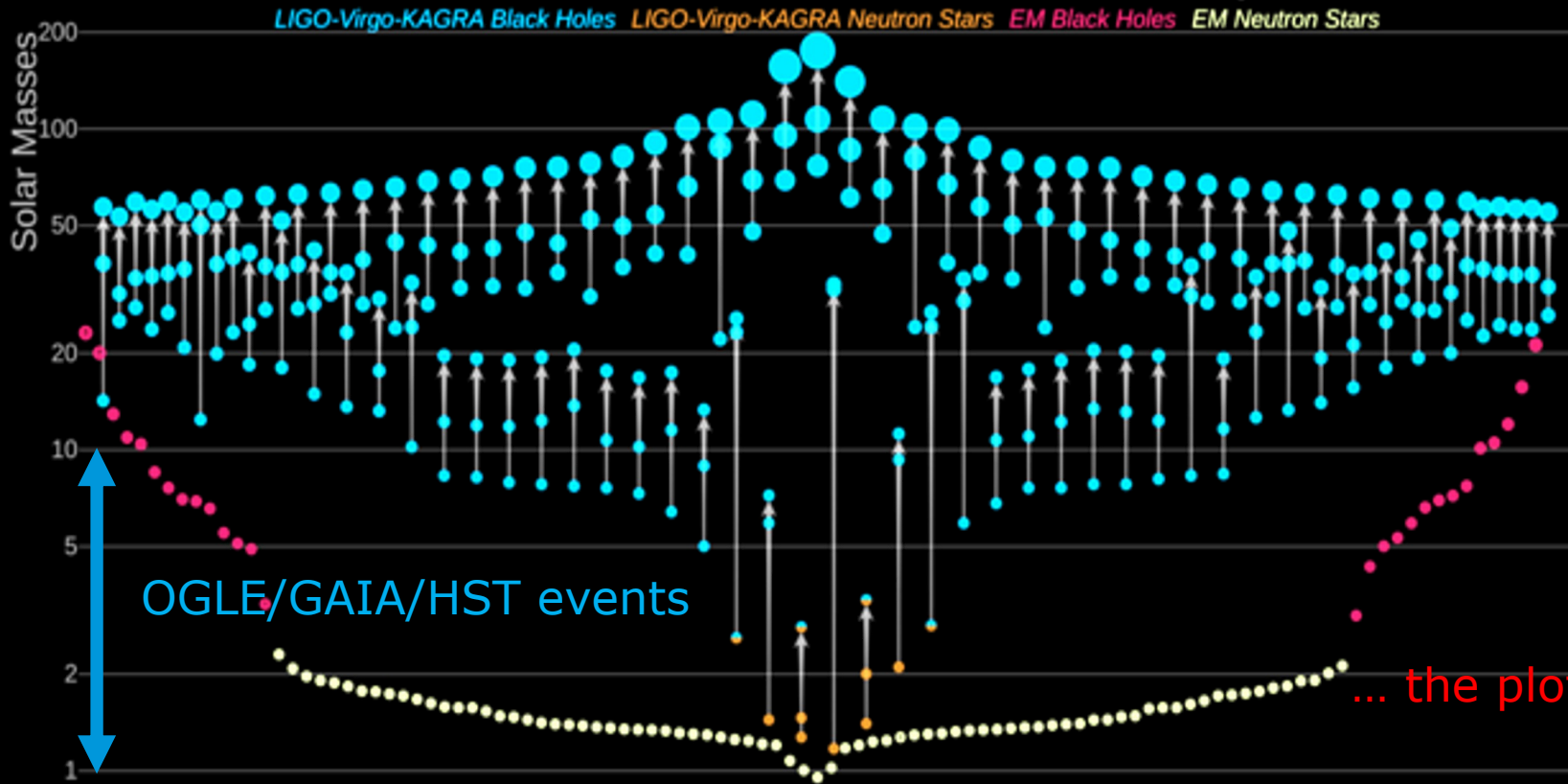


Simulation of the merger between two Black Holes



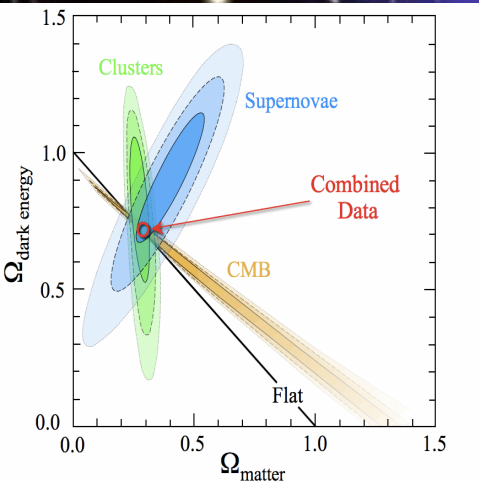
LIGO/Virgo/
Kagra/GEO 600
observations
start in May
2023!

Masses in the Stellar Graveyard



LIGO-Virgo-KAGRA | Aaron Geller | Northwestern

What is Dark Matter?



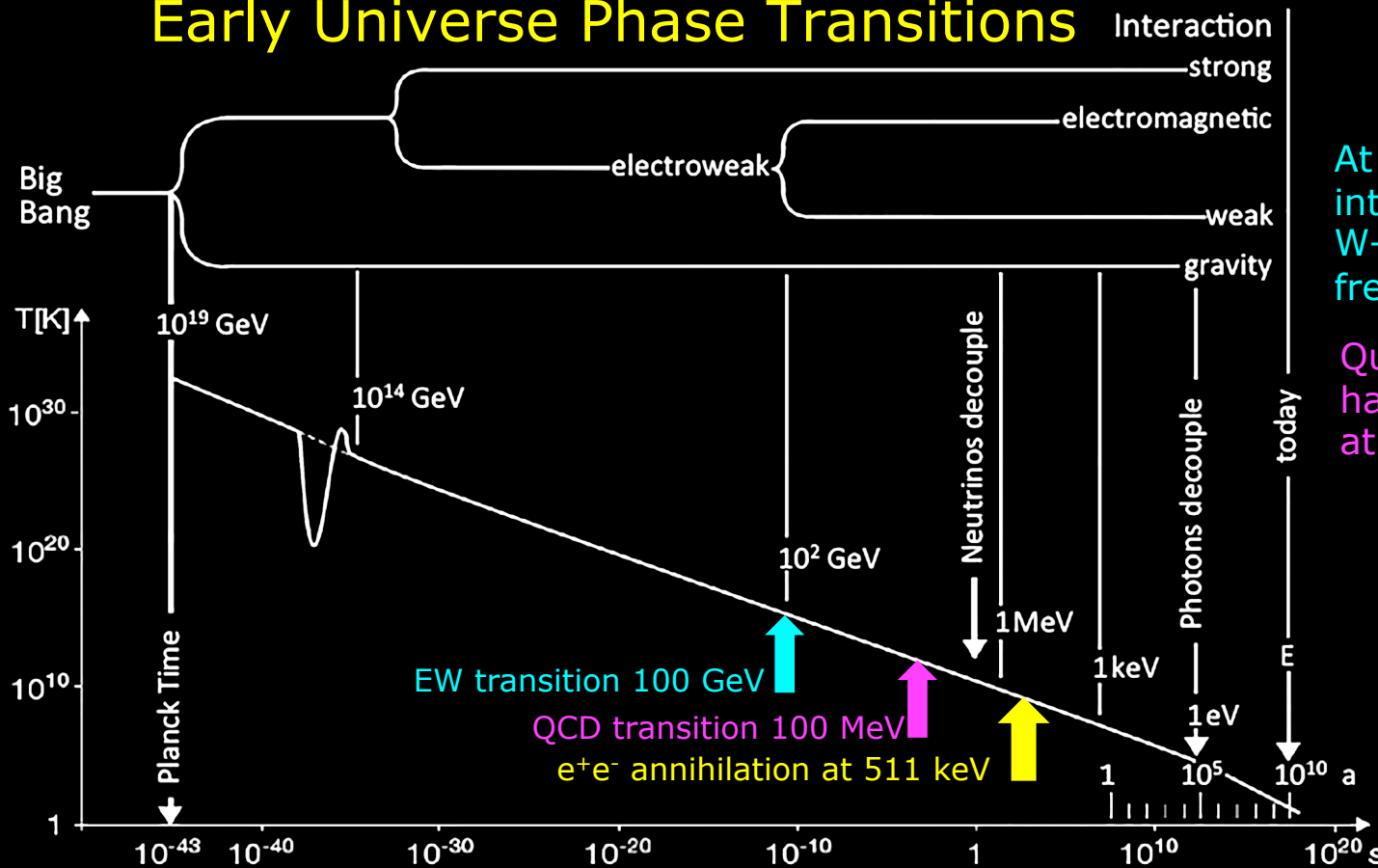
Generations of physicists are desperately searching the Dark Matter particle. Primordial Black Holes may be an alternative.

Early Universe Phase Transitions



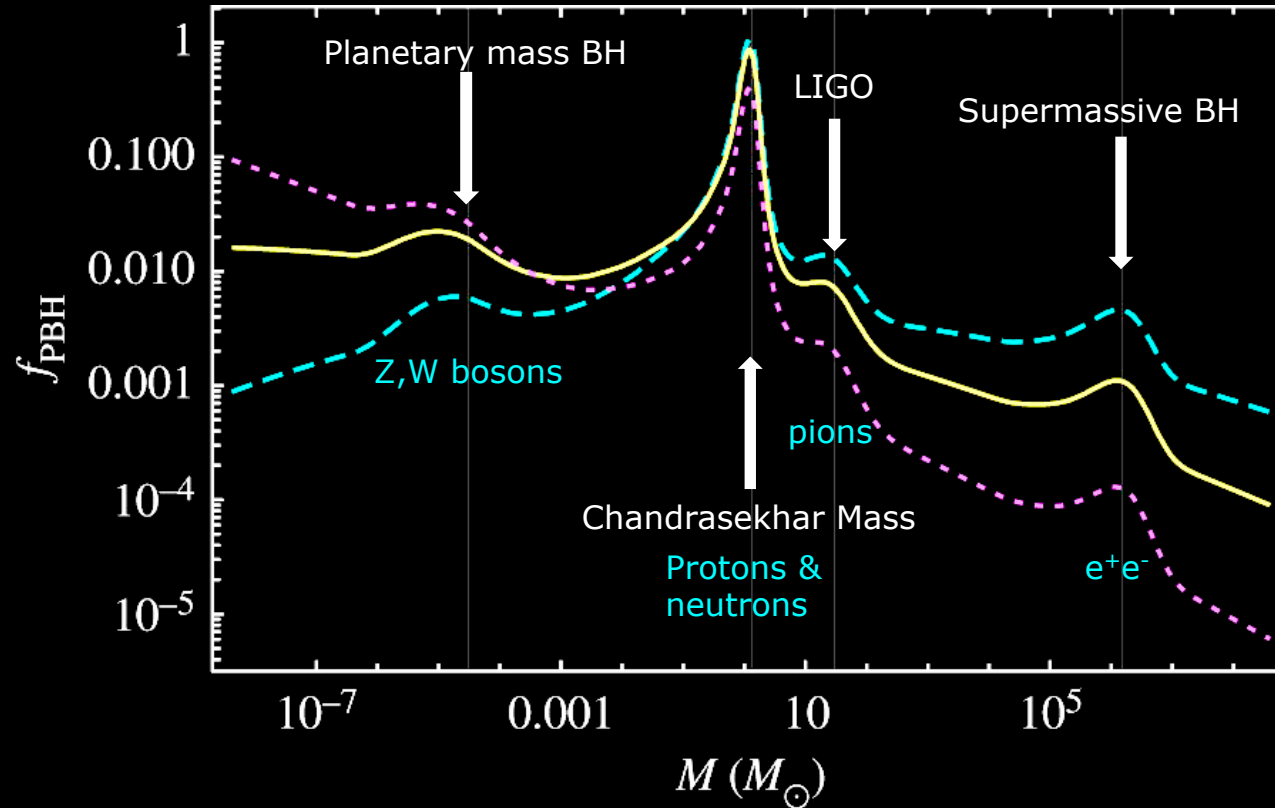
At the electroweak interaction scale the first W- and Z-bosons are freezing out.

Quarks freeze out to form hadrons (baryons, pions) at the QCD transition.



Each of these transitions leaves an imprint on the thermal history of the universe and enables black hole collapse.

PBH Mass Spectrum



Different peaks correspond to different particles created at the early universe phase transitions and the corresponding reduction in the sound velocity.

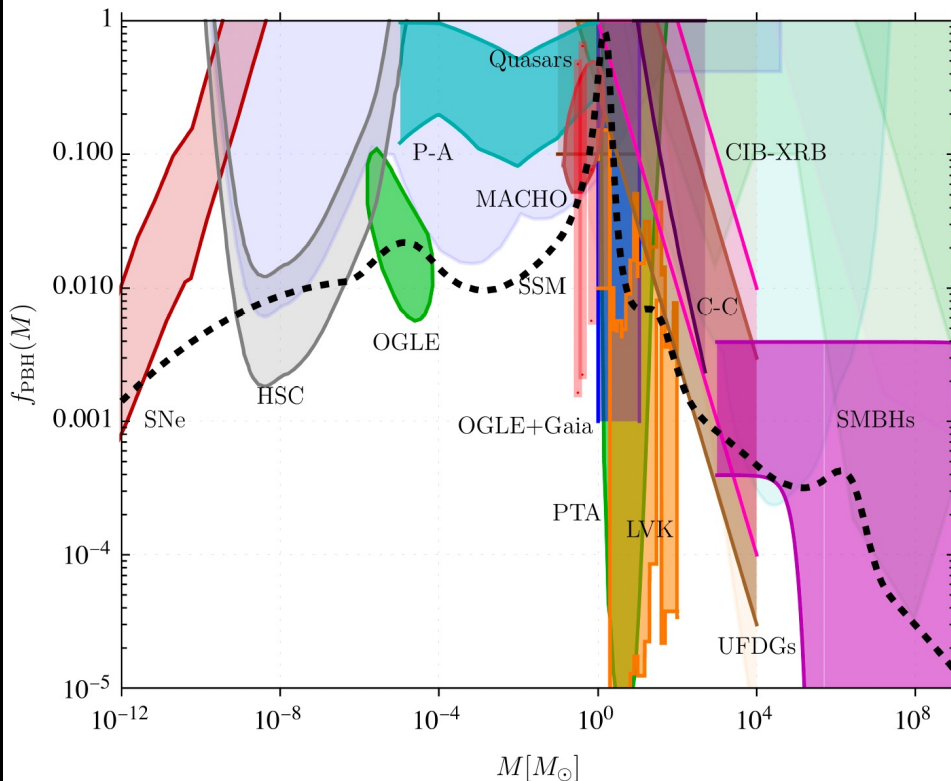
BH mass corresponds to the horizon size at each time.

Only requirement is enough fluctuation power in a volume fraction of 10^{-9} of the early Universe.

Carr, Clesse, García-Bellido 2019

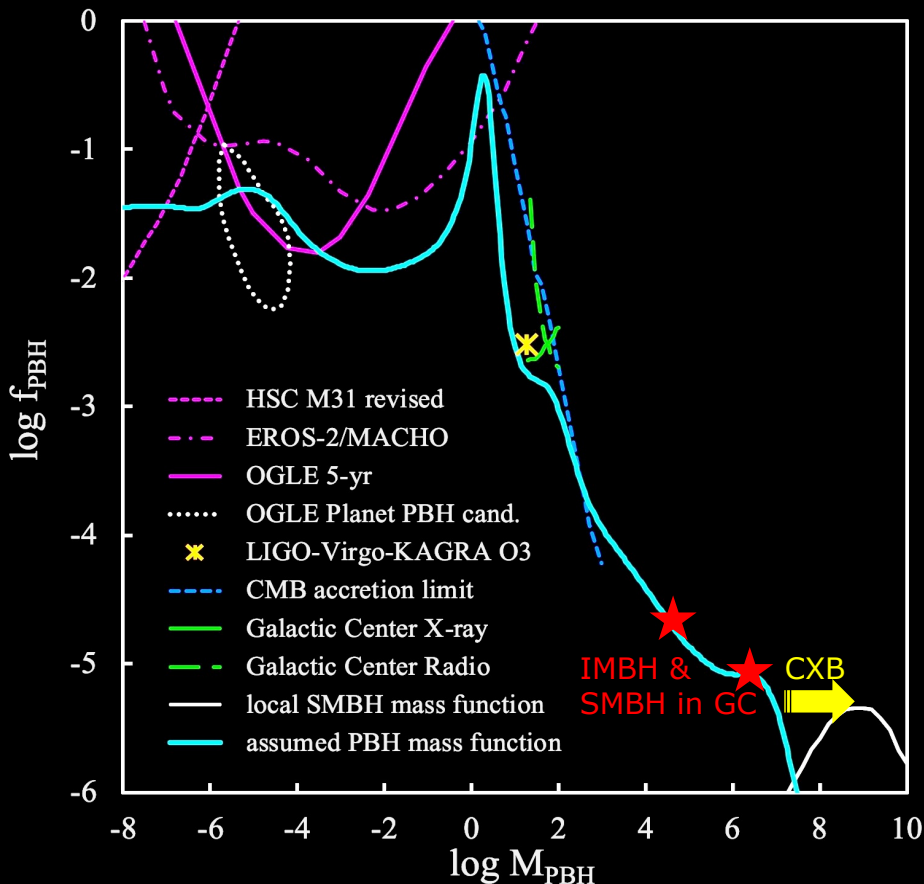
Observational Evidence for Primordial Black Holes: A Positivist Perspective

B. J. Carr,^{1,*} S. Clesse,^{2,†} J. García-Bellido,^{3,‡} M. R. S. Hawkins,^{4,§} and F. Kühnel^{5,¶}



arXiv:2306.03903v1 [astro-ph.CO] 6 Jun 2023

PBH mass spectrum assumed for our work



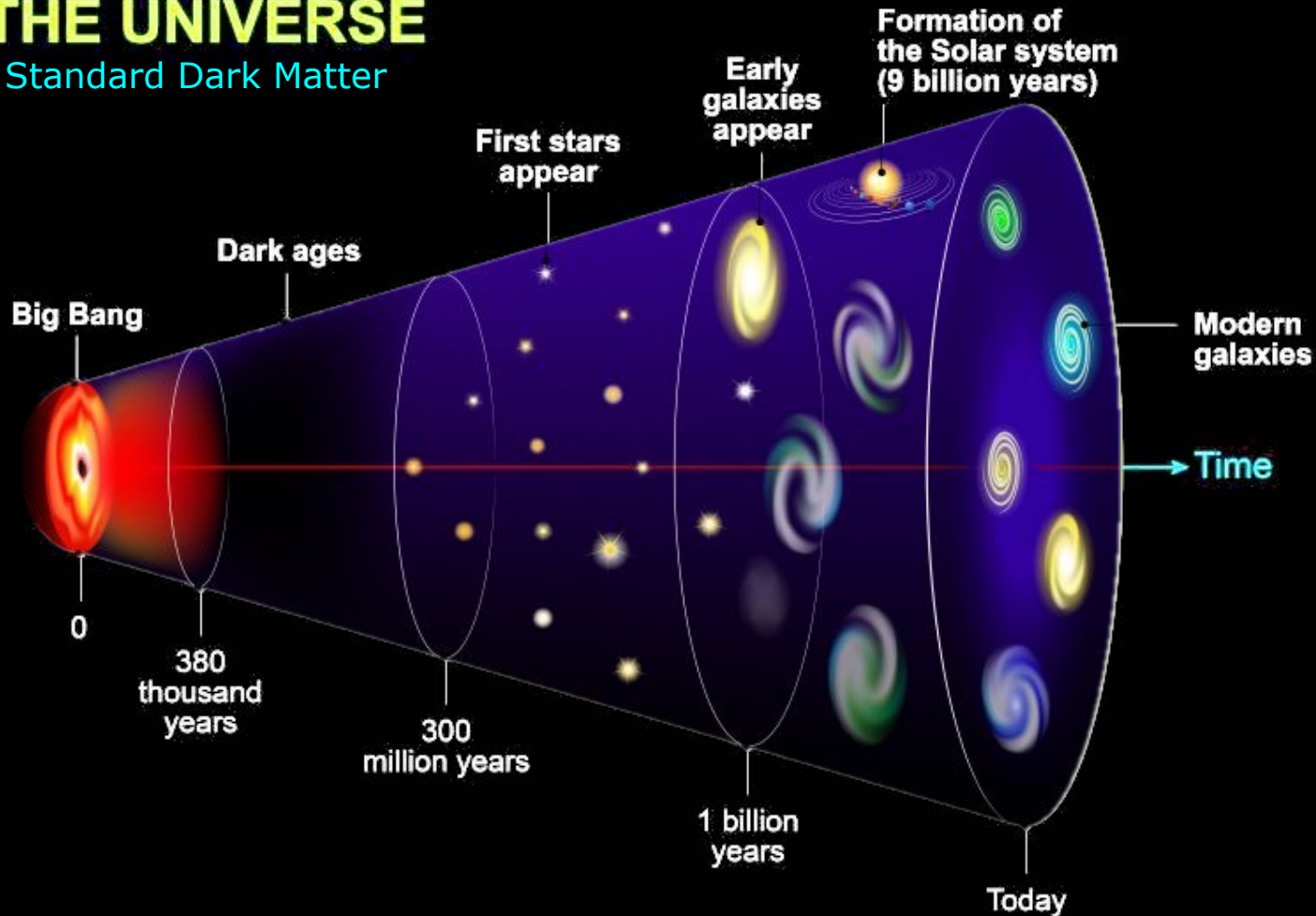
Bernard Carr, Juan García-Bellido et al. are working on a new version of their PBH mass spectrum, which assumes a rolling index of the primordial power spectrum and thus has a steeper decline at large PBH masses. This is now fully consistent with all observational constraints.

This is, what we use to estimate the PBH contribution to the extragalactic backgrounds.

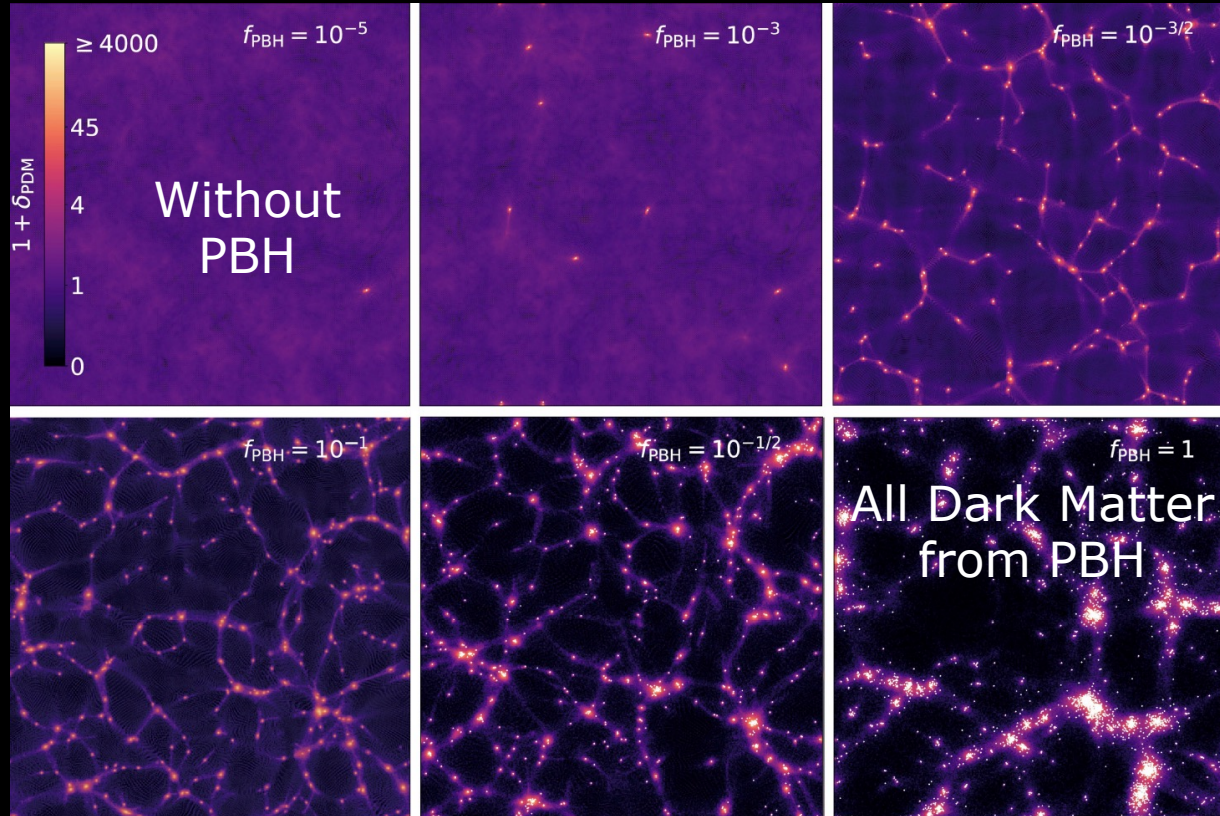
Hasinger 2020

EVOLUTION OF THE UNIVERSE

Standard Dark Matter

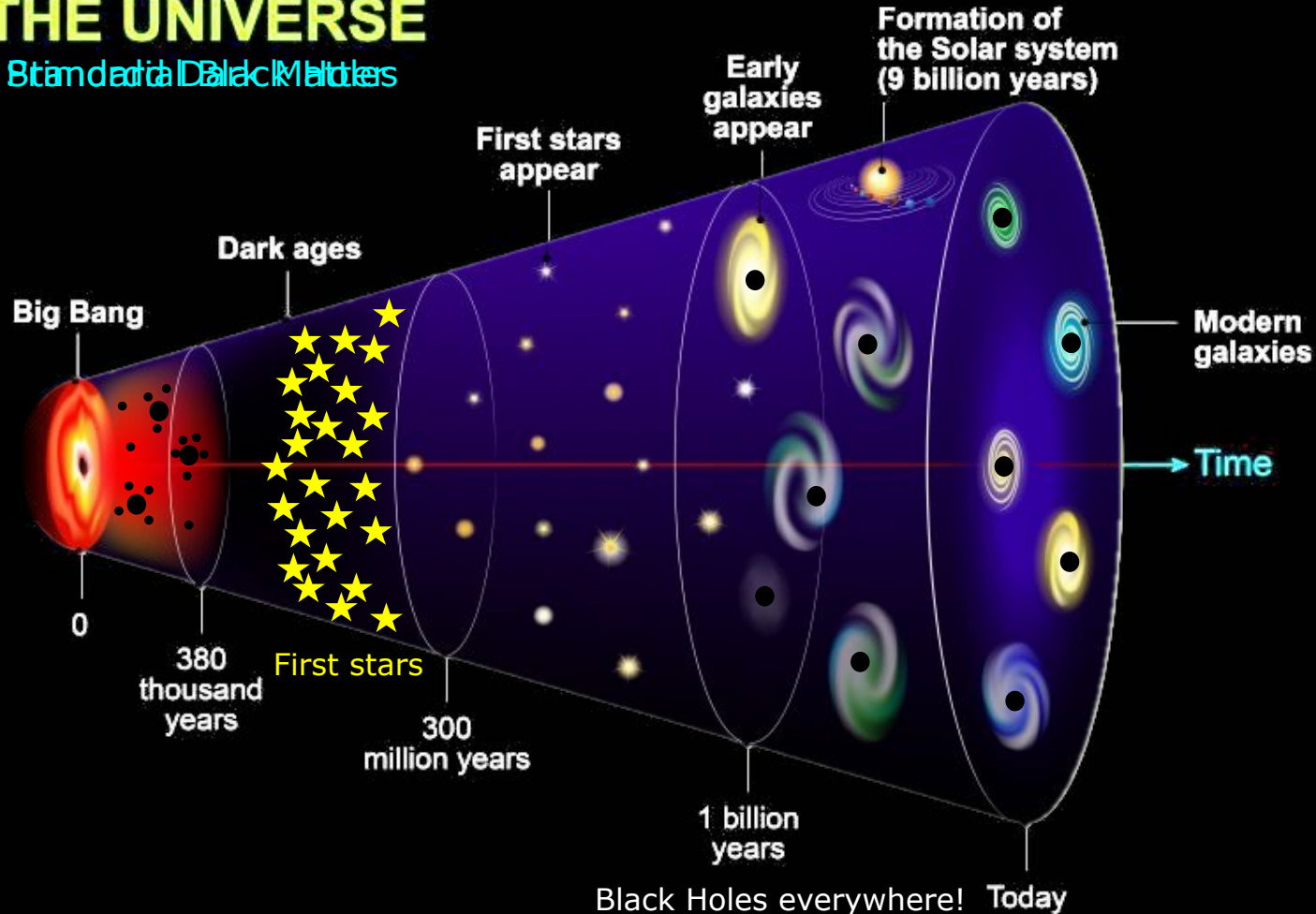


Growth of Large-Scale Structure at $z=10$



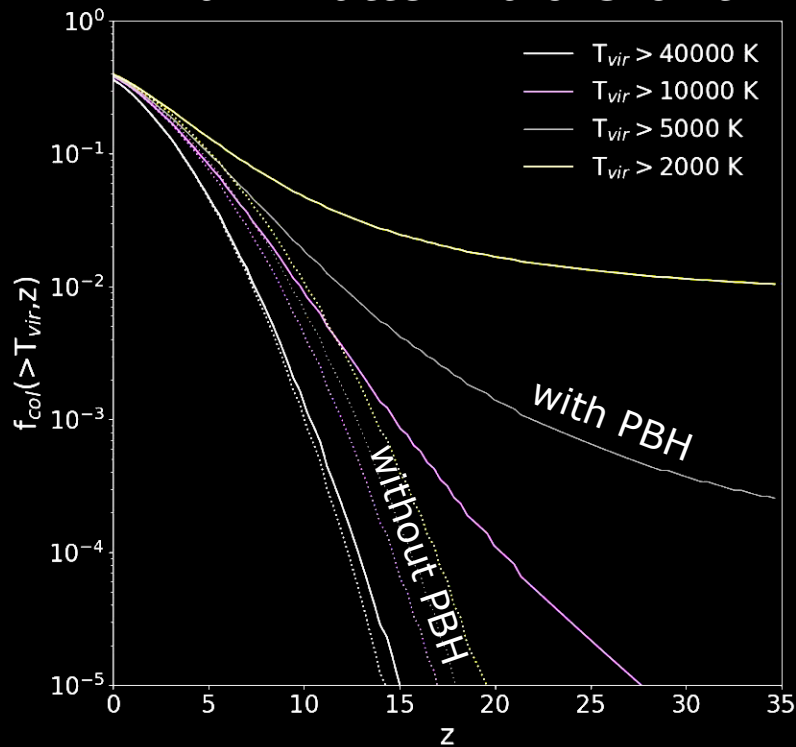
EVOLUTION OF THE UNIVERSE

Standard Model

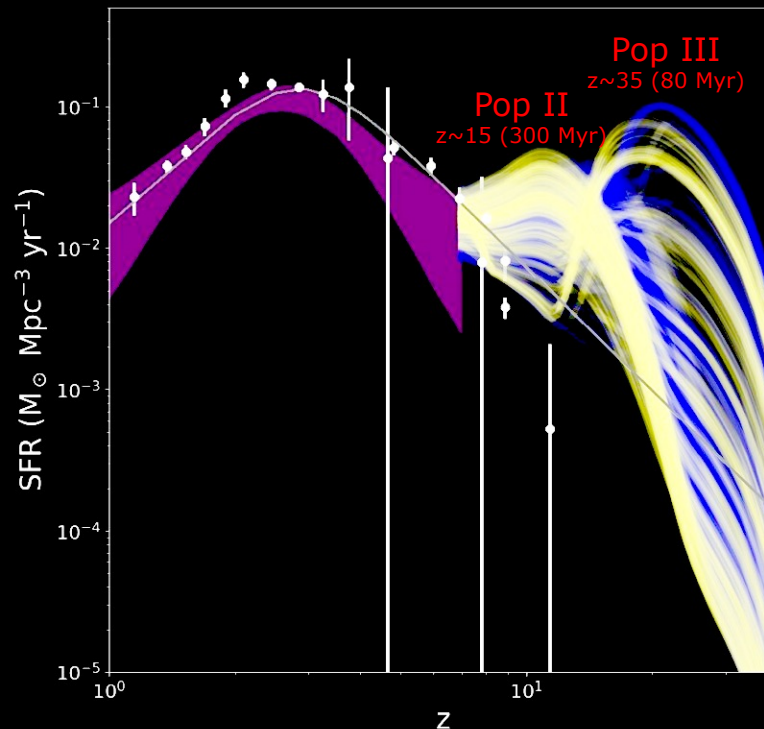


PBH add small DM haloes and early star formation

Dark Matter Halo Growth



Star Formation History

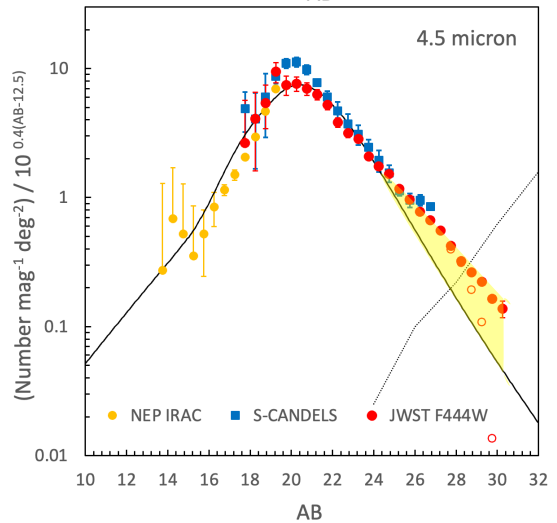
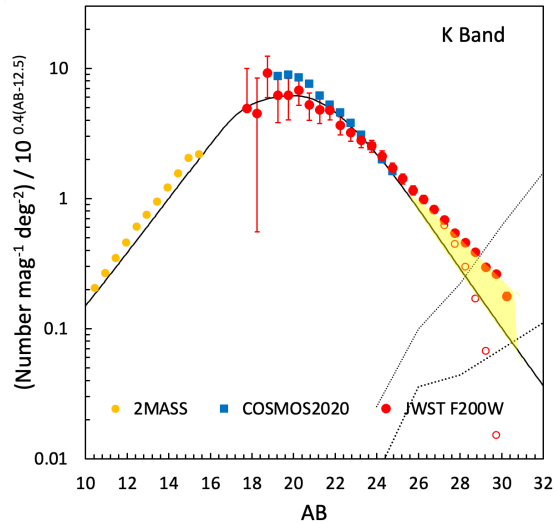
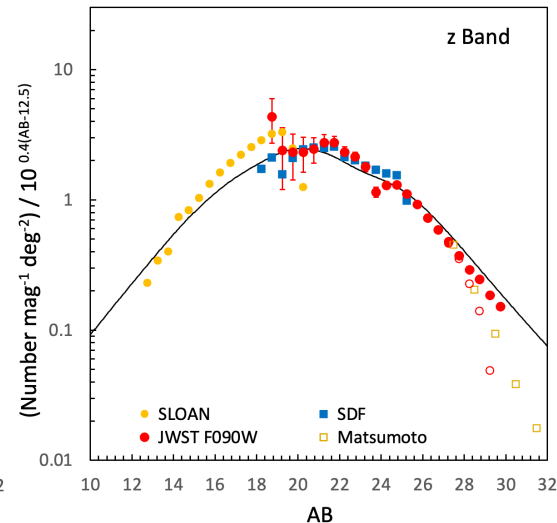
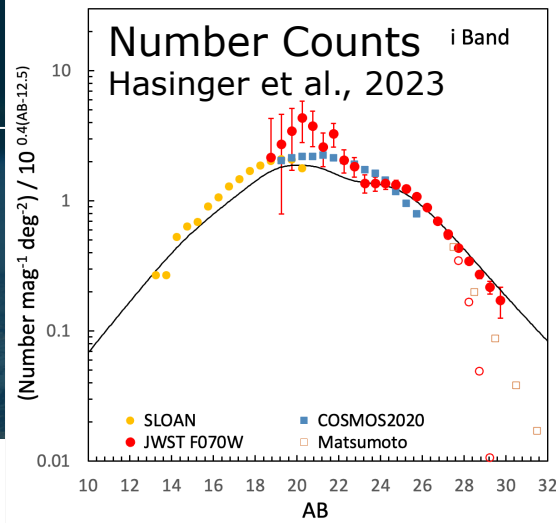
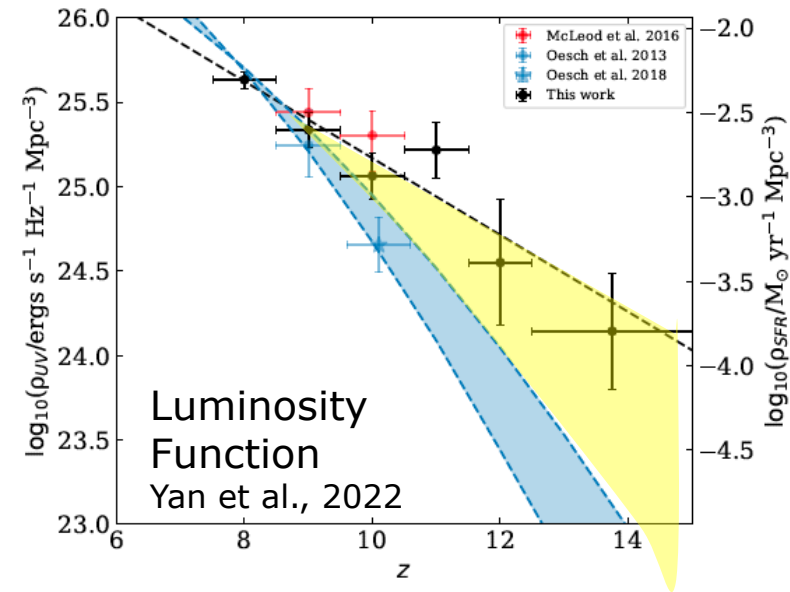


Cappelluti, Hasinger, Natarajan, 2022, ApJ

Number of Galaxies in JWST Medium-Deep Fields

First hints for a new
population of early galaxies.

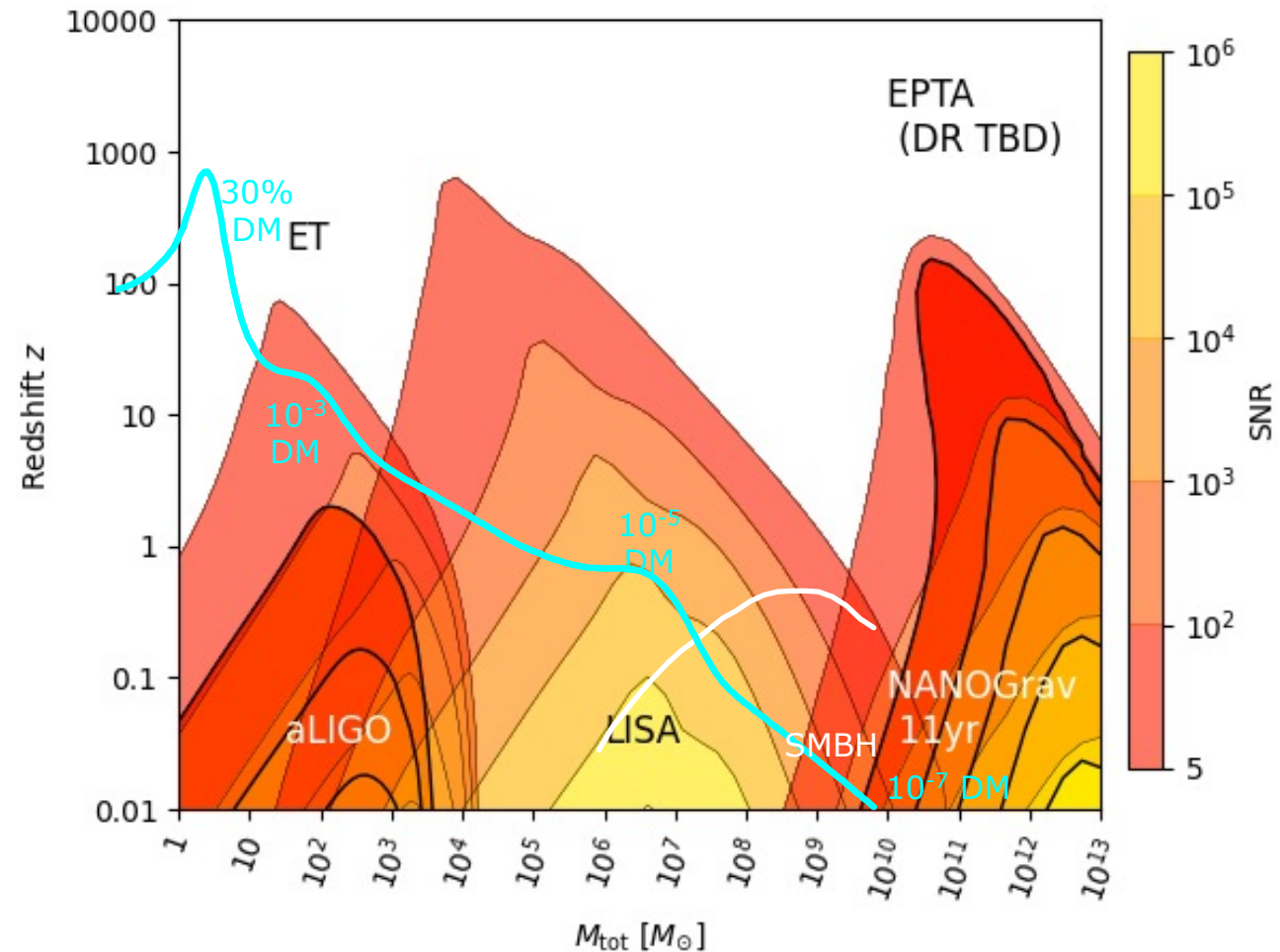
This has been predicted by
the primordial black hole dark
matter model!



Sensitivity to BH-BH Mergers

So far, all of these “fingerprints” are tantalizing, but only circumstantial evidence.

However, future Gravitational Wave observations can uniquely discriminate between astrophysical and primordial black holes!



THE SPECTRUM OF GRAVITATIONAL WAVES

Observatories
& experiments

Ground-based
experiment



Space-based observatory



Pulsar timing array



Cosmic microwave
background polarisation



Timescales

milliseconds

seconds

hours

years

billions of years

Frequency (Hz)

100

1

10^{-2}

10^{-4}

10^{-6}

10^{-8}

10^{-16}

Cosmic fluctuations in the early Universe

Cosmic
sources



Supernova



Pulsar



Compact object falling
onto a supermassive
black hole



Merging supermassive black holes



Merging neutron
stars in other galaxies

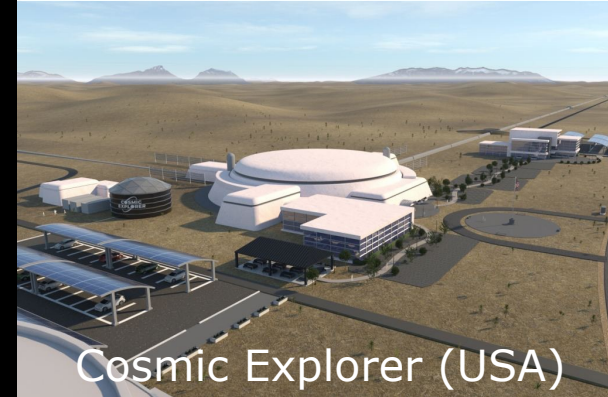
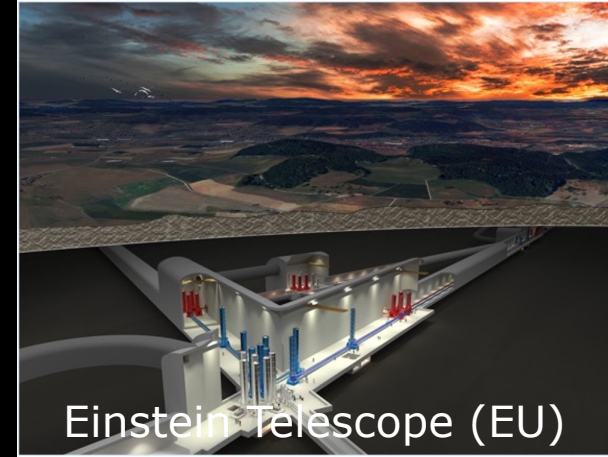
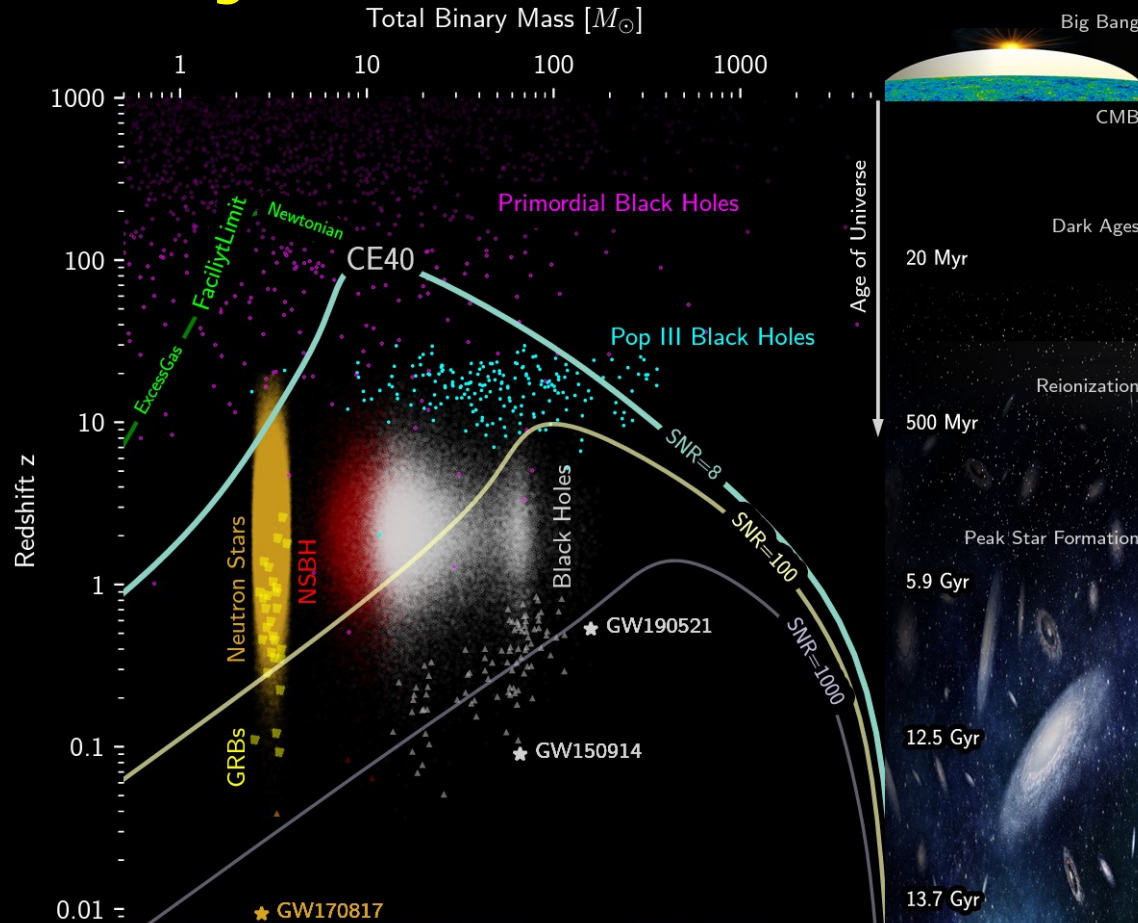


Merging stellar-mass black holes
in other galaxies



Merging white dwarfs
in our Galaxy

Sensitivity of the next generation groundbased gravitational wave interferometers



Forschung.
Technologie.
Digitalisierung.

**The new national center
for astrophysics in
Lusatia (DZA)**



Deutsches Zentrum für Astrophysik



Thank you very much!